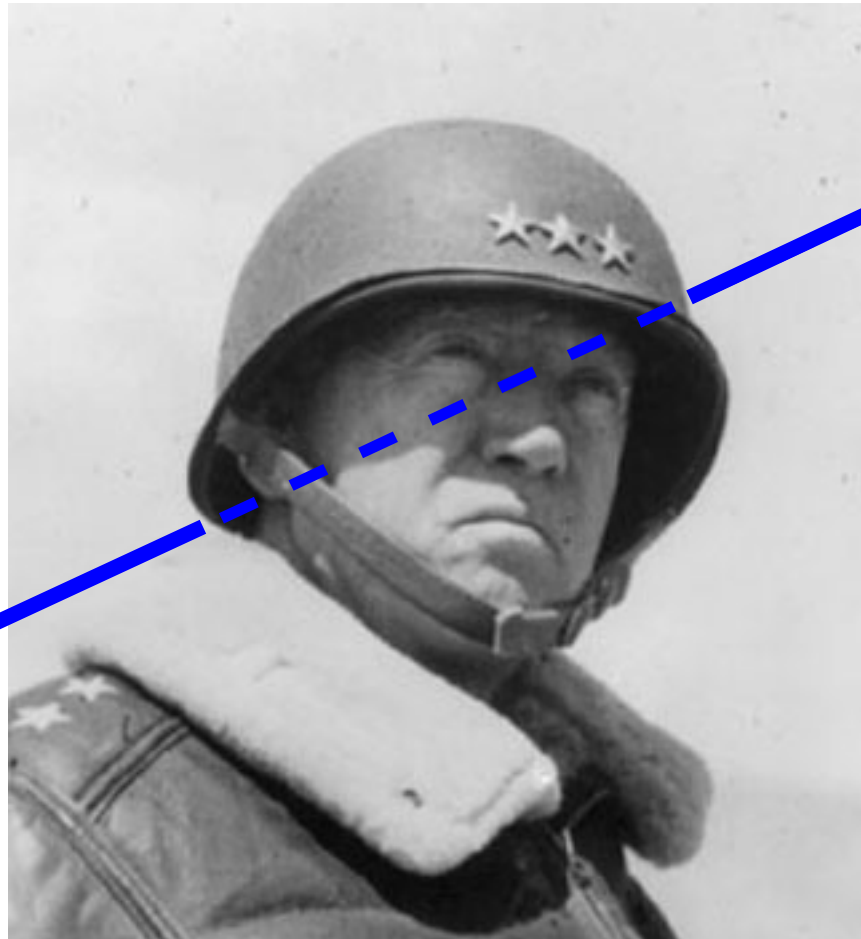


Muons

I: Muons in General



μ

	I	II	III	
Quarks	u	c	t	γ
	d	s	b	g
Leptons	ν_e	ν_μ	ν_τ	Z
	e	μ	τ	W

Three Generations of Matter

Force Carriers



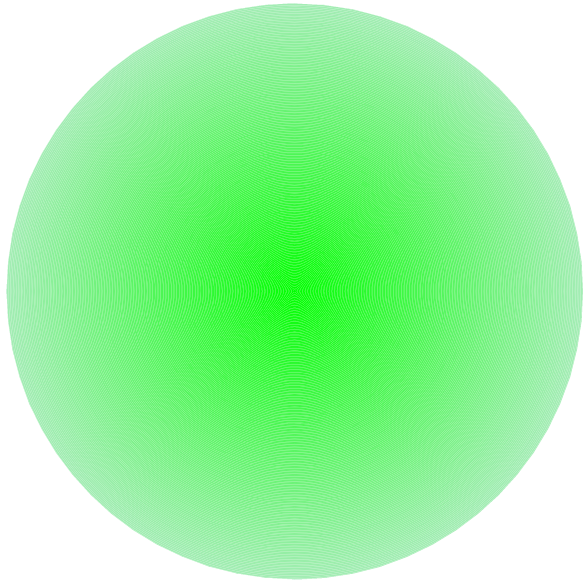
!

e^+

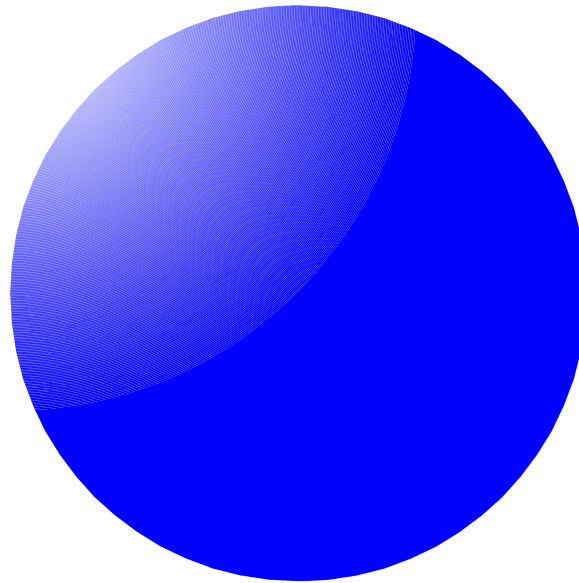


μ

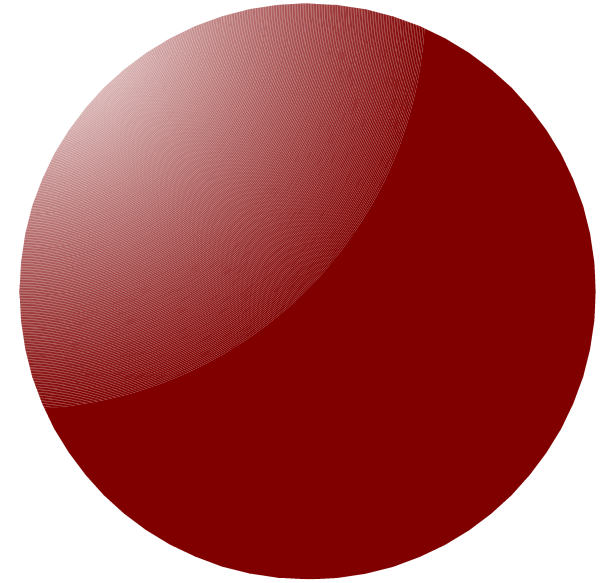
C.D. Anderson, 1936



$e: 1q_e$



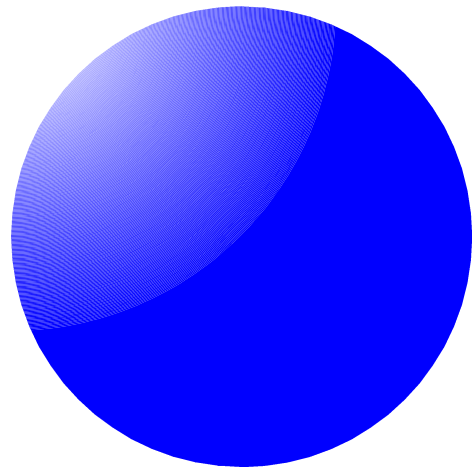
$\mu: 1q_e$



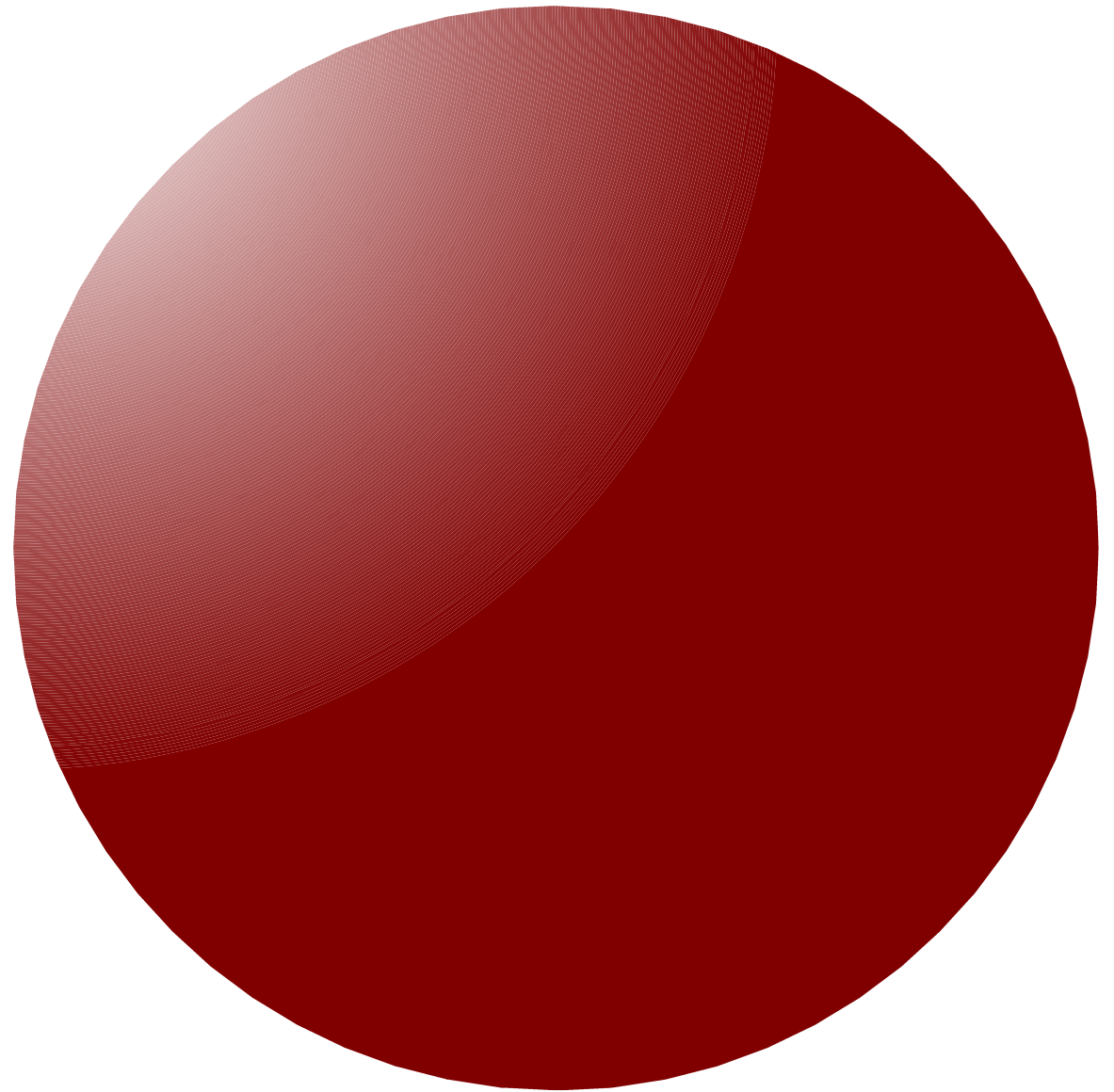
$\tau: 1q_e$

$$1q_e = 1.602 \times 10^{-19} \text{C}$$

e: 512keV/c²



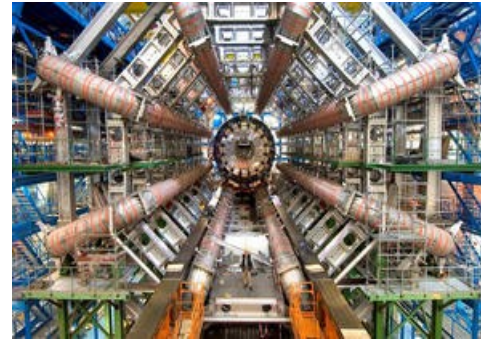
μ: 106MeV/c²



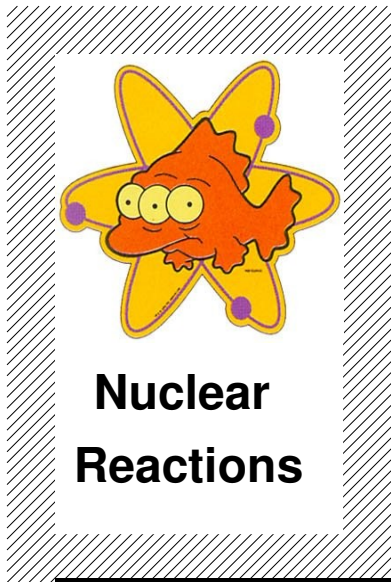
τ: 1.777GeV/c²

1eV = 1.602 × 10⁻¹⁹J

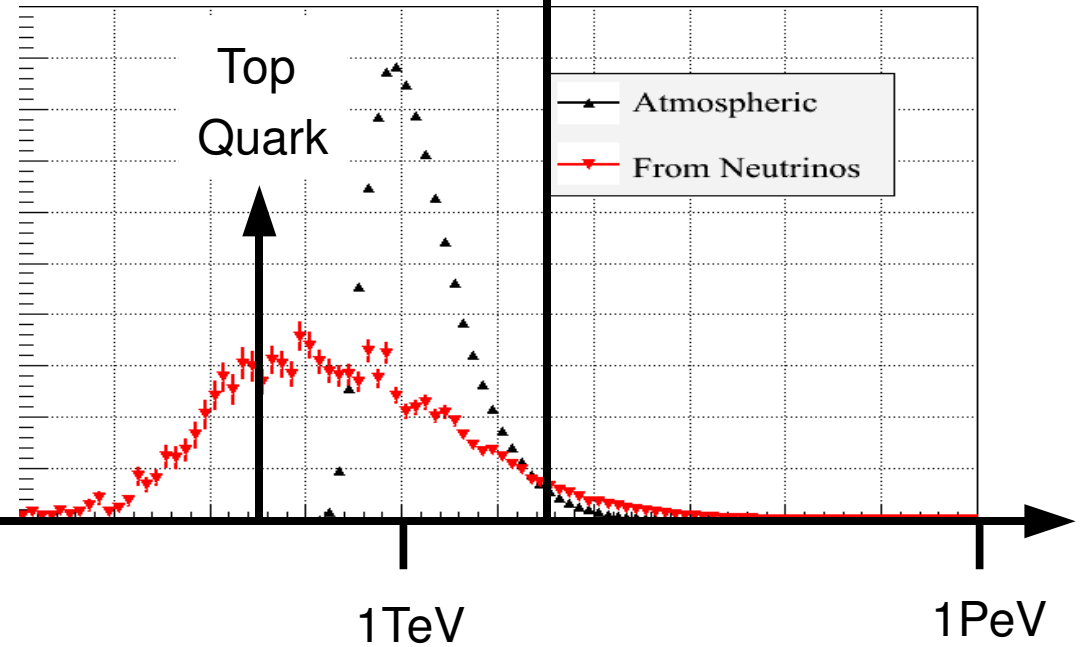
LHC



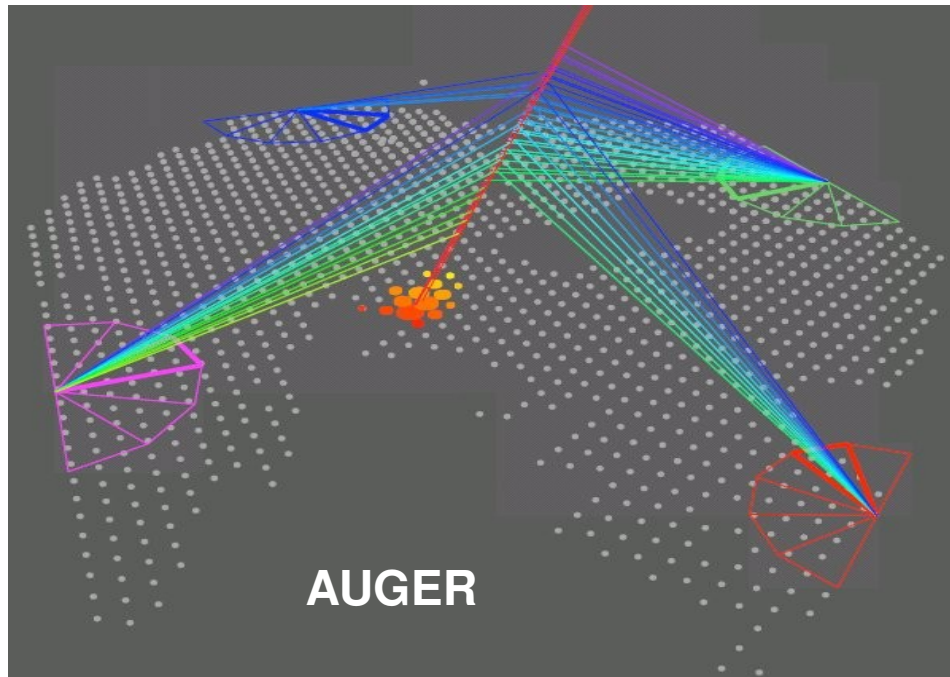
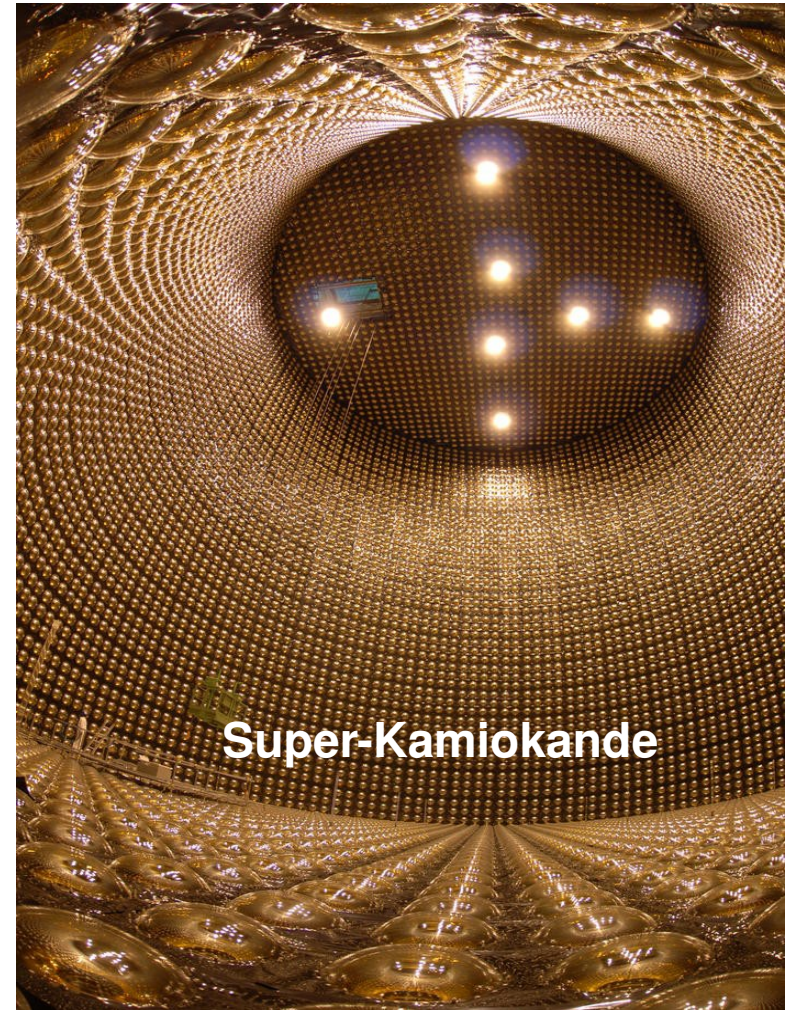
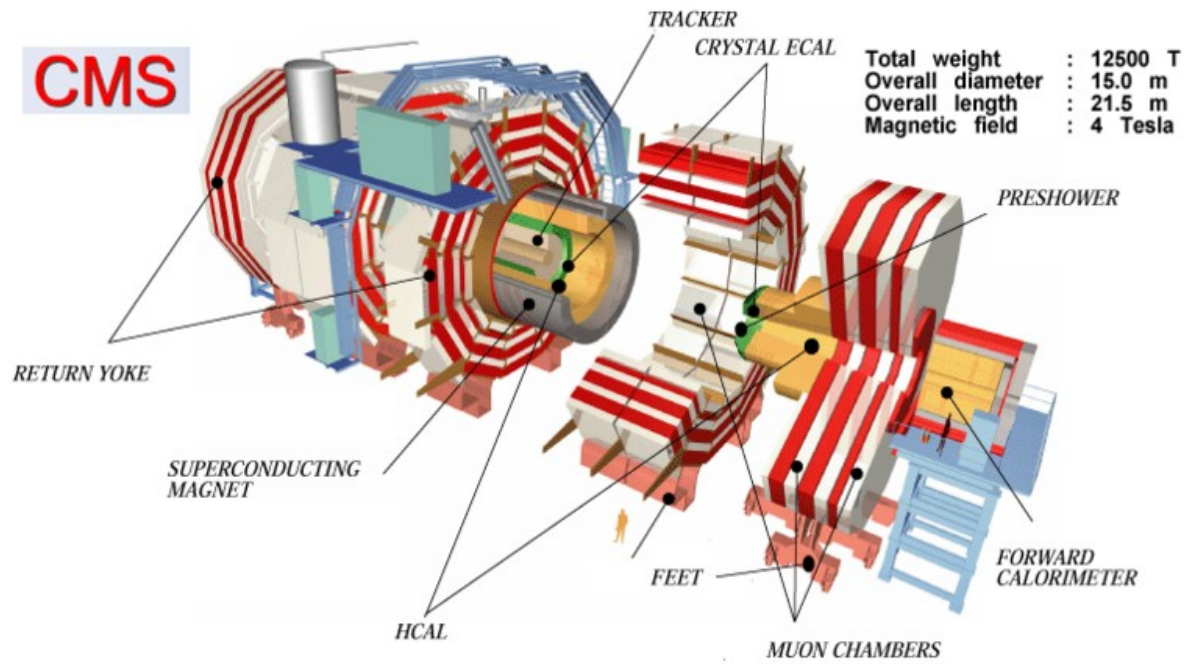
$$E = m_0 c^2$$



