

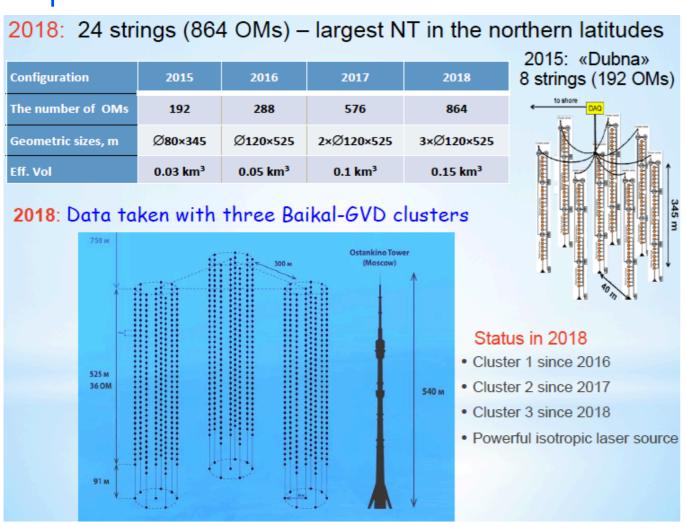


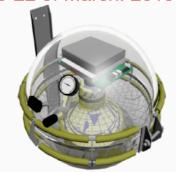
Baikal-GVD: first clusters

Details see Olga Suvorova @ XVIII International Workshop on Neutrino Telescopes - Venezia - 18-22 of March. 2019

- GVD detector construction underway in Lake Baikal
- Currently clusters #2 and #3 are in operation while cluster #1 is subject to maintenance works
- Baikal-GVD expedition on-going (deployments until April 11th)
 Plan was to deploy and commission two new GVD-clusters (clusters 4 and 5), well underway

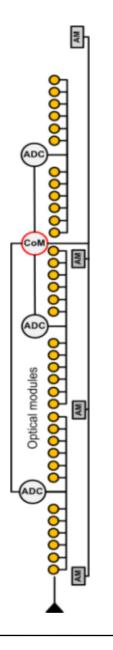


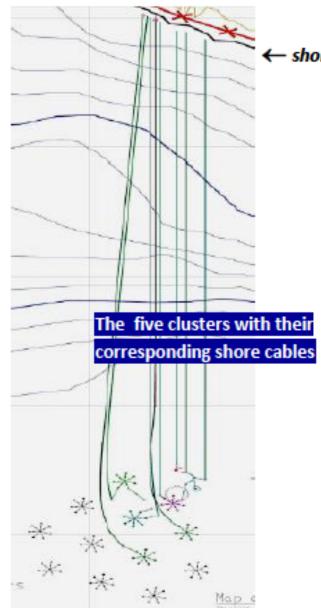




Optical module PMT: R7081-100



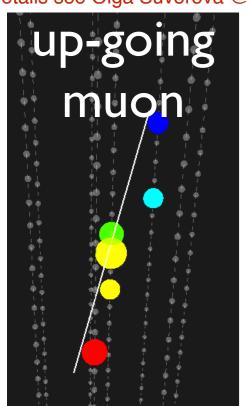


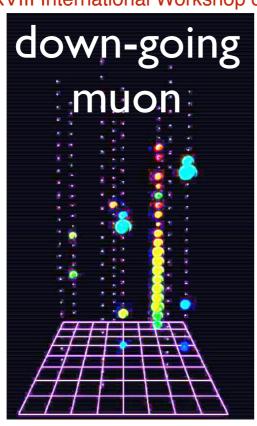


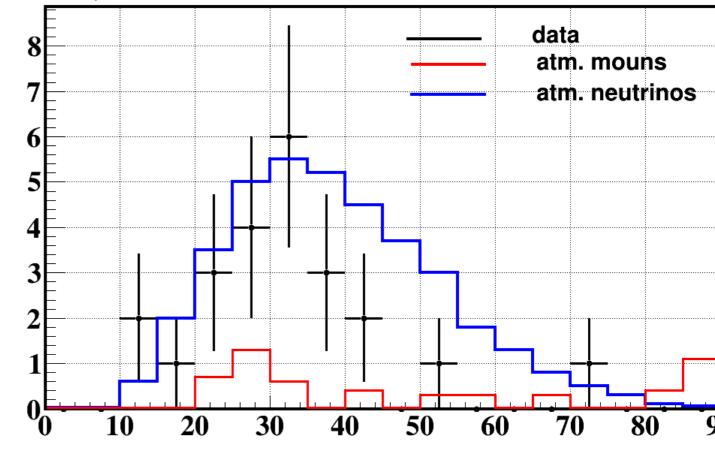


Atmospheric Neutrino Analysis at GVD

Details see Olga Suvorova @ XVIII International Workshop on Neutrino Telescopes - Venezia - 18-22 of March, 2019