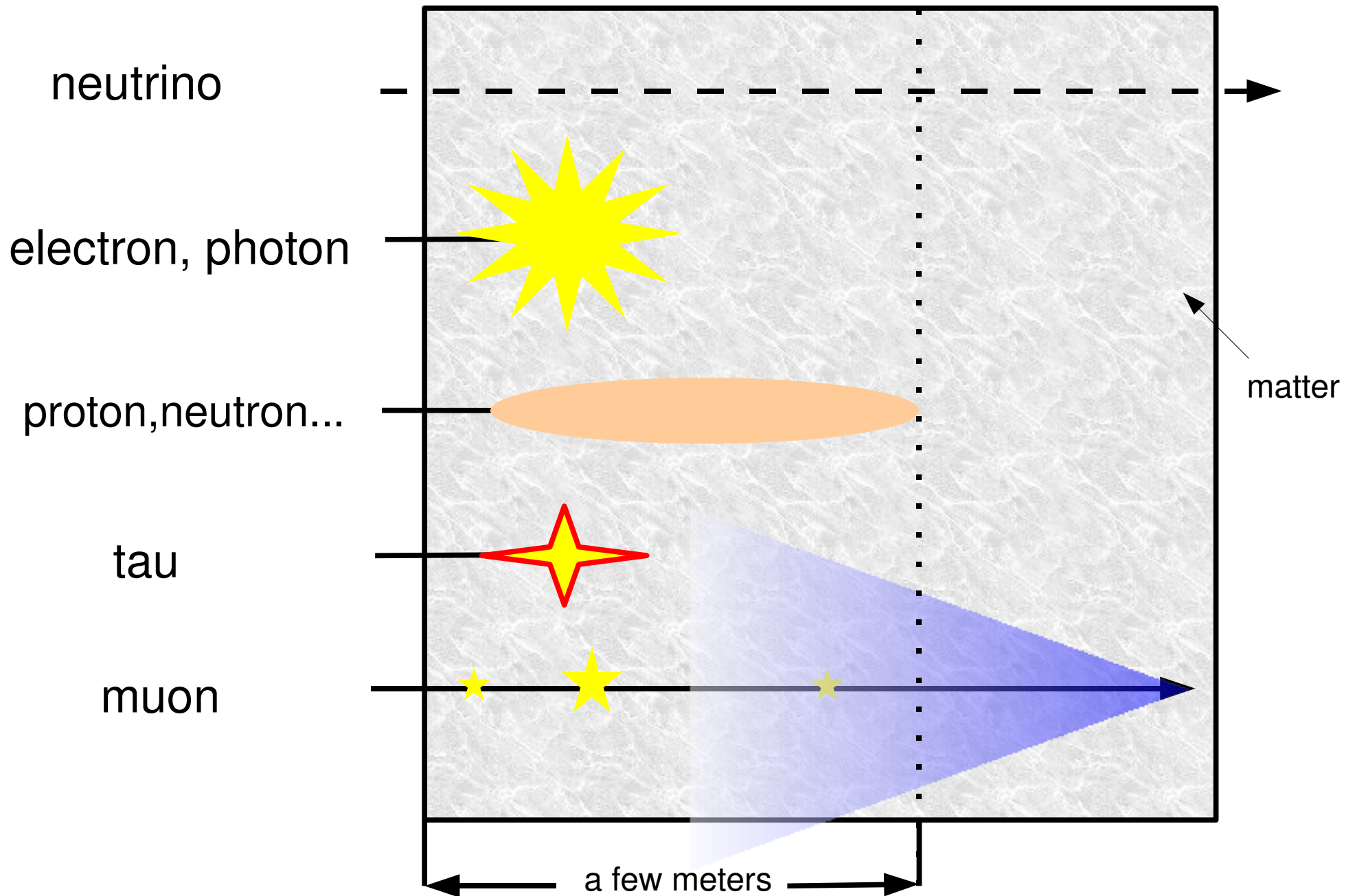
Muon
Proton/
Neutron



CMS



High (but not too high!) Energies

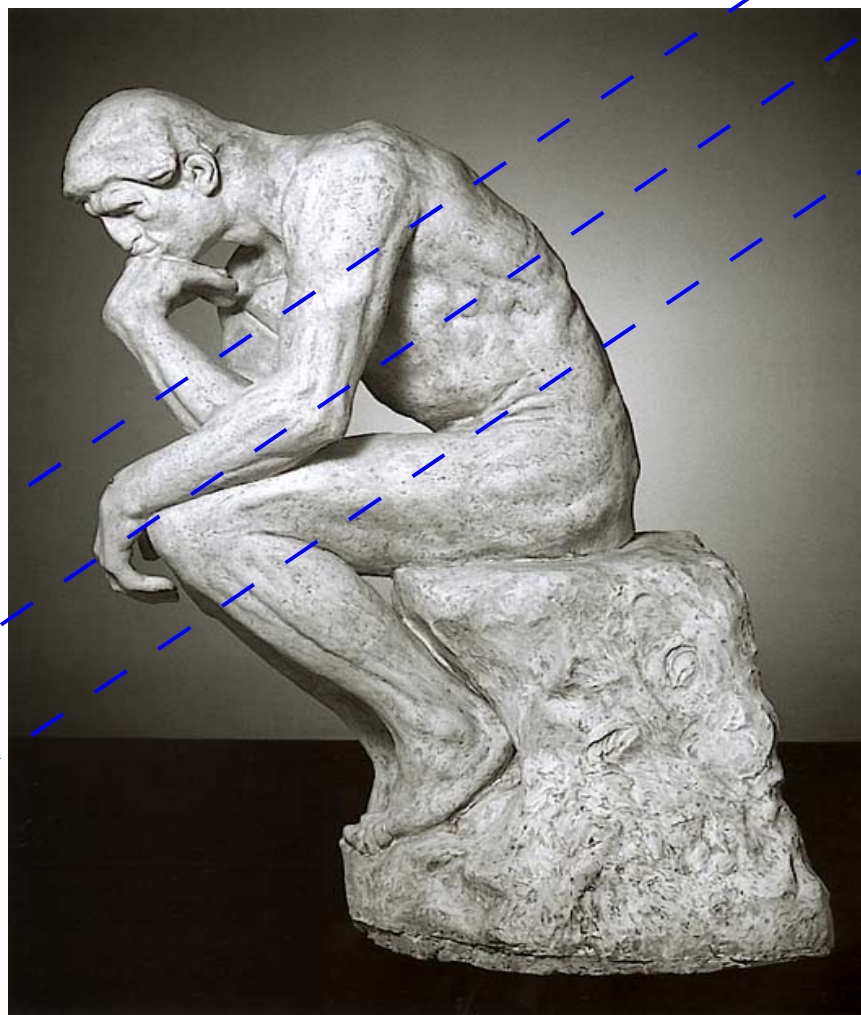




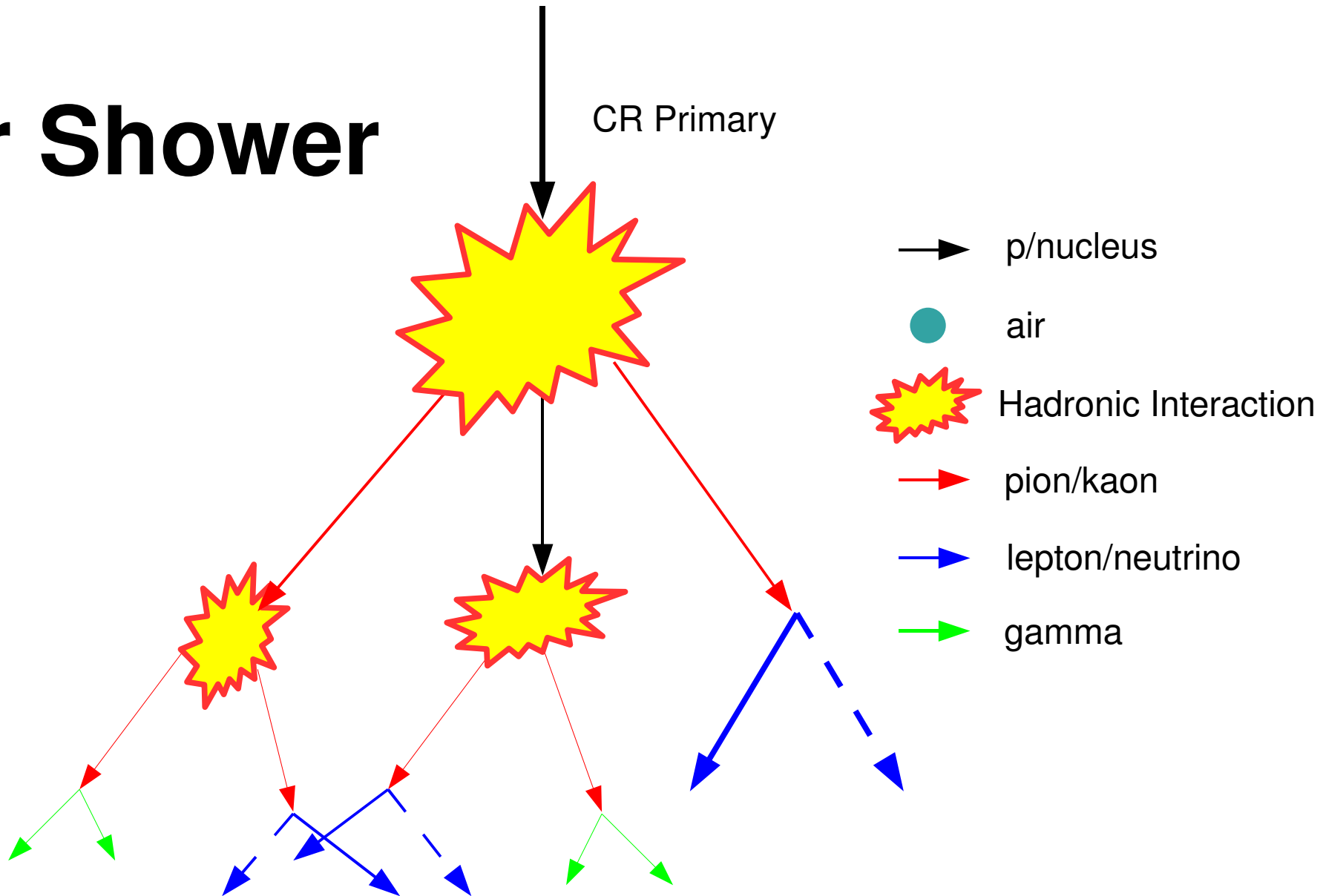
Total flux	Muons	neutrons	Electrons	Protons, pions
$3 \cdot 10^{-2} \text{ cm}^{-2} \text{ s}^{-1}$	63%	21%	15%	< 1%

nucl-ex/0601019

$\approx 100 \mu\text{/s}$

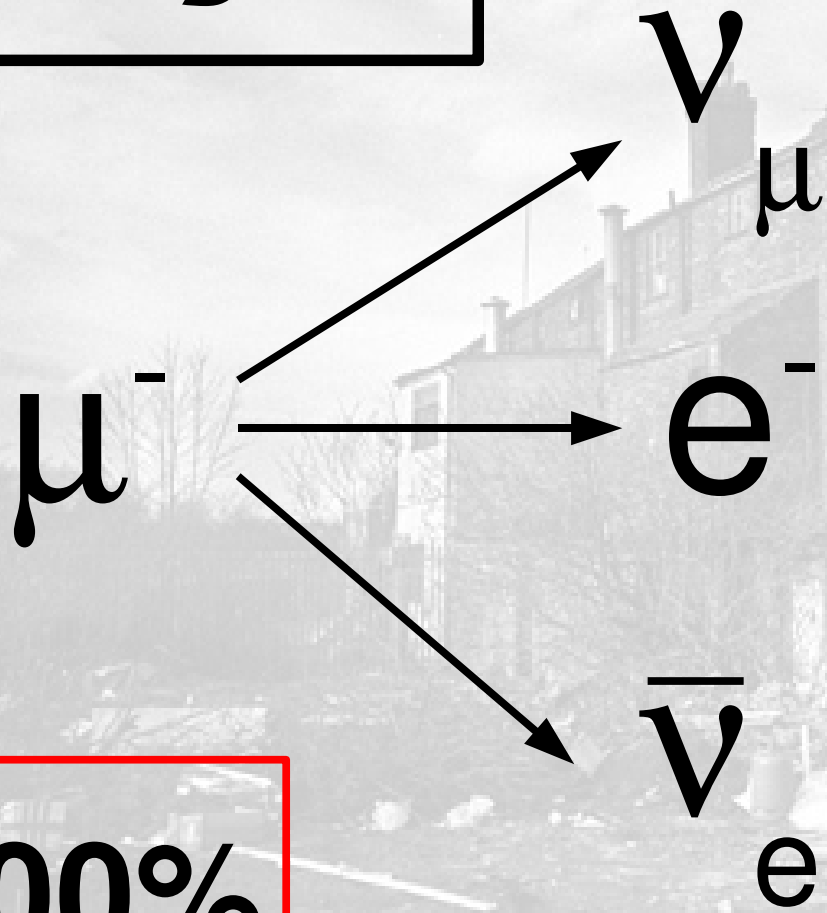


Air Shower



Conservation Law

Decay...