- 33live days were analyzed of the first GVD data
- Event selection via BDT
- Expectations:
 - ~ 3 events estimation of atm. muons background
 - ~36 events estimation of signal atm. neutrinos
- Observation in data:
 - 23 events were selected in the signal region

Atmospheric Neutrino Physics starting at GVD

GVD-1 to reach 0.4 km3 by 2021 with 9 clusters and 2592 OMs

About 0.6 astrophysical events/year are expected per GVD cluster

 $\theta(deg)$

Baikal-GVD aims on search for astrophysical neutrinos

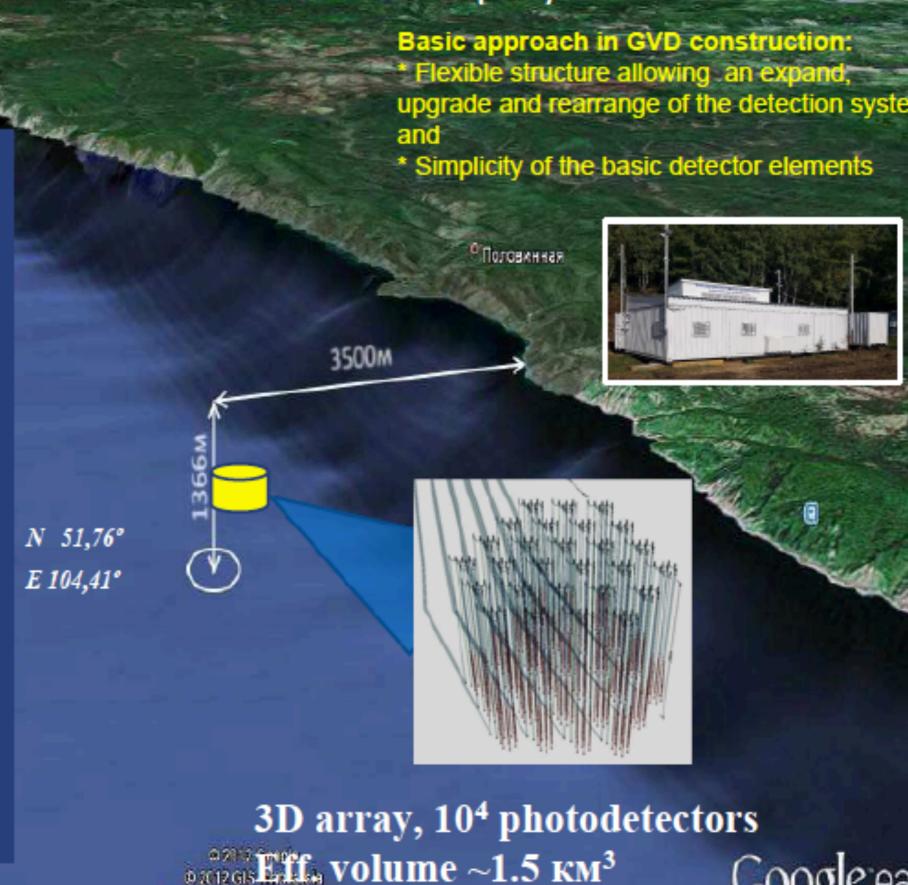
Tista INTAS Project 95-1569 Image Crottz TeraMétors

Details see Olga Suvorova @ XVIII International Workshop on Neutrino Telescopes - Venezia -18-22 of March, 2019

- 1370 m maximum depth
- Distance to shore ~4 km
- Absence of high luminosity from biology and K⁴⁰ background
- Water properties:
 Abs. length: 22 ± 2 m
 Scatt. length: L_s ~ 30-50 m
 L_s/(1- <cosθ>) ~ 300-500 m