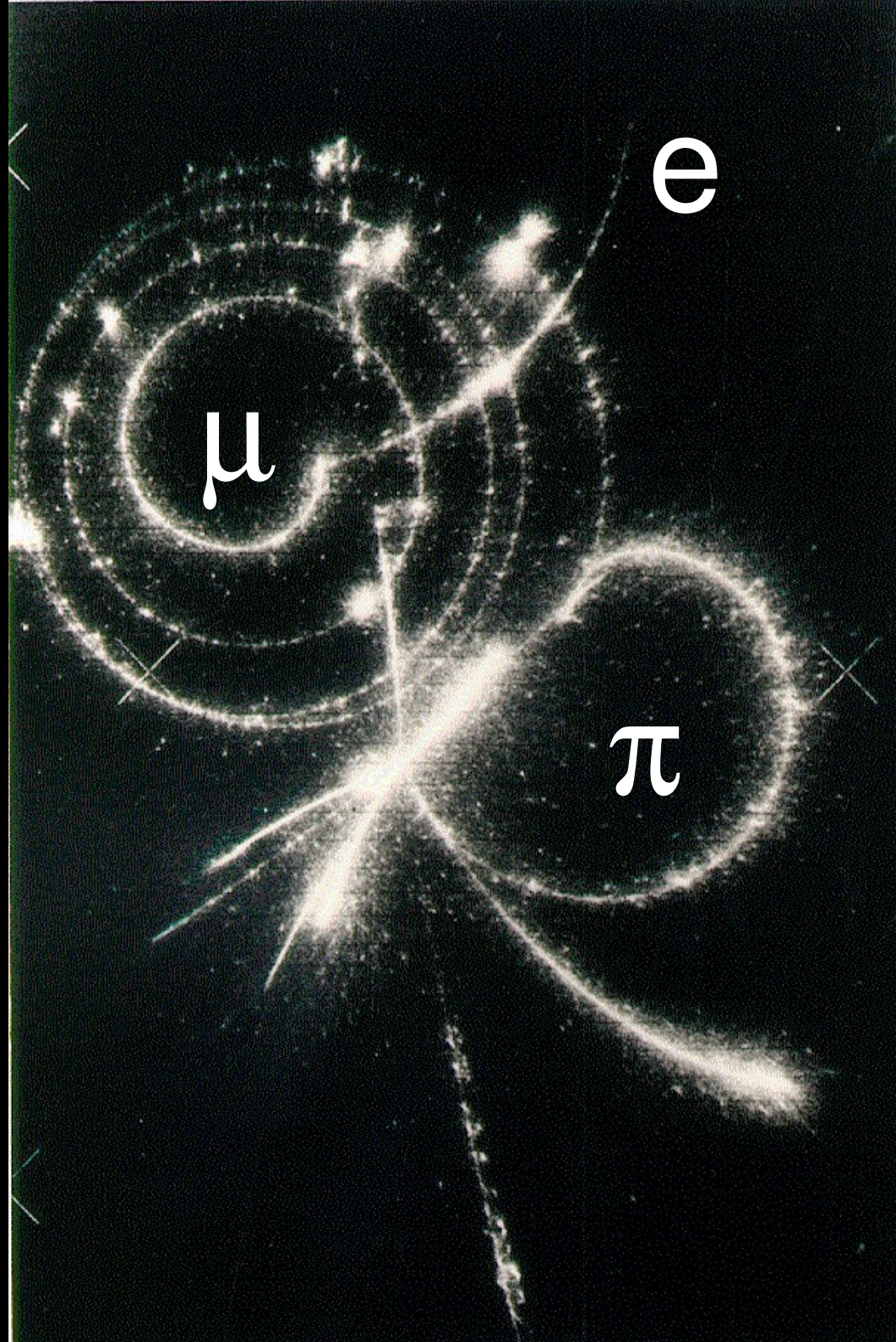


muonic number
(approx.!)
Charge

lepton number

$\approx 100\%$

Σ : Energy



$$3 \times 10^{-8} \text{ m/s} \rightarrow \gamma \text{CT} \leftarrow 2.197 \times 10^{-6} \text{ s}$$

$$E/m_0 c^2 \rightarrow \gamma \text{CT}$$

→ E=105MeV ($\gamma=1$): 660m

→ Shower Maximum to Detector Center (vertical): $\approx 10\text{km}$

→ Shower Maximum to Detector Center (horizontal): $\approx 100\text{km}$

→ E=400GeV ($\gamma=4,000$): 2,640km

→ Earth Diameter: 12,740km

→ Earth-Moon: 300,000km

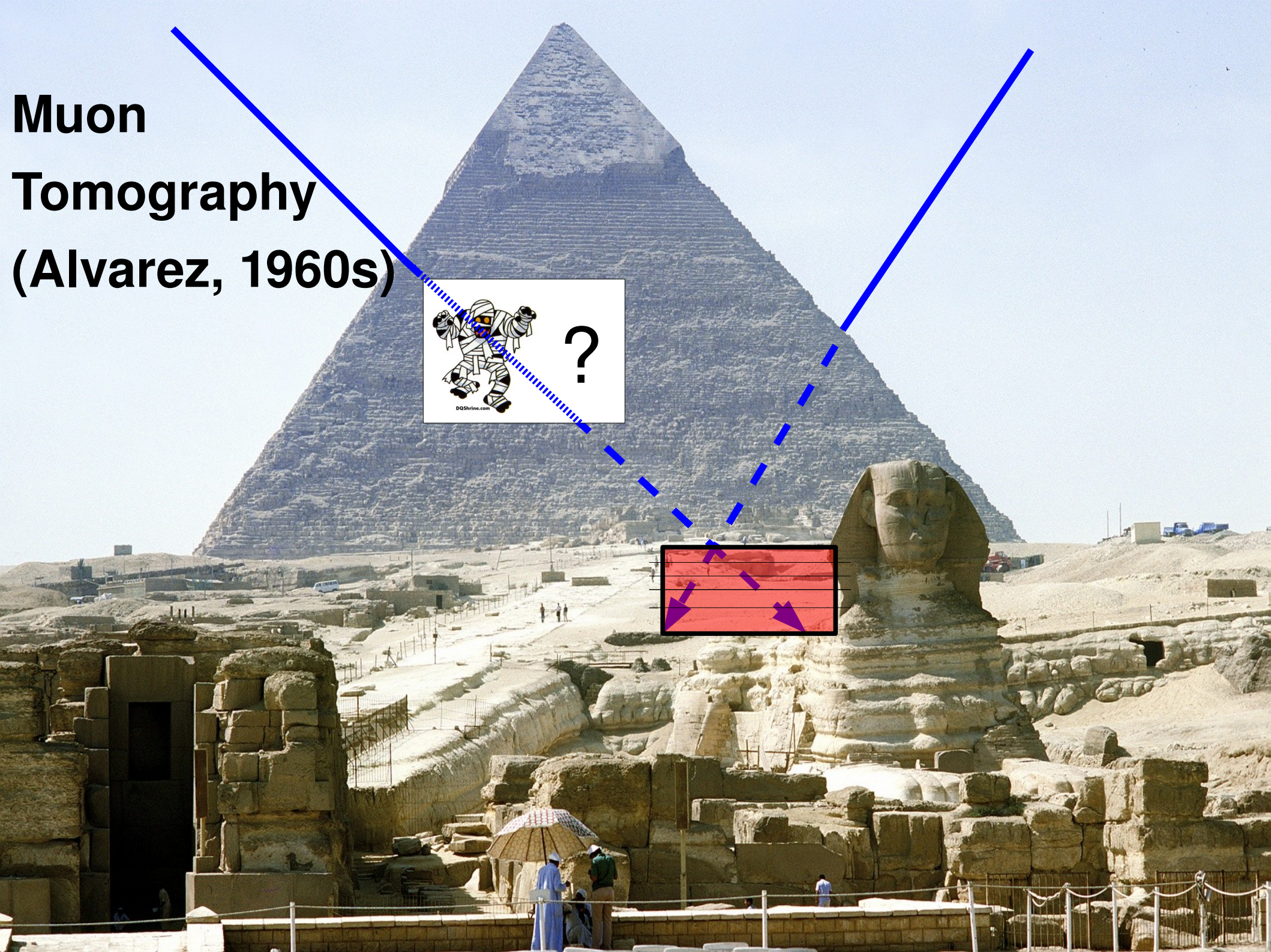
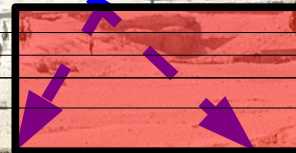
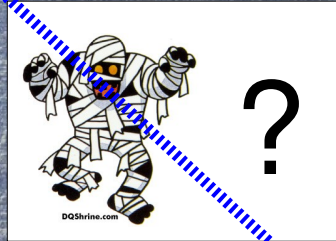
→ E=100TeV ($\gamma=1,000,000$): 660,000km

→ log d

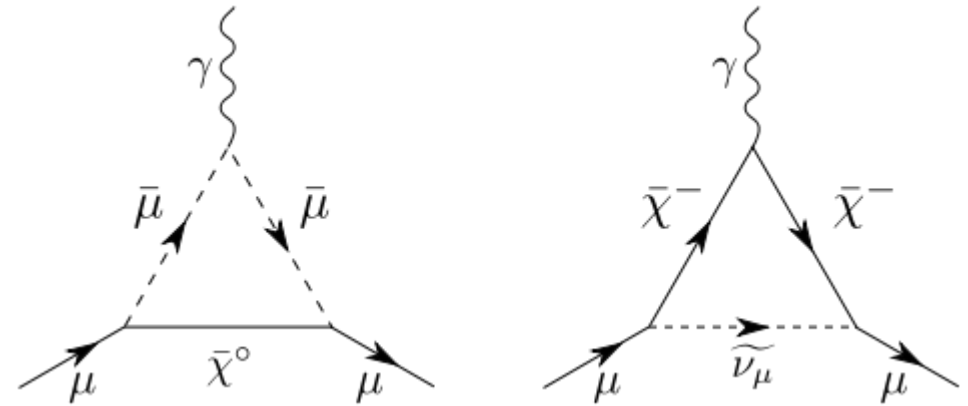
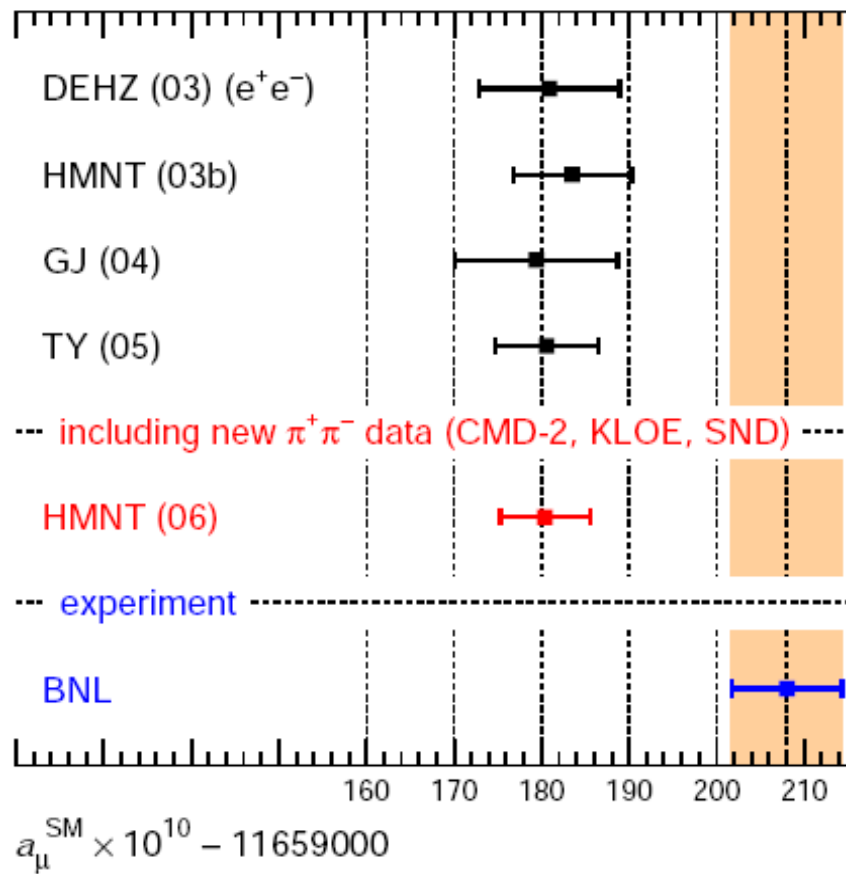
Outside

HEP/Astro-Ph

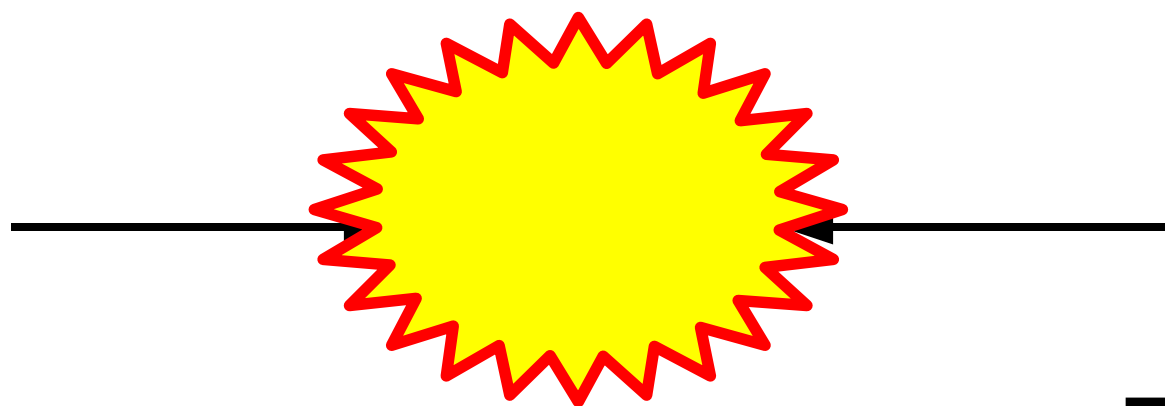
Muon Tomography (Alvarez, 1960s)



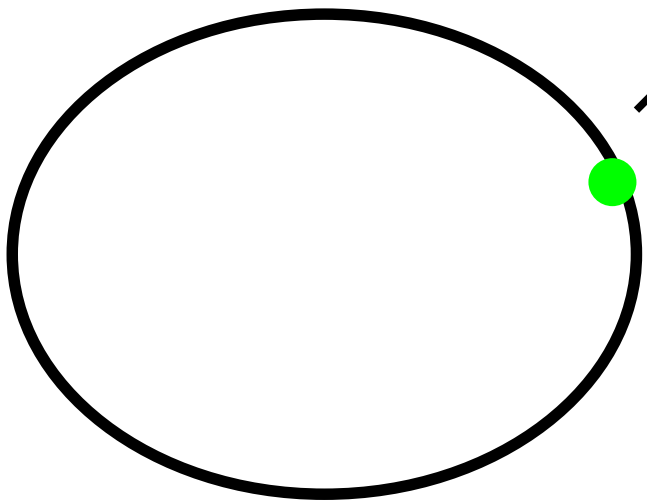
Anomalous Magnetic Moment: $a_\mu \equiv \frac{g_\mu - 2}{2}$



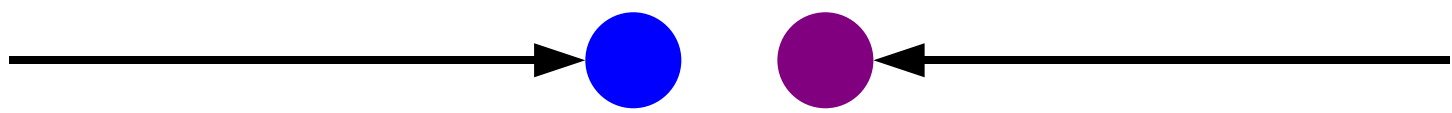
$p\bar{p}$



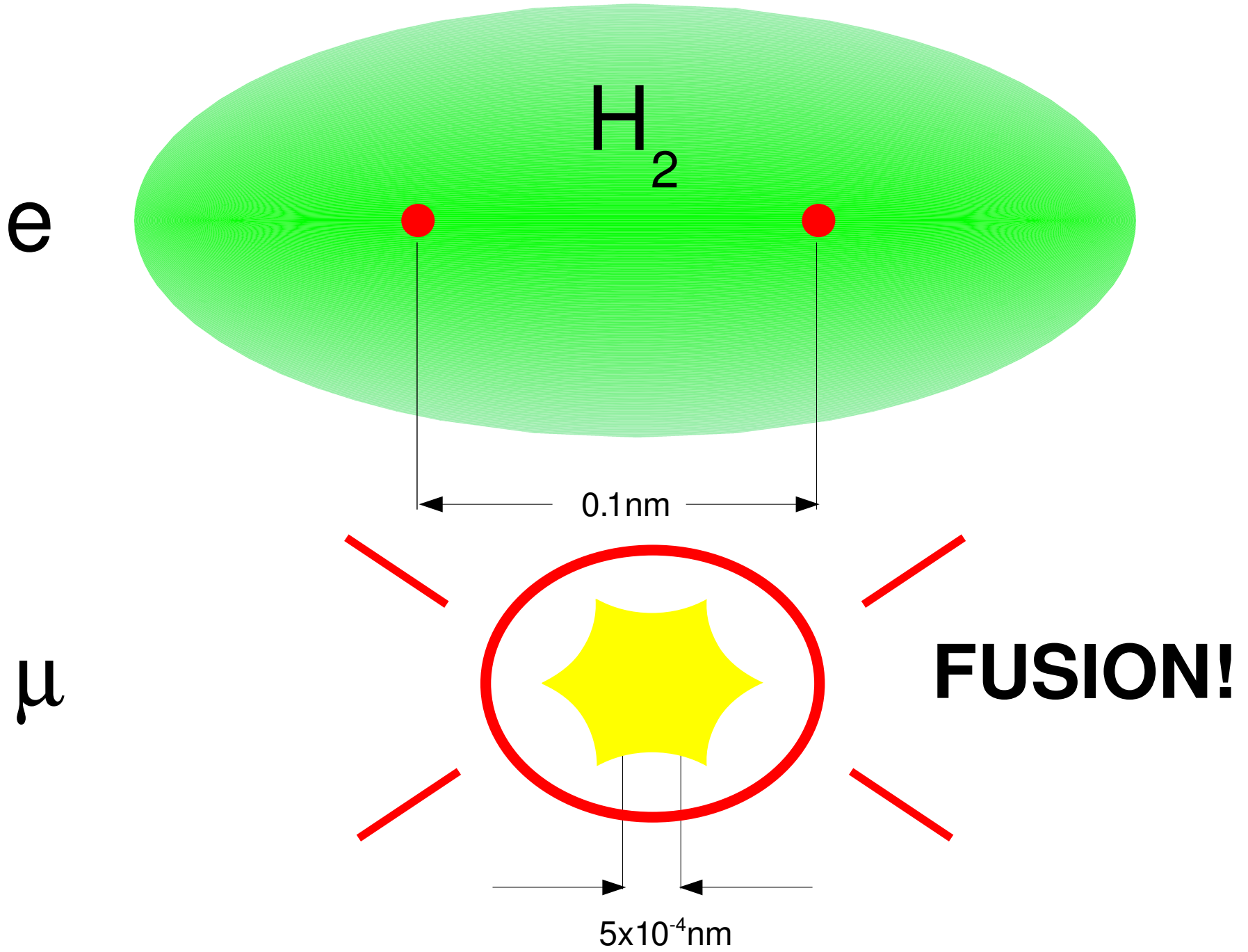
electrons



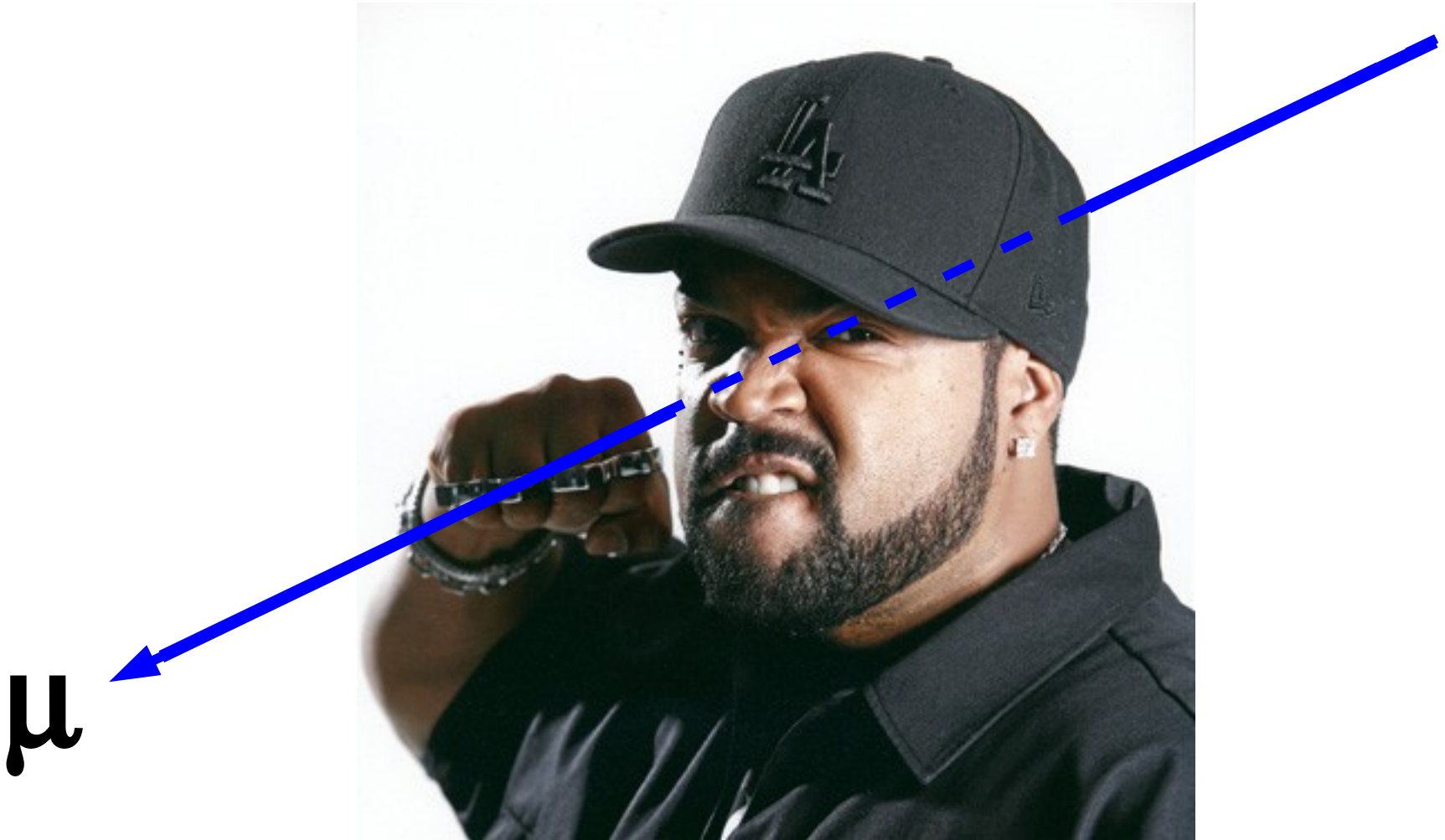
$P \propto m^{-4}$

Two black arrows originate from the green dot on the electron-positron collider diagram, pointing outwards and upwards to the right, representing the production of particles.

Muon Collider?



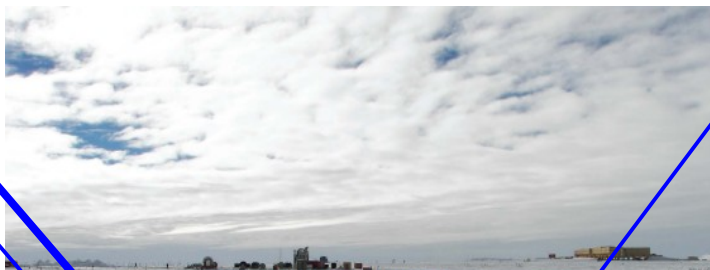
II: Muons in IceCube



Coincident Muons

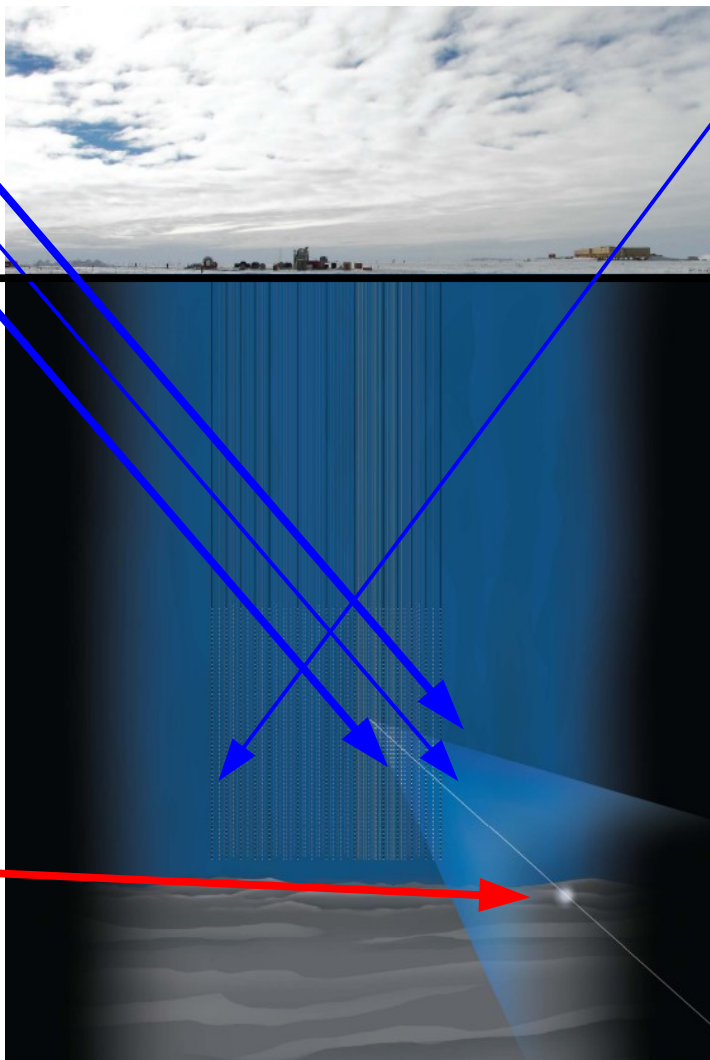
25Hz

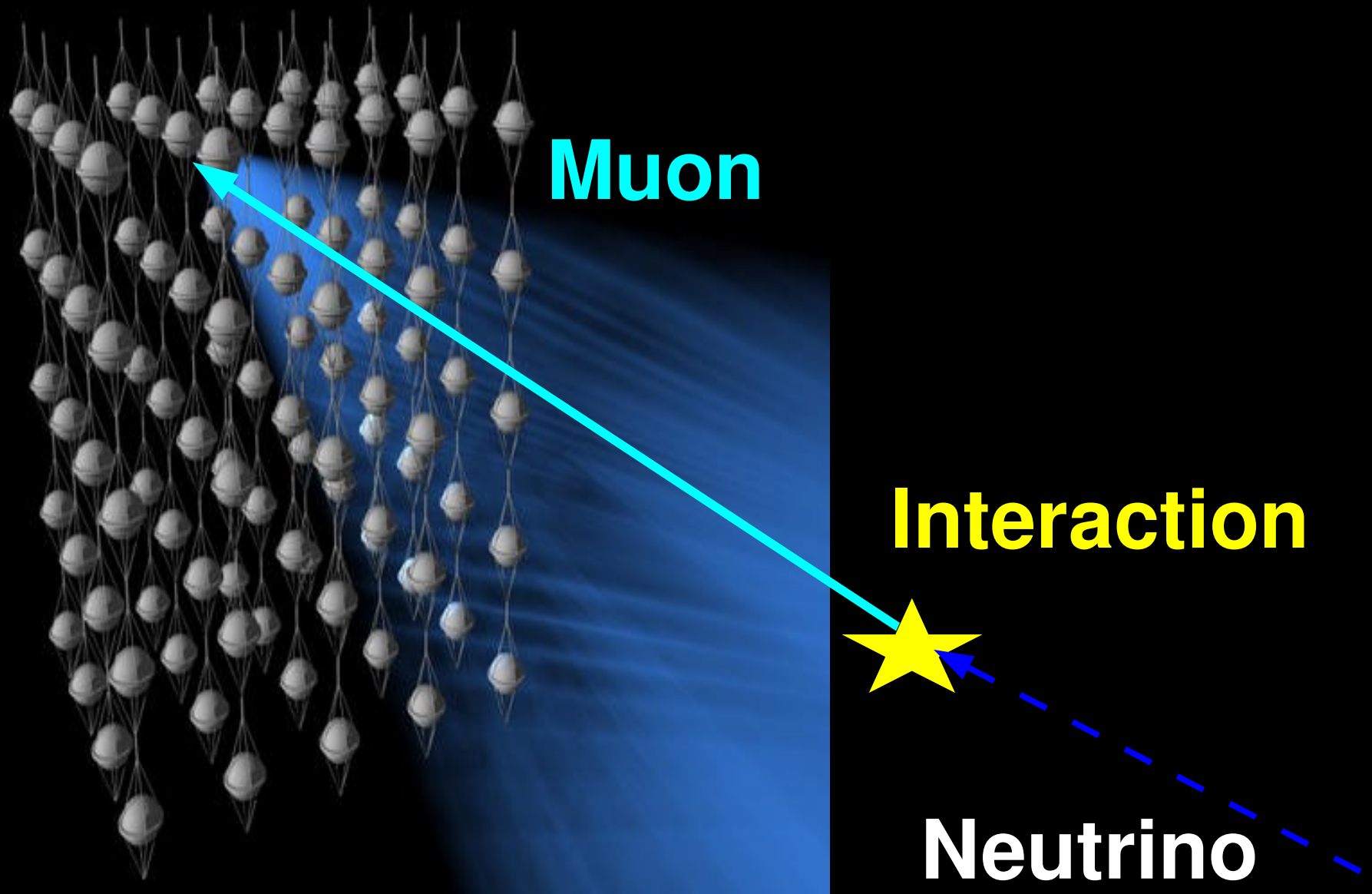
10s of meters



Muons From
Neutrinos

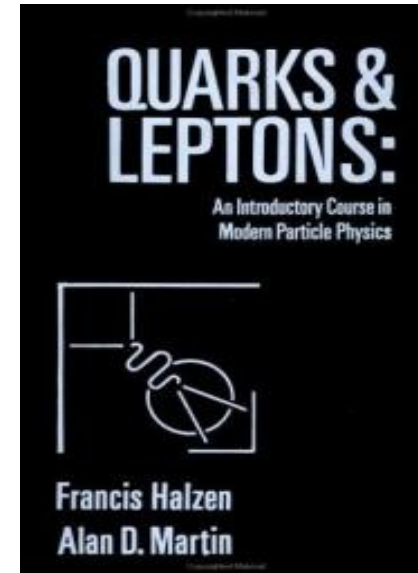
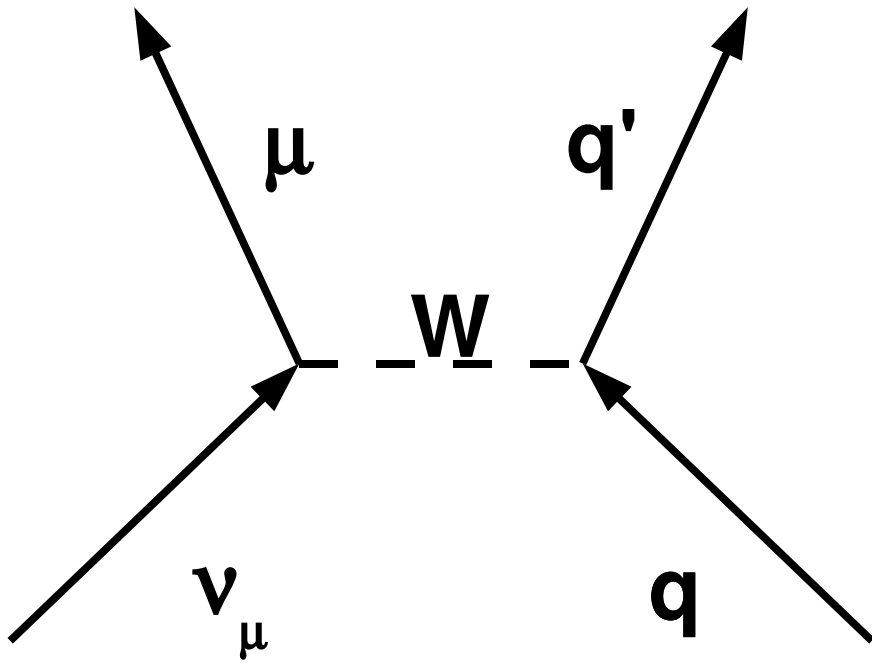
$\approx 20/h$





also: τ , hadron decay to μ (rare)

Deep Inelastic Scattering

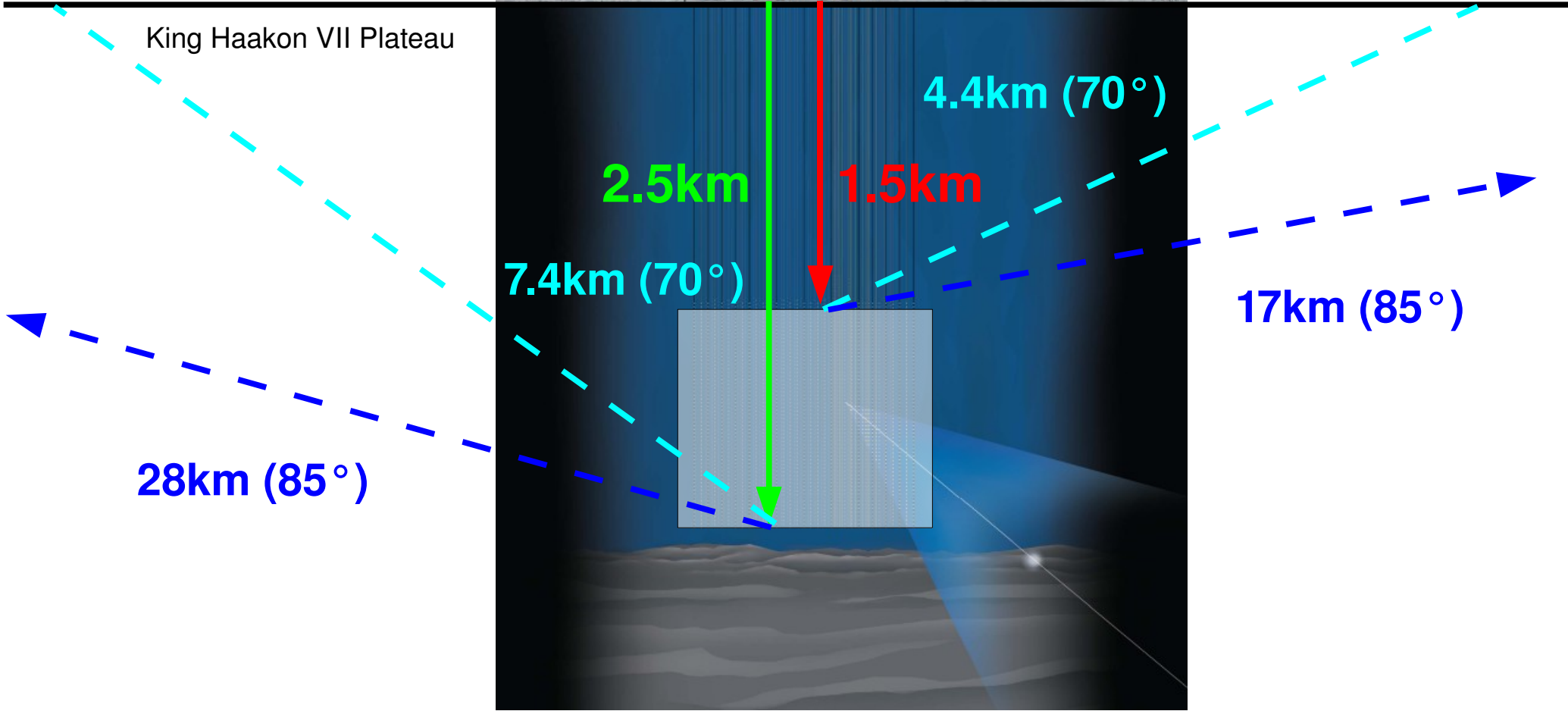


?



!