Strongly anisotropic phase function: $<\cos\theta> \sim 0.9$

 Possibility to deploy the detector from the ice of the lake

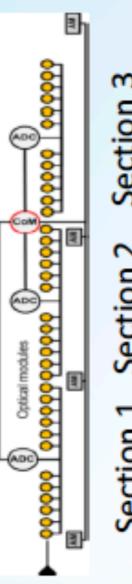


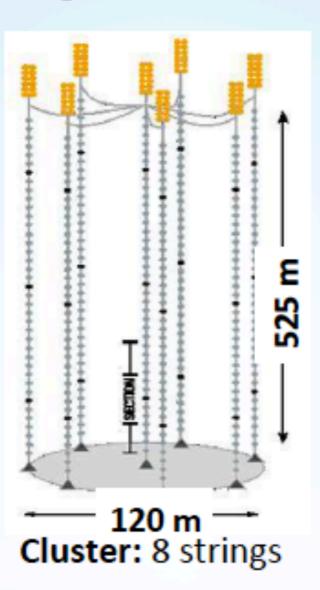
2768 m

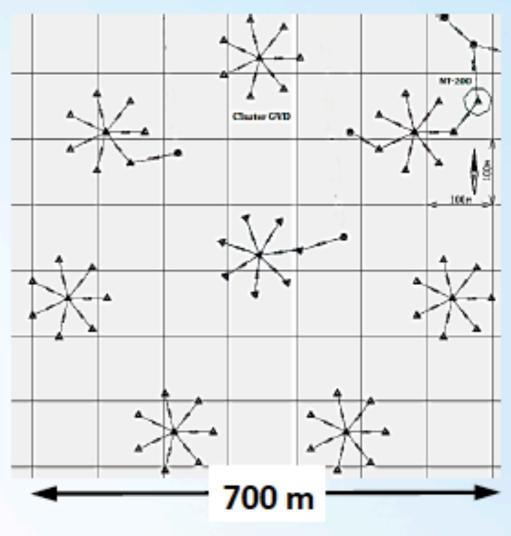
Baikal-GVD: phase 1 (2020-2021)



Optical module PMT: R7081-100







GVD-1: 8 clusters

GVD-1	
OMs	2304
Clusters (8 Strings)	8
Depths, m	750 - 1275
Eff. Volume	0.4 km ³

Directional resolution Energy resolution	
Cascades: 3.5° – 5.5°	δ(E/Esh) ~ 0.15
Muons: 0.25° - 0.5°	δ(lgE) ~ 0.4

Details see Olga Suvorova @ XVIII International Workshop on Neutrino Telescopes - Venezia - 18-22 of March, 2019 XVIII INTERNATIONAL WORKSHOP ON NEUTRINO TELESCOPES I VENEZIA I 18-22 OF MARCH, 2019







NEW ENTRY ON THE NEUTRINO MAP - @ONC

BRAINSTORMING AROUND A SEGMENTED DETECTOR FOR HE HORIZONTAL TRACKS

→ STARTING UP CONCEPTUAL DESIGN

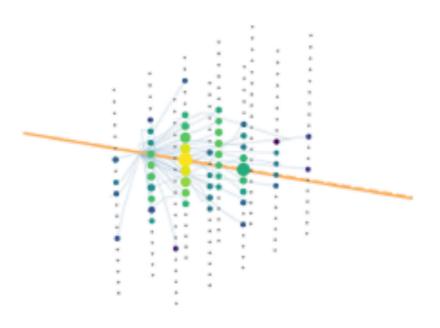
rectangular bundle

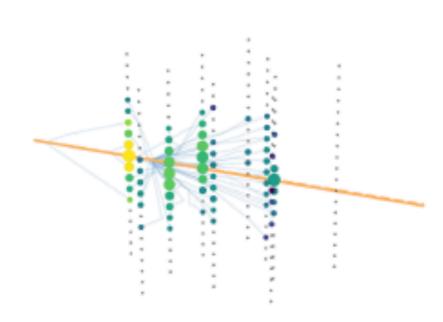
Water model from Antares

 $E_{\nu_i} = 50 \text{ TeV}, E_{J, \text{vertex}} = 28 \text{ TeV}$

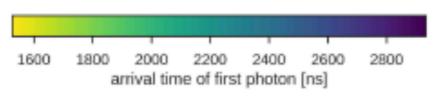
pentagonal bundle

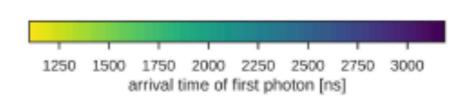
 $E_{\nu_i} = 50 \text{ TeV}, E_{l, \text{ vertex}} = 28 \text{ TeV}$





Study on going from K. Krings (TUM)





22



- STRAW STRings for Absorption length in Water
 - Pathfinder mission towards a possible large scale neutrino telescope
 - Deployed in June 2018 at the Cascadia Basin site operated by Ocean Networks Canada (ONC)
 - depth of about 2600meters
 - two STRAW120 meters tall mooring lines
 - instrumented with 3 POCAMs and 5 sDOMs