Intermezzo



Muon Monte Carlo: a high-precision tool for muon propagation through matter

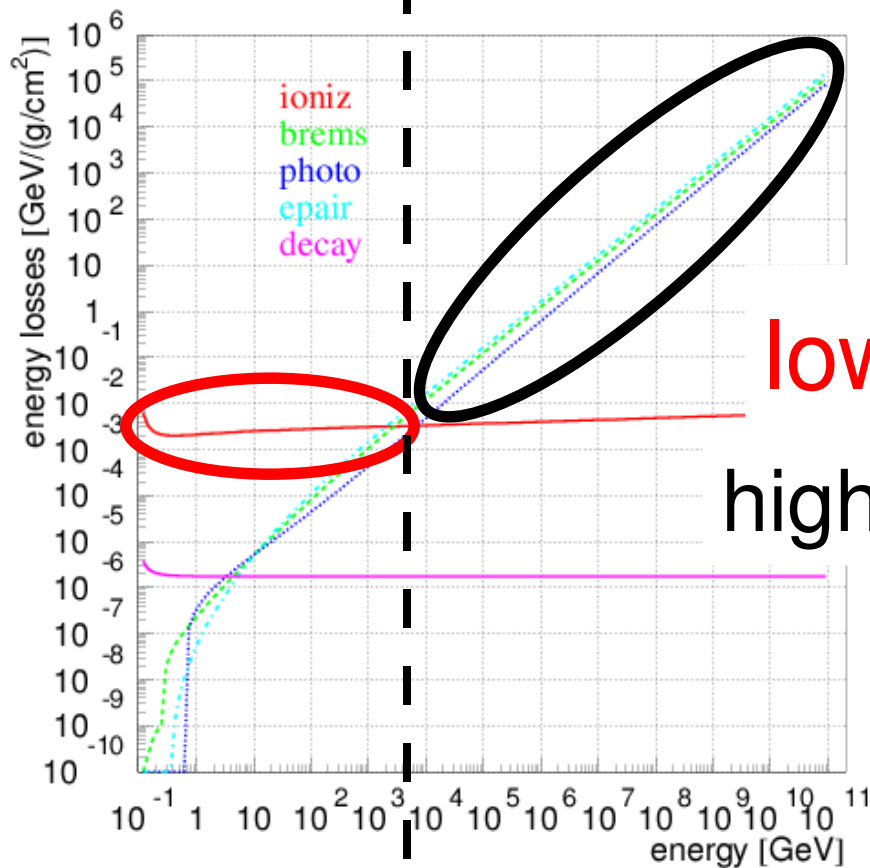
hep-ph/0407075

Dmitry Chirkin¹, Wolfgang Rhode²

chirkin@physics.berkeley.edu

rhode@uni-wuppertal.de

few TeV



low E: flat

high E: steep

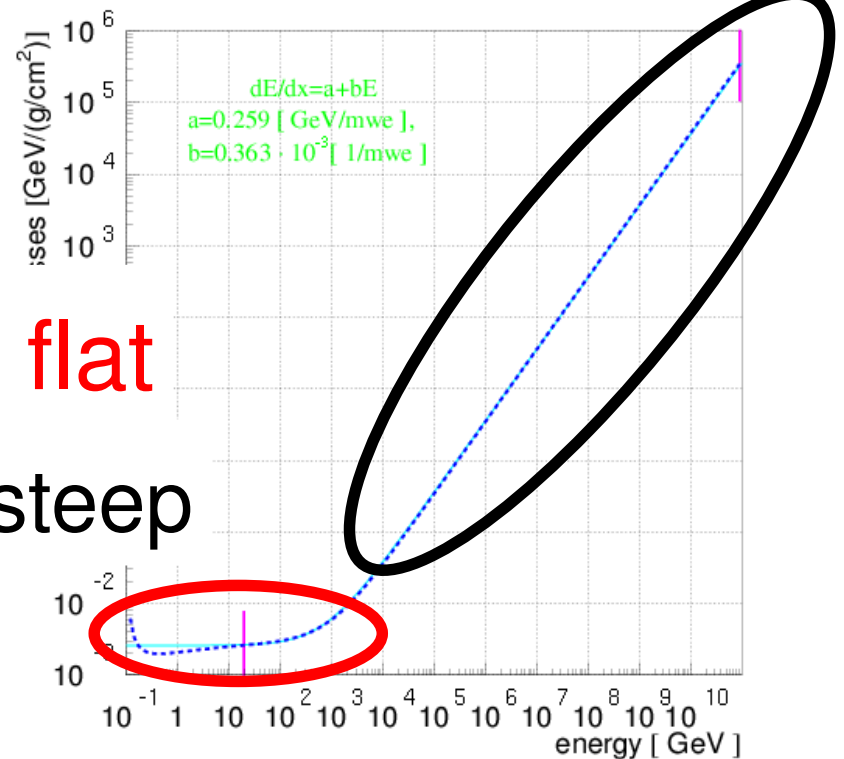


Fig. 21. Fit to the energy losses in ice

Threshold Energy

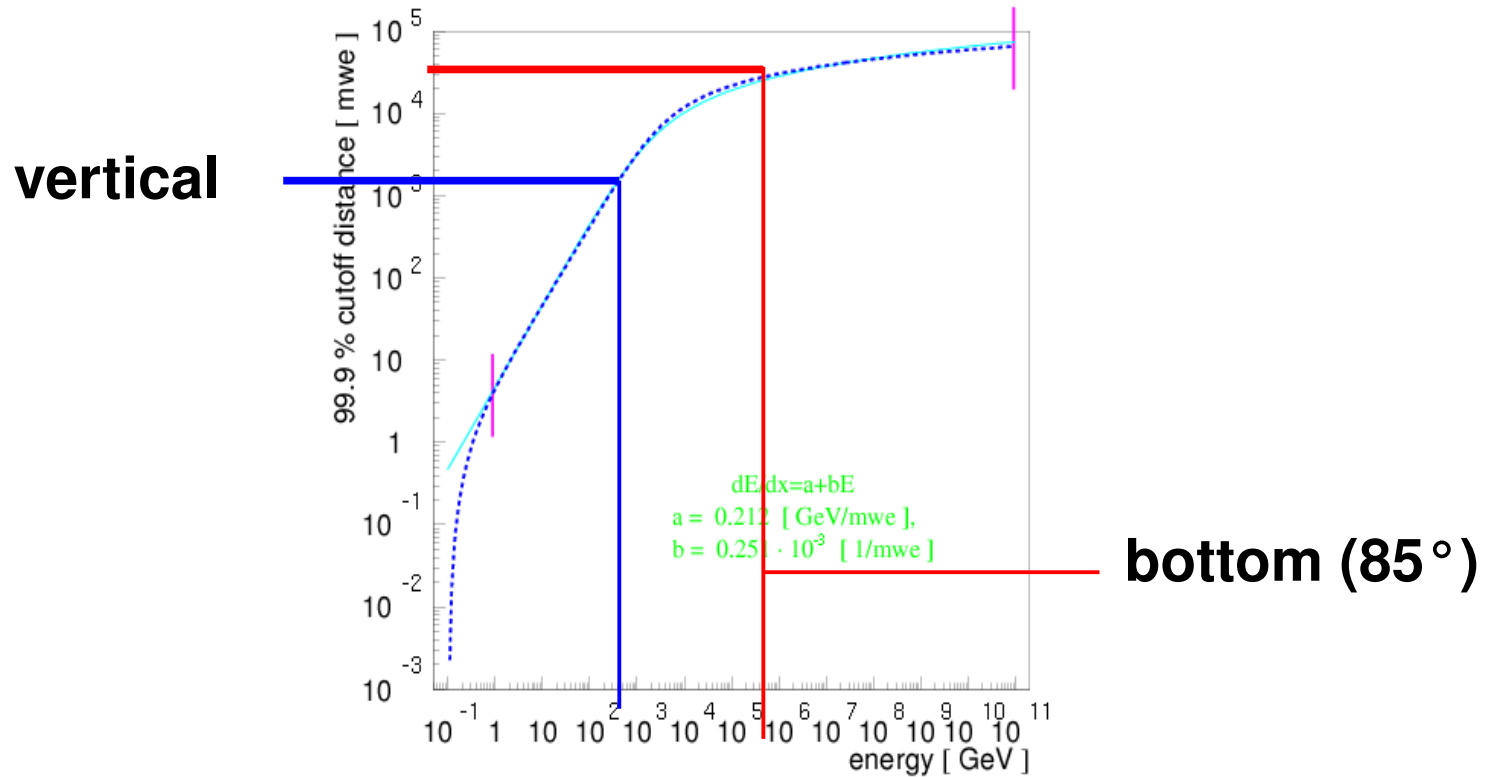
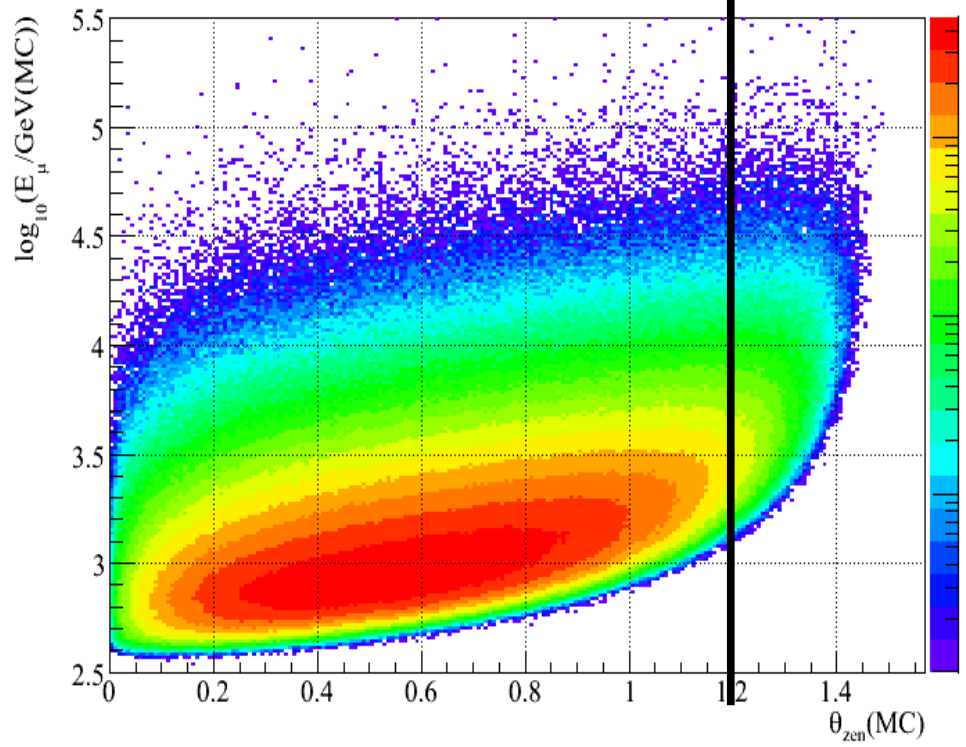
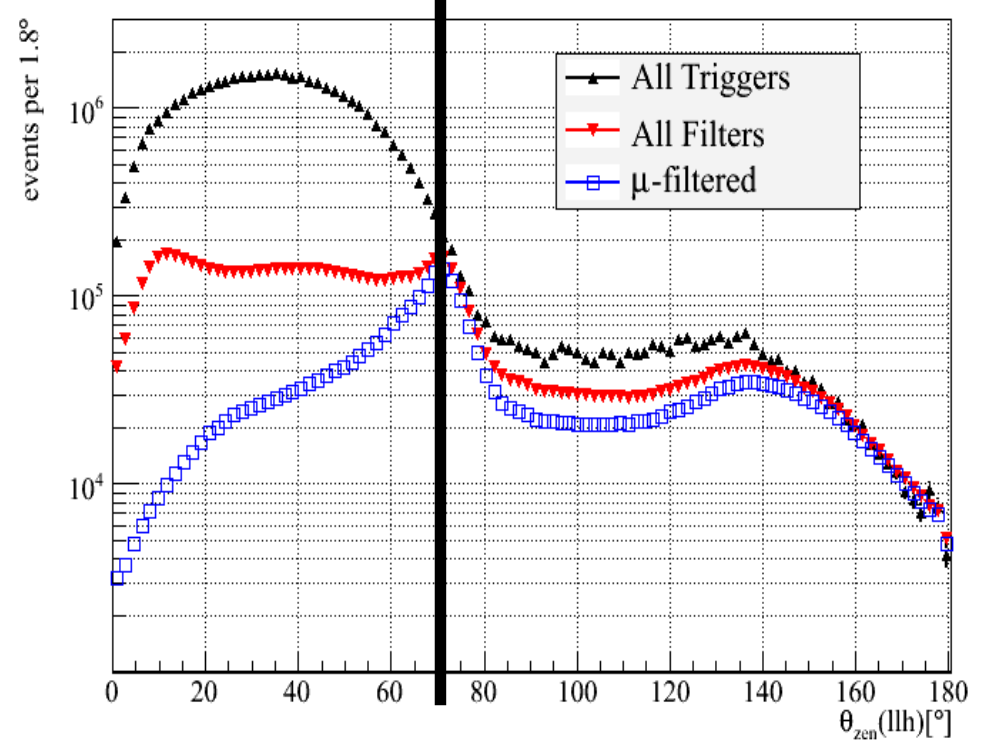


Fig. 29. Fit to the $E_{cut}(x)$

Zenith Angle vs. Muon Surface Energy

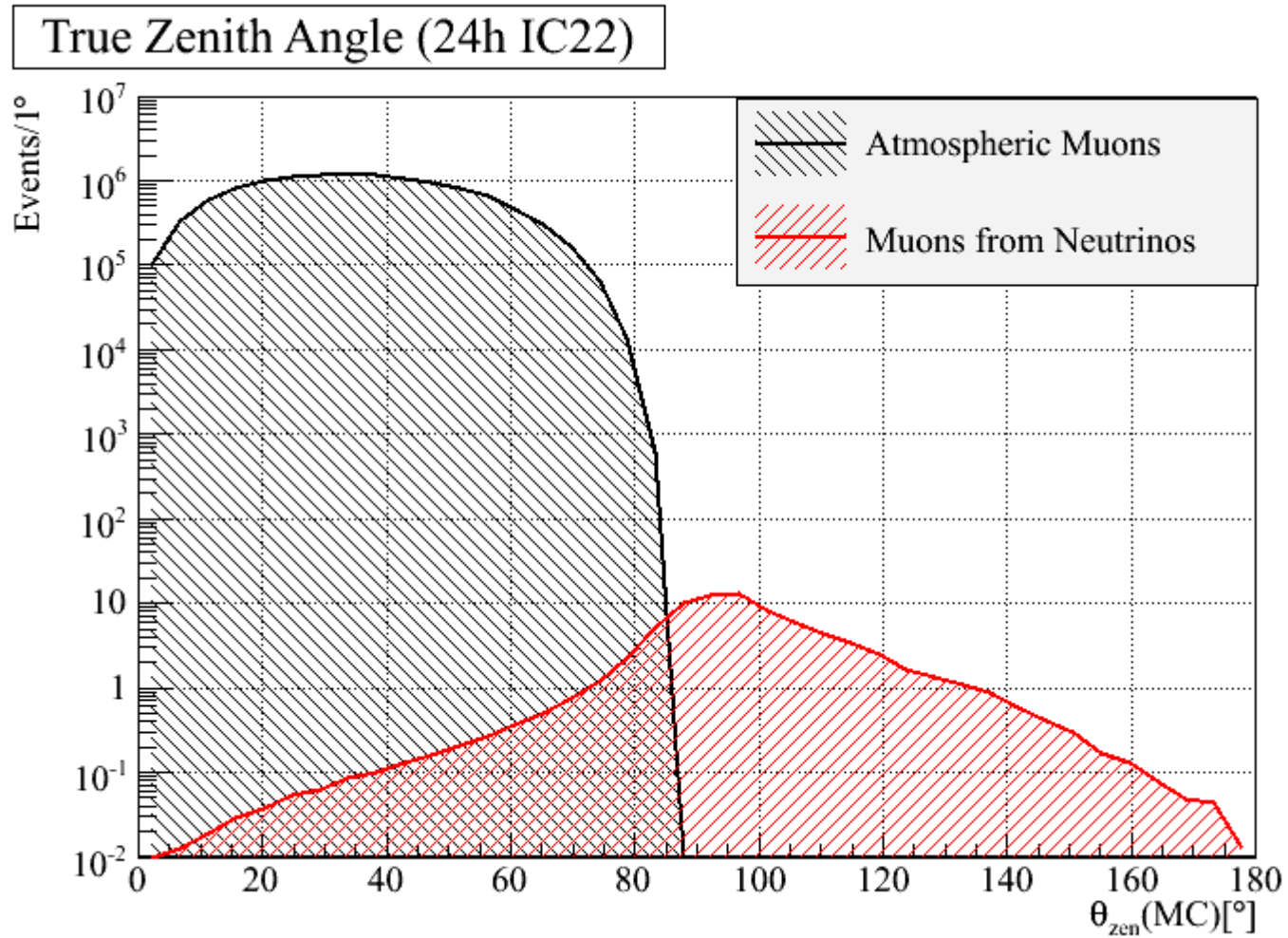


Llh Zenith Angle (12h IC22)

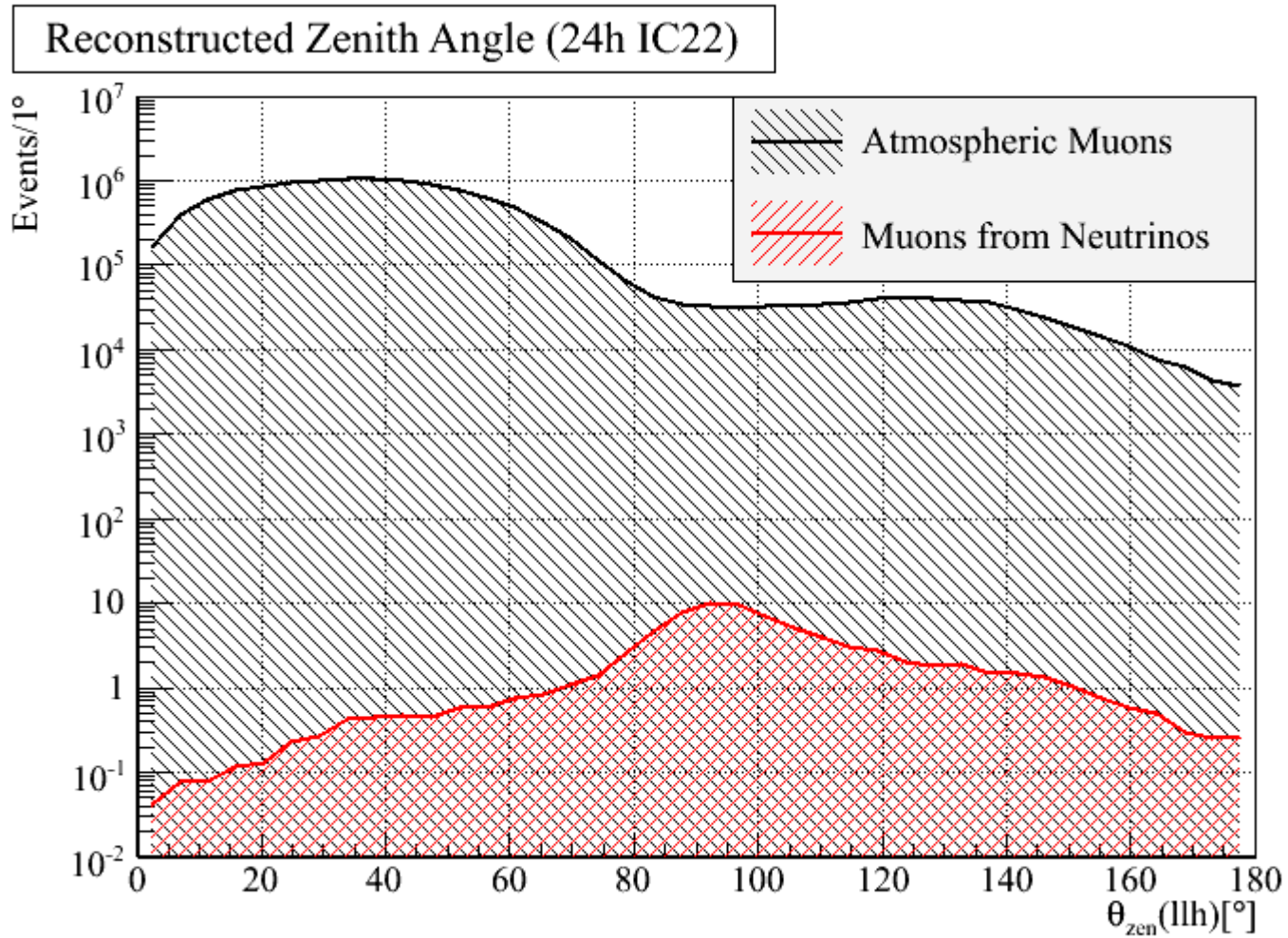


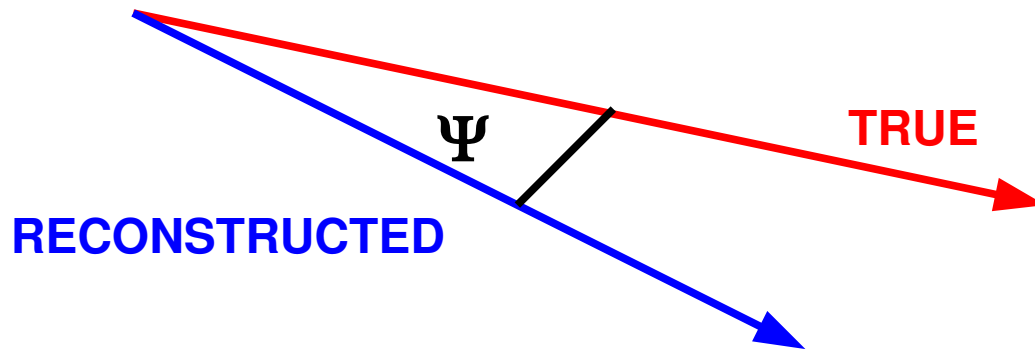
70°

Ideal

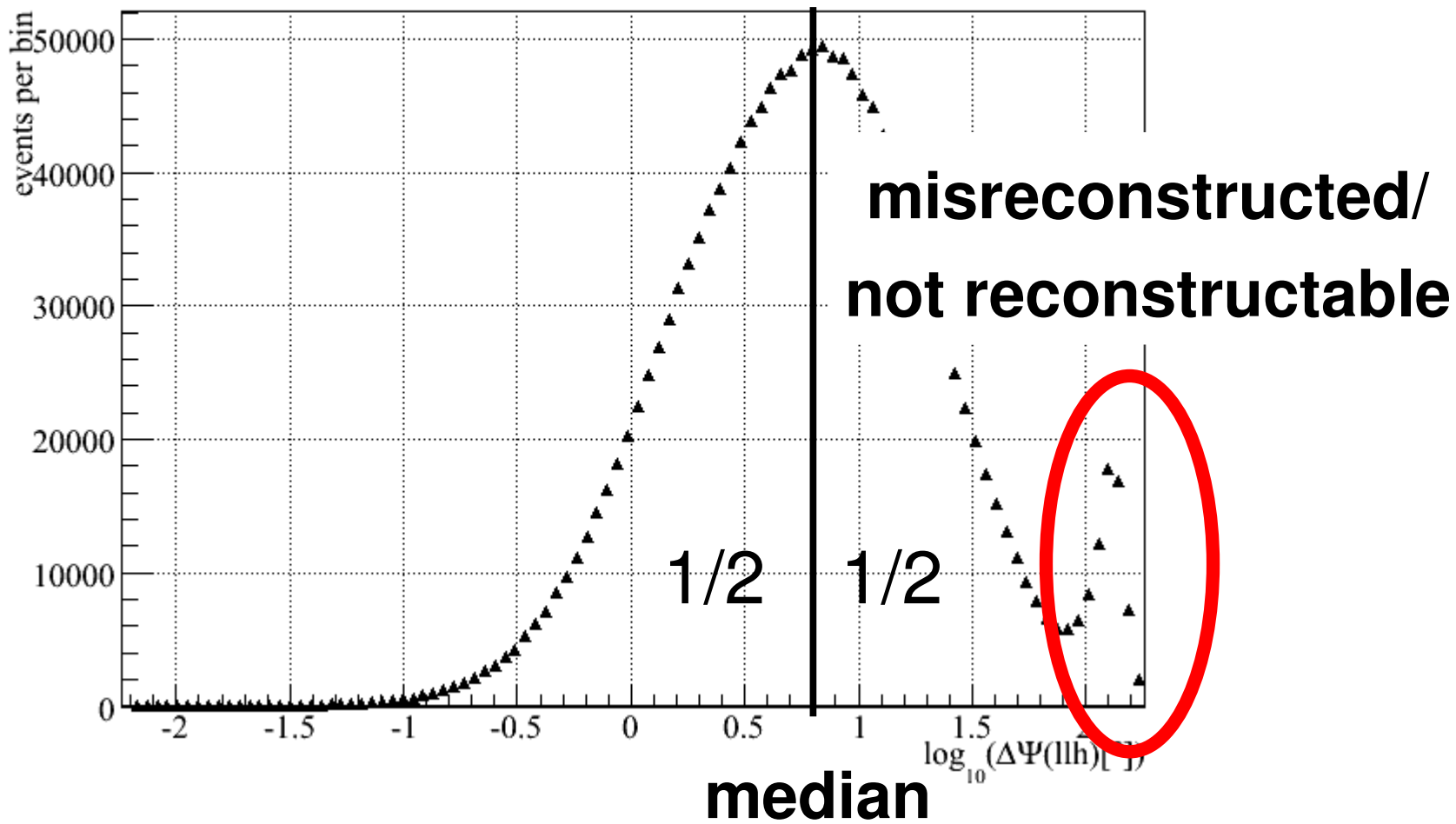


Actual



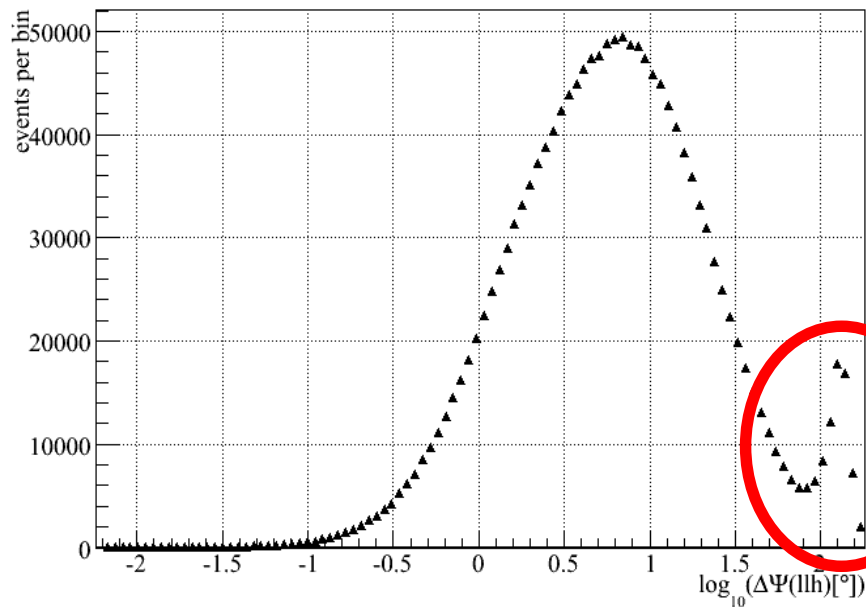


Point Spread Function

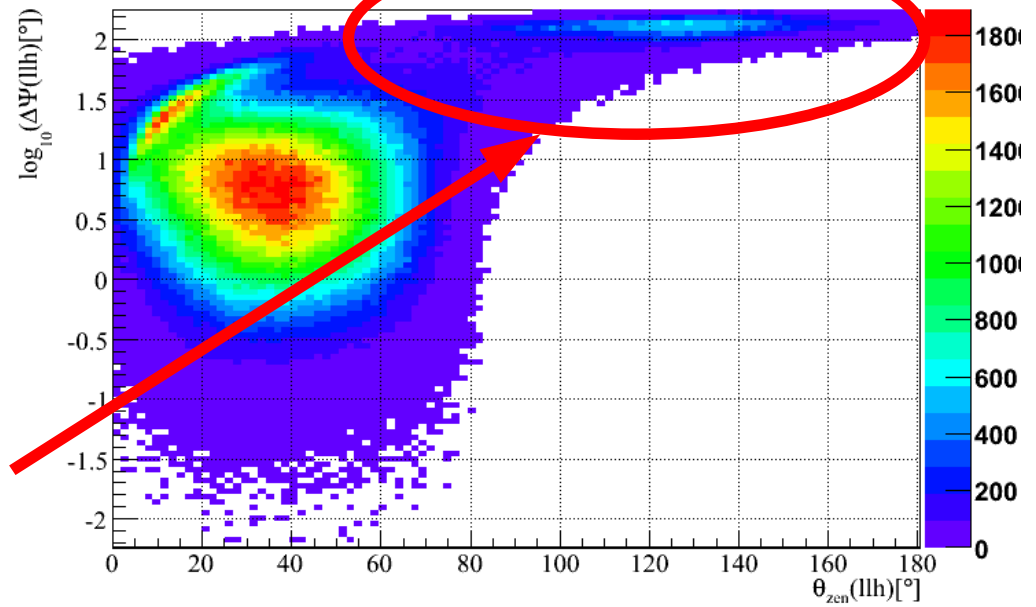


Raw

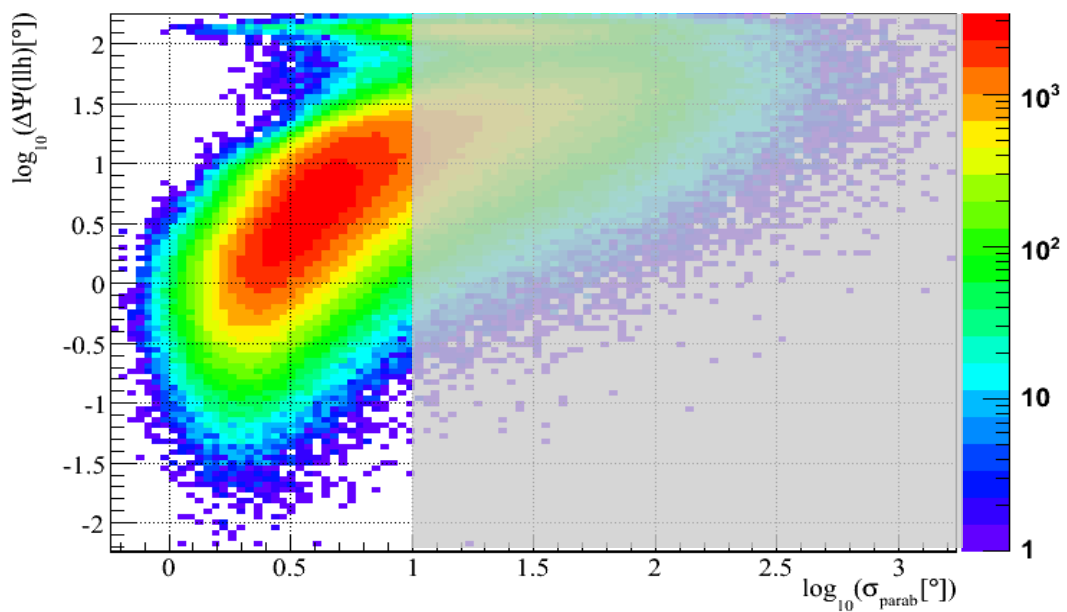
Point Spread Function



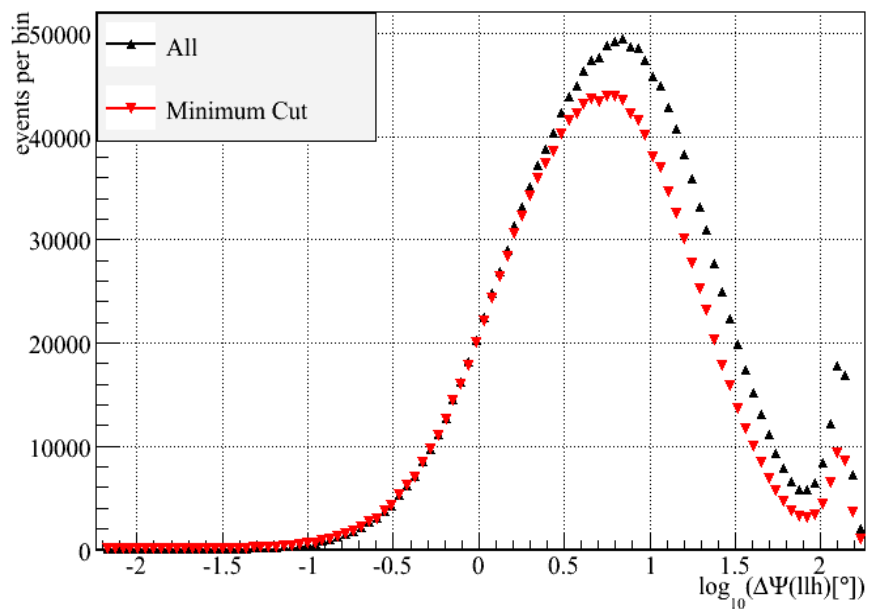
Zenith Angle and Point Spread Function



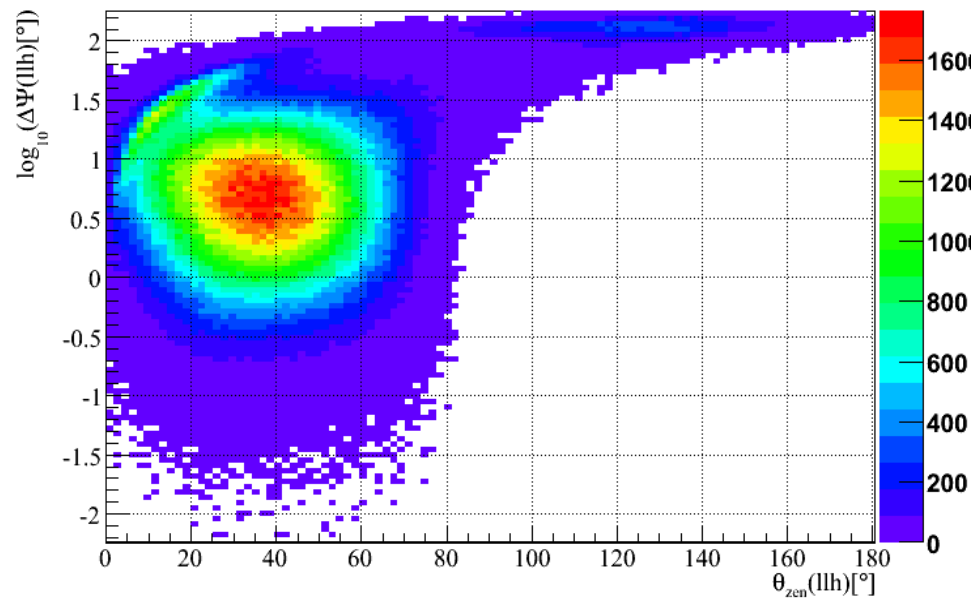
Paraboloid Error Ellipse and Point Spread Function



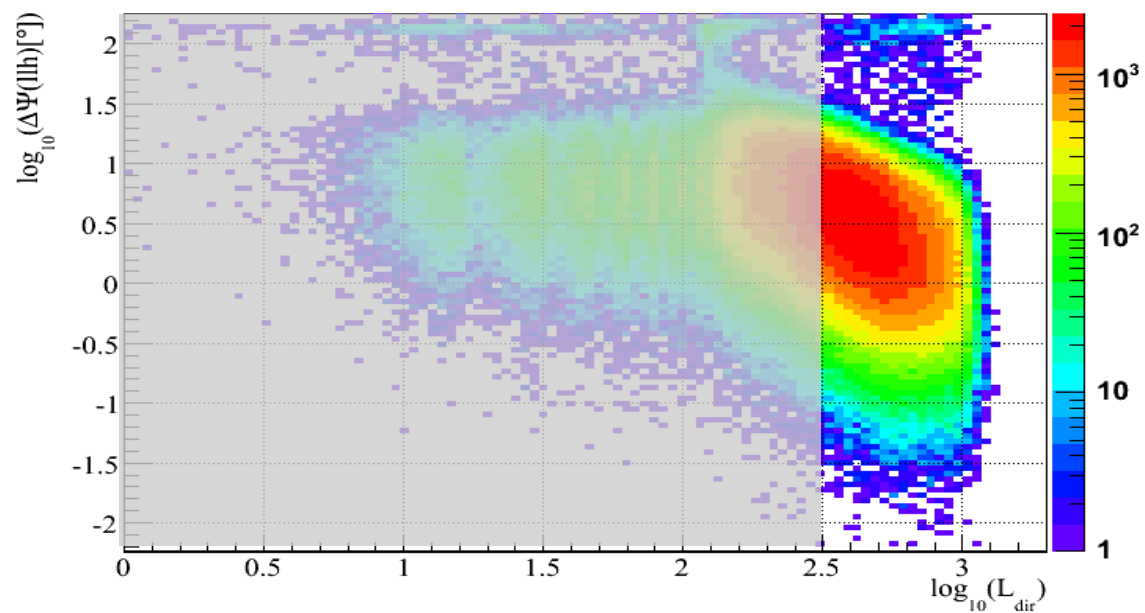
Point Spread Function



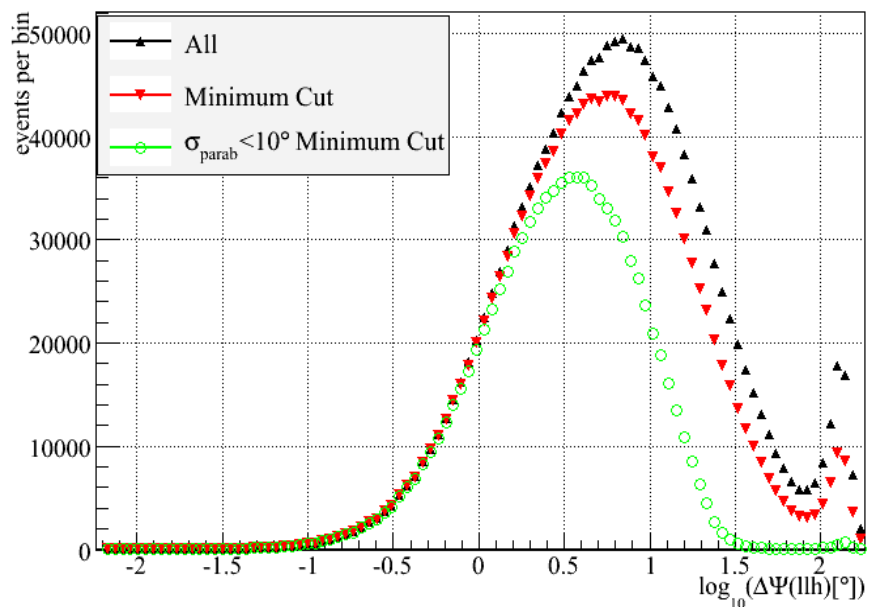
Zenith Angle and Point Spread Function



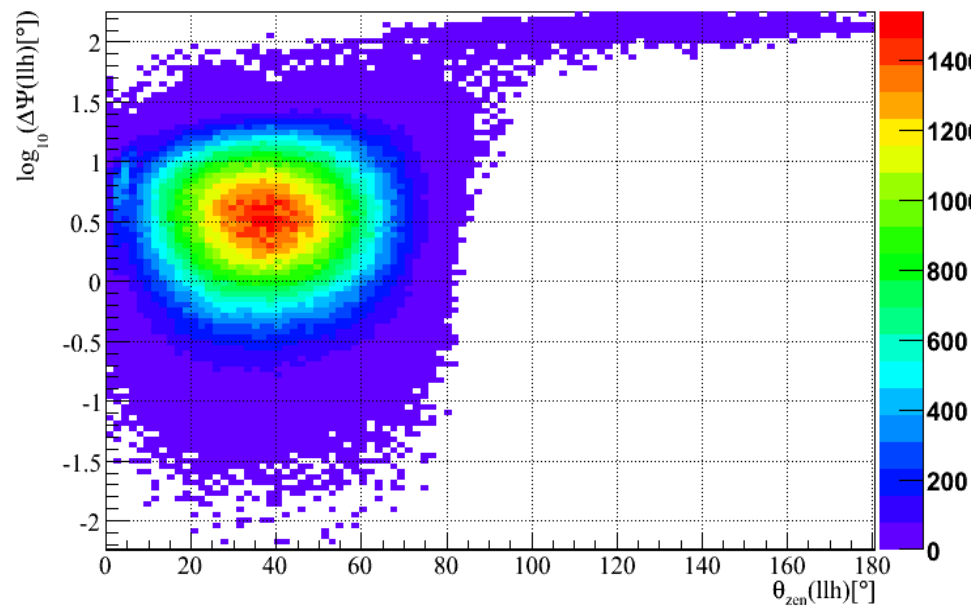
Track Length and Point Spread Function



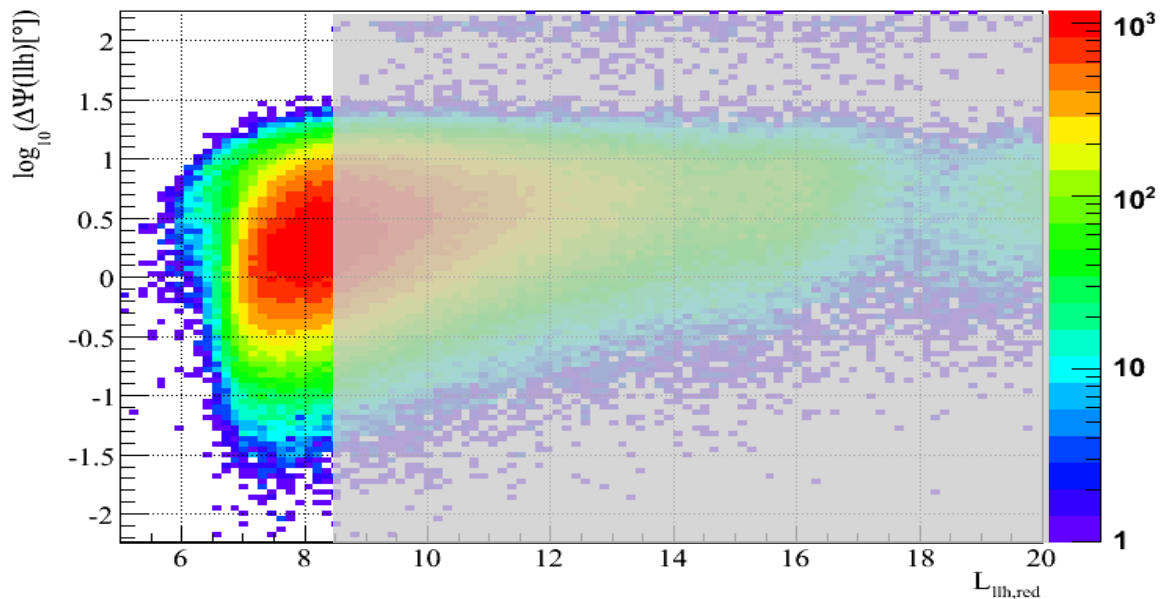
Point Spread Function



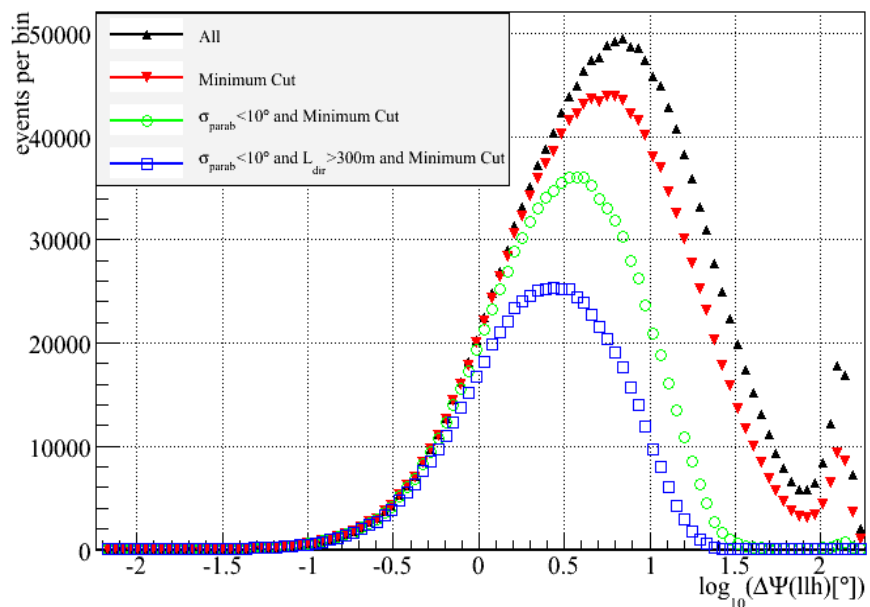
Zenith Angle and Point Spread Function



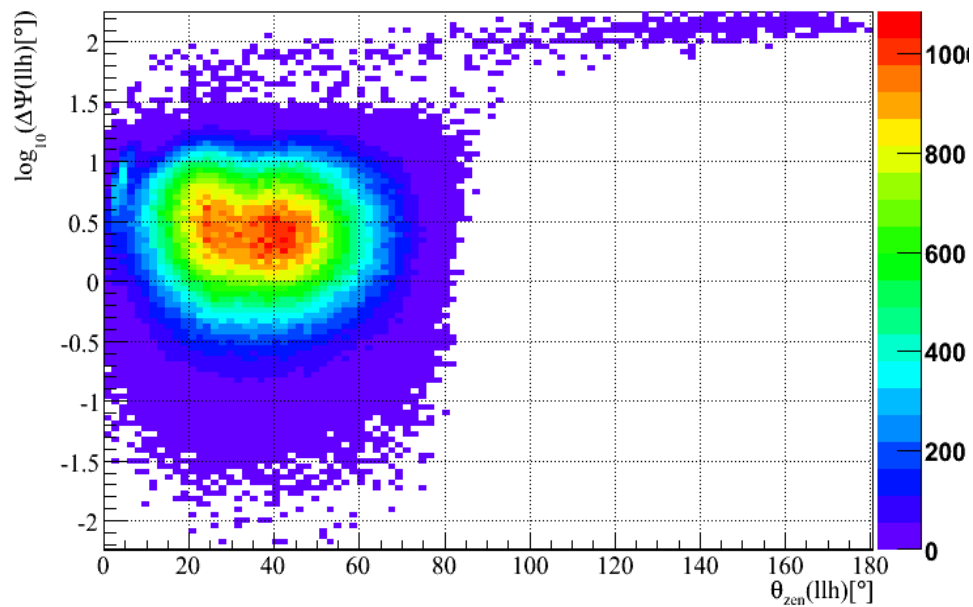
Reduced Likelihood and Point Spread Function



Point Spread Function

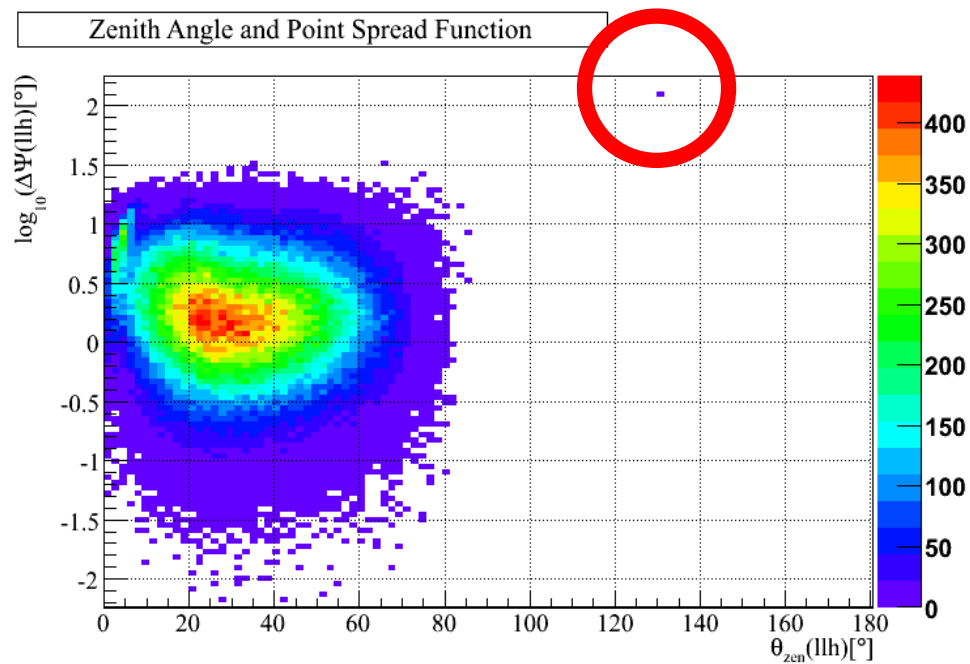
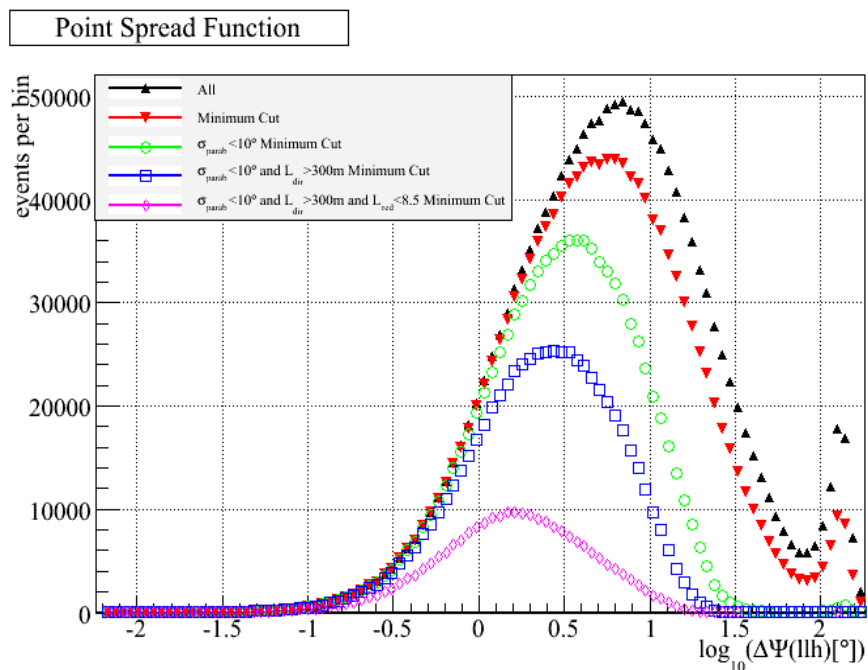


Point Spread Function and Zenith Angle

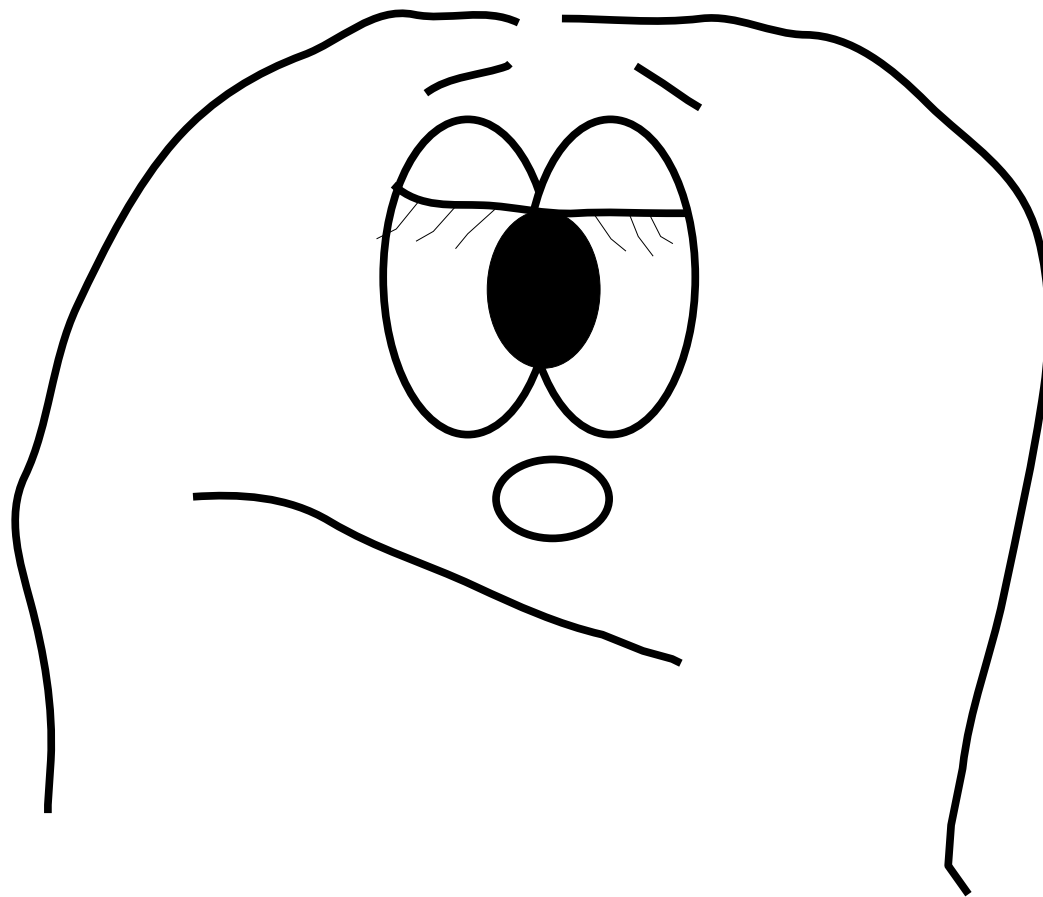


Final

SIGNAL!!!

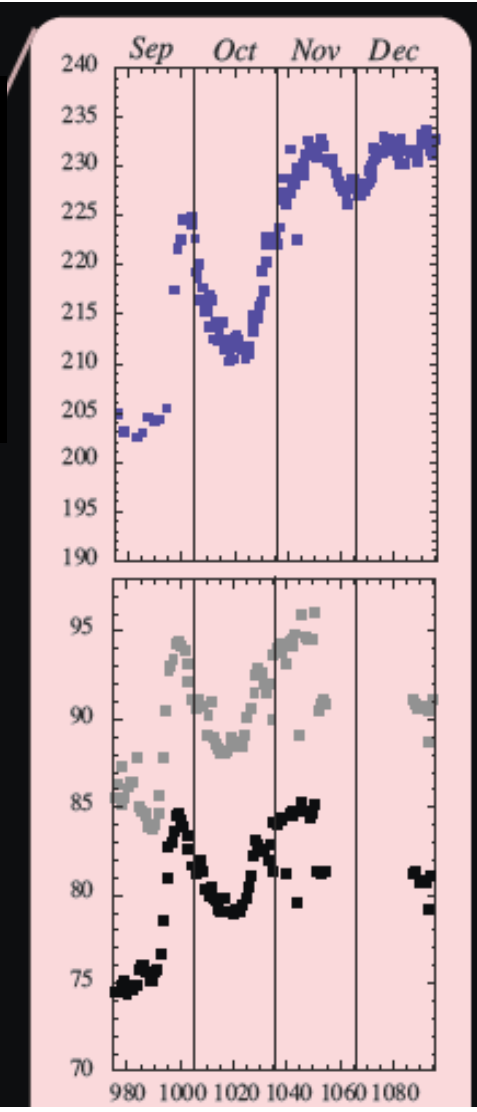
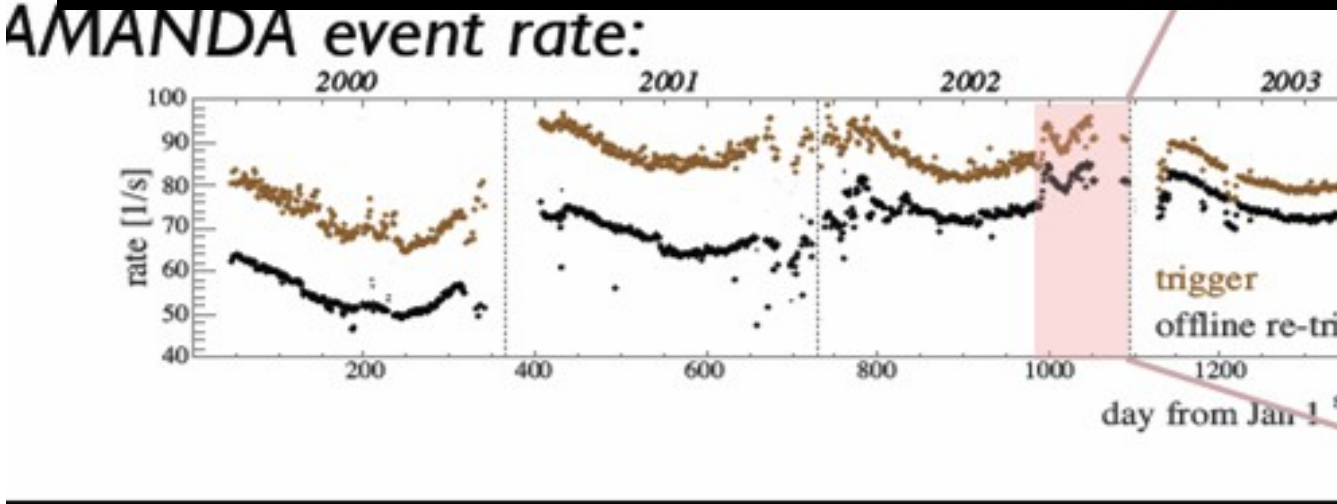


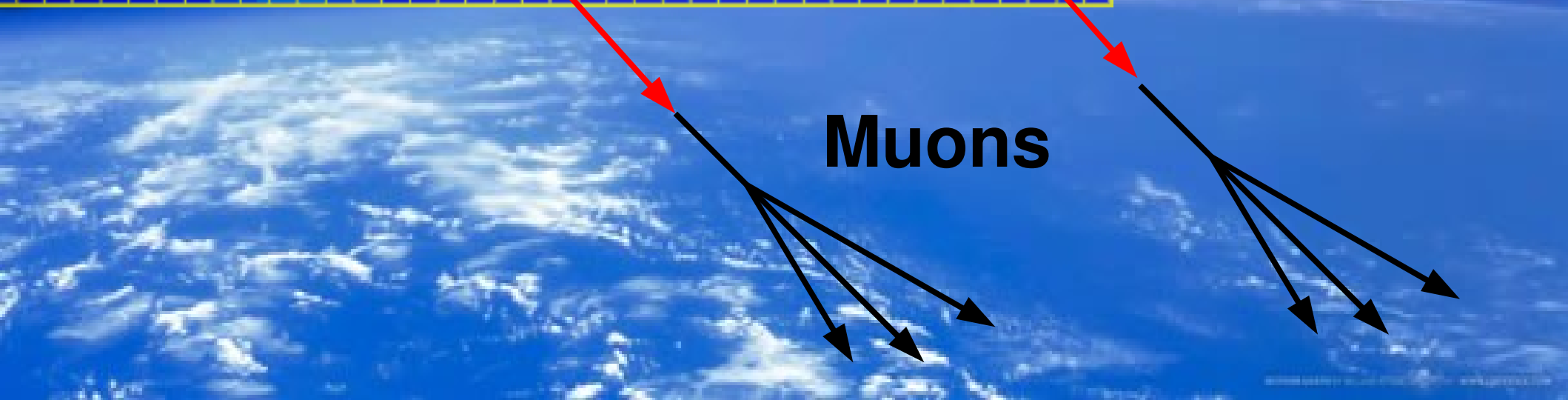
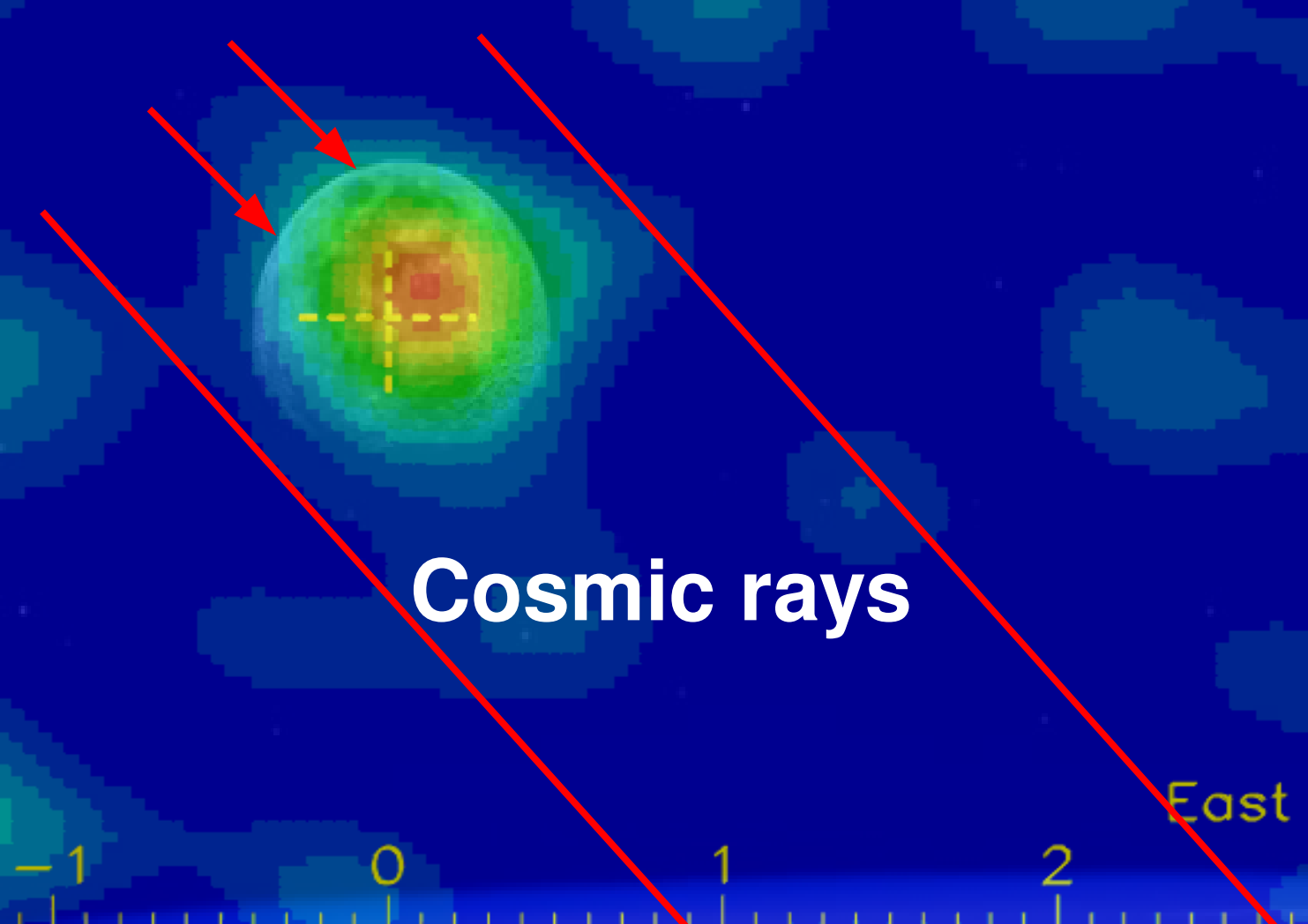
Applications



Omg!
Neutrinos!
Excitement!

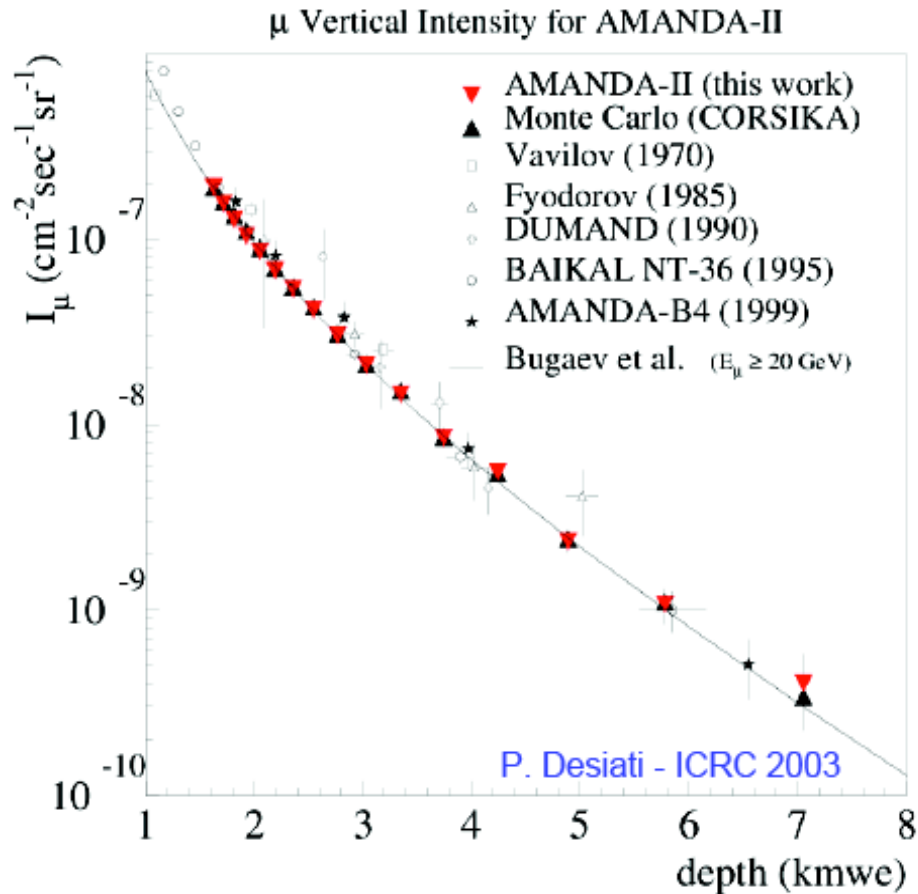
Meteorology





Muons

Muon Rate



Energy Spectrum

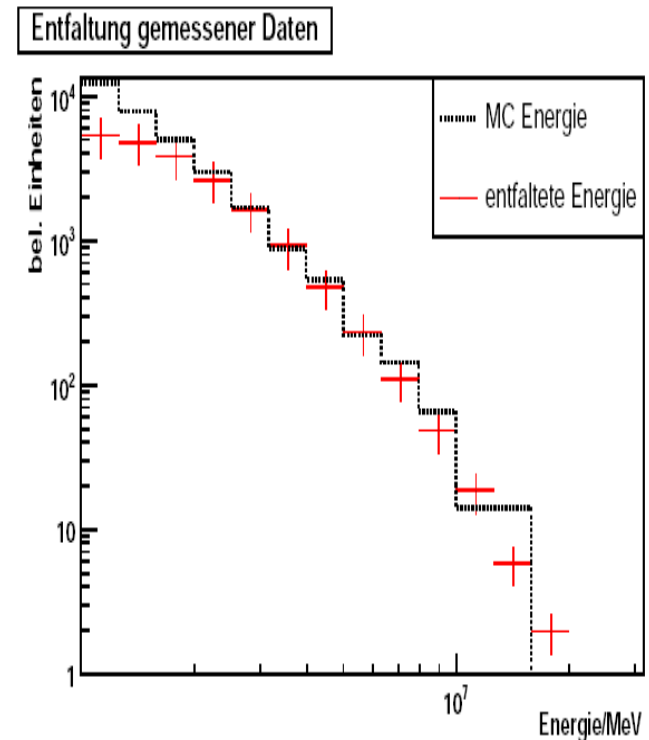


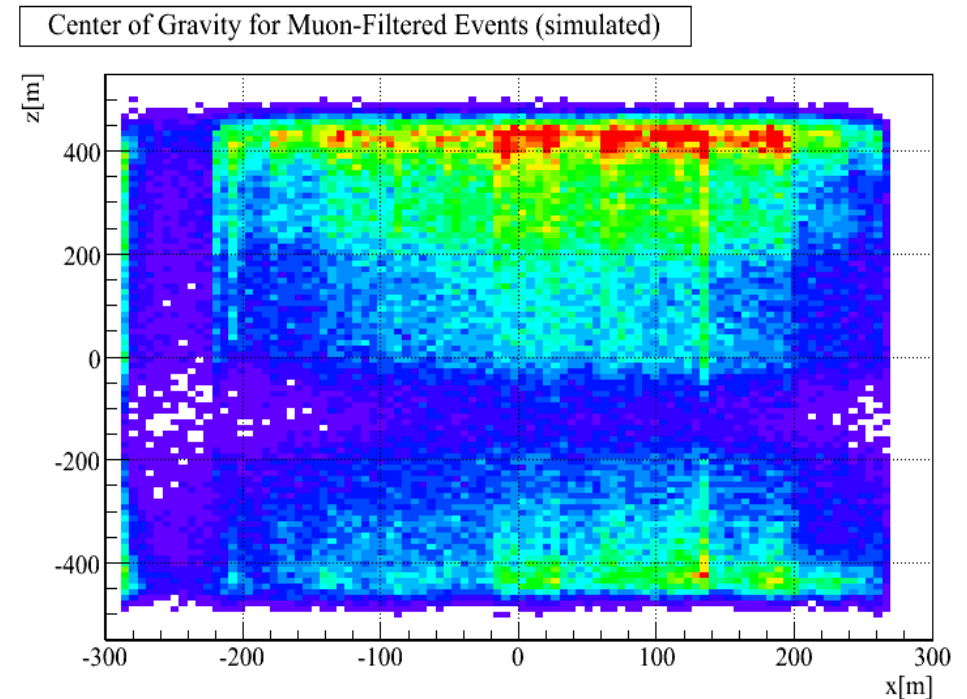
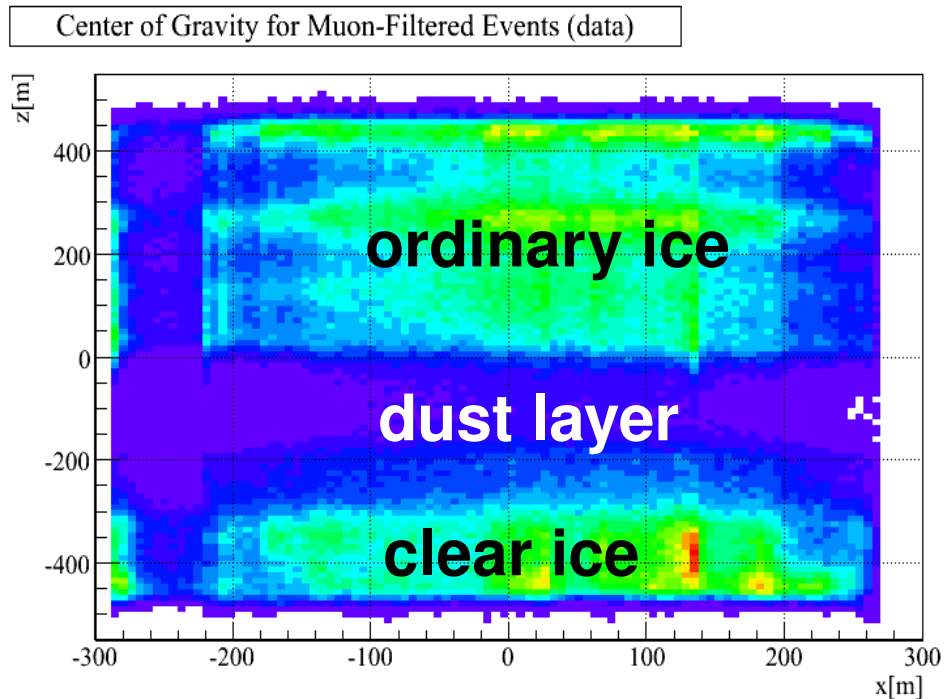
Abbildung 7.13: Entfaltung von gemessenen Daten aus dem Jahr 2000, verglichen mit simulierten Ereignissen (gestricheltes Histogramm)

AMANDA!

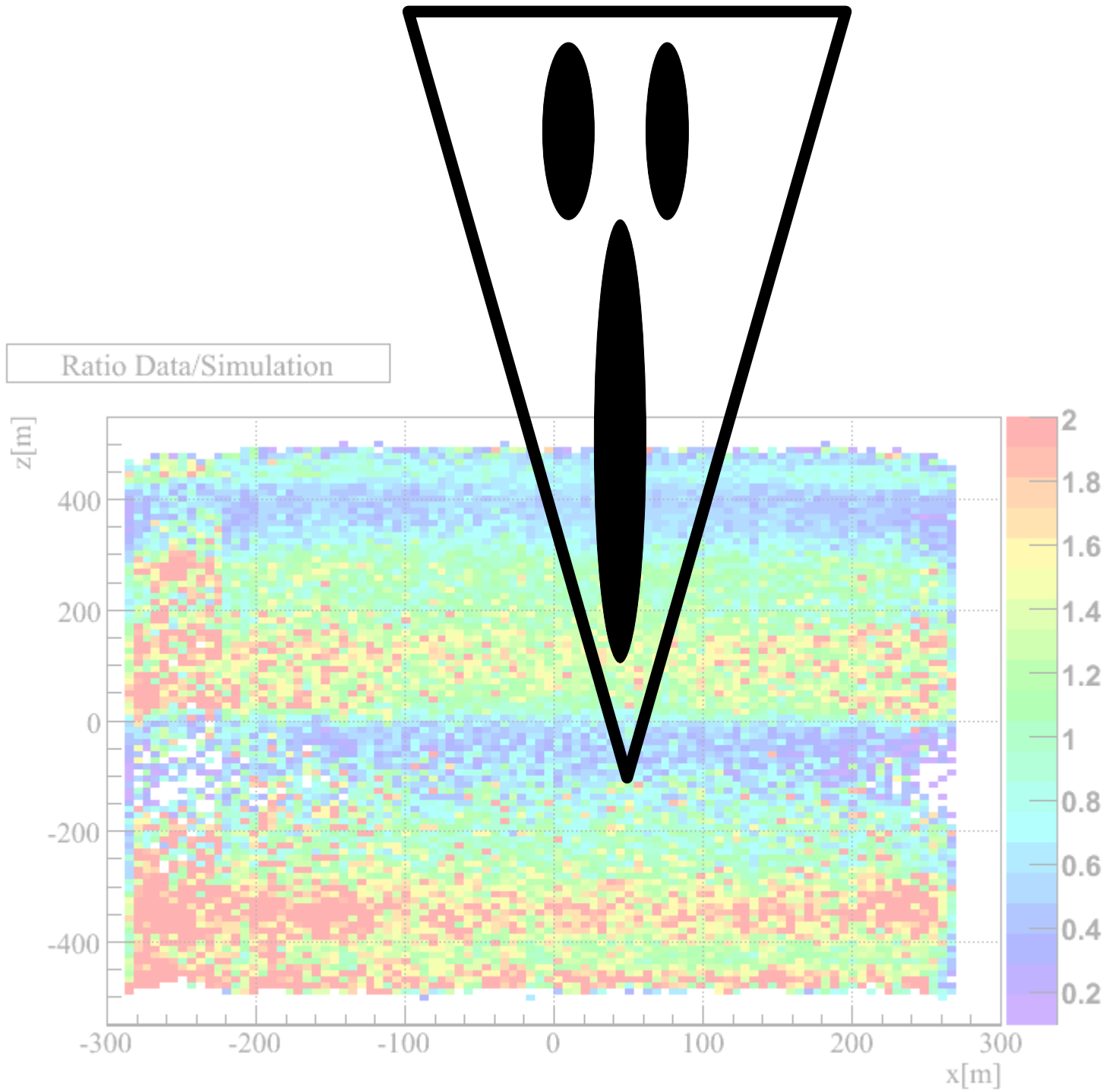
Detector Calibration

Real

Simulated



Near-Horizontal Muons
(below 70 degrees)



end