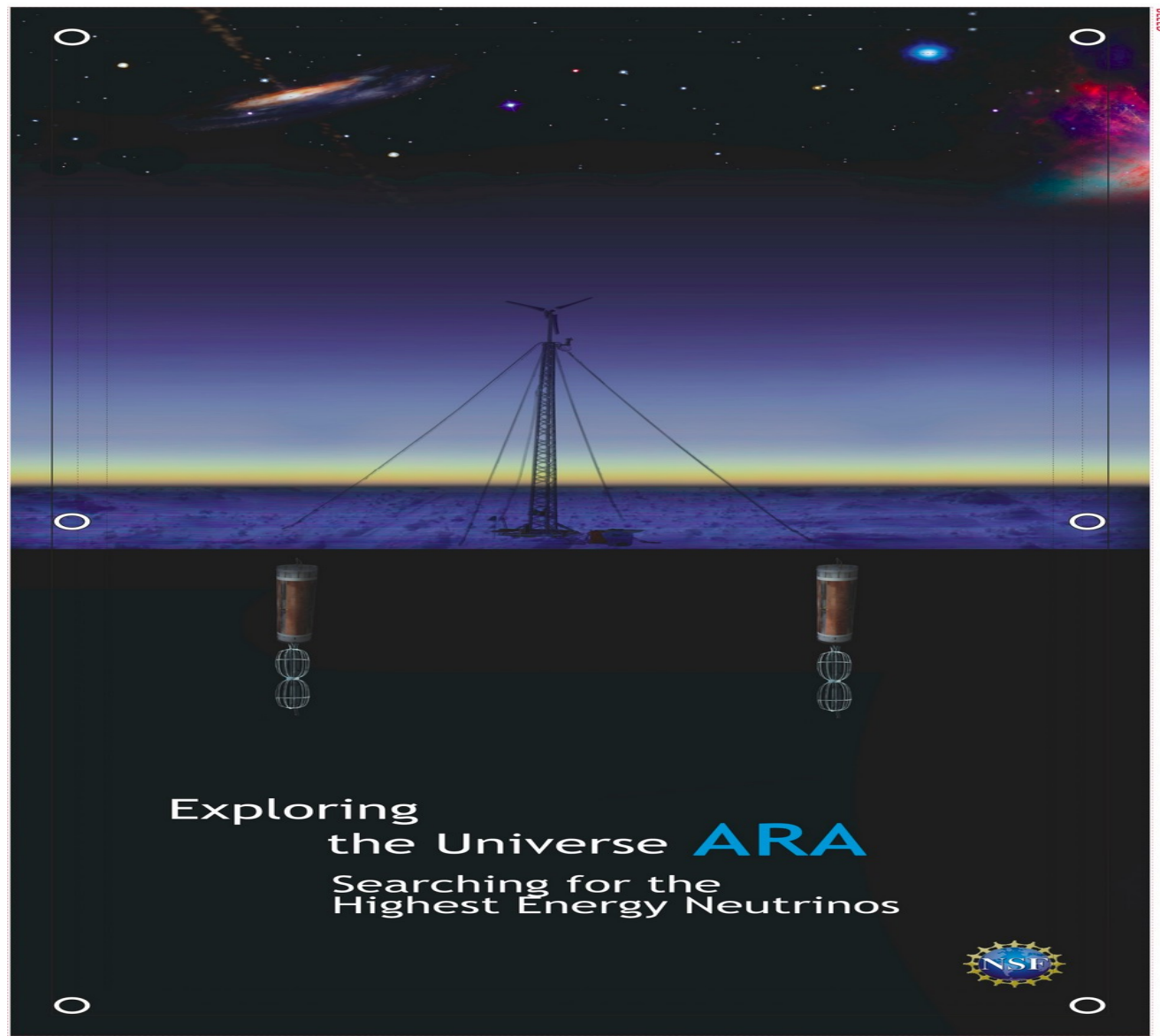
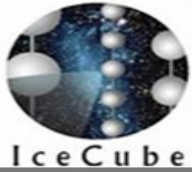


# Radio-wave detection of neutrinos

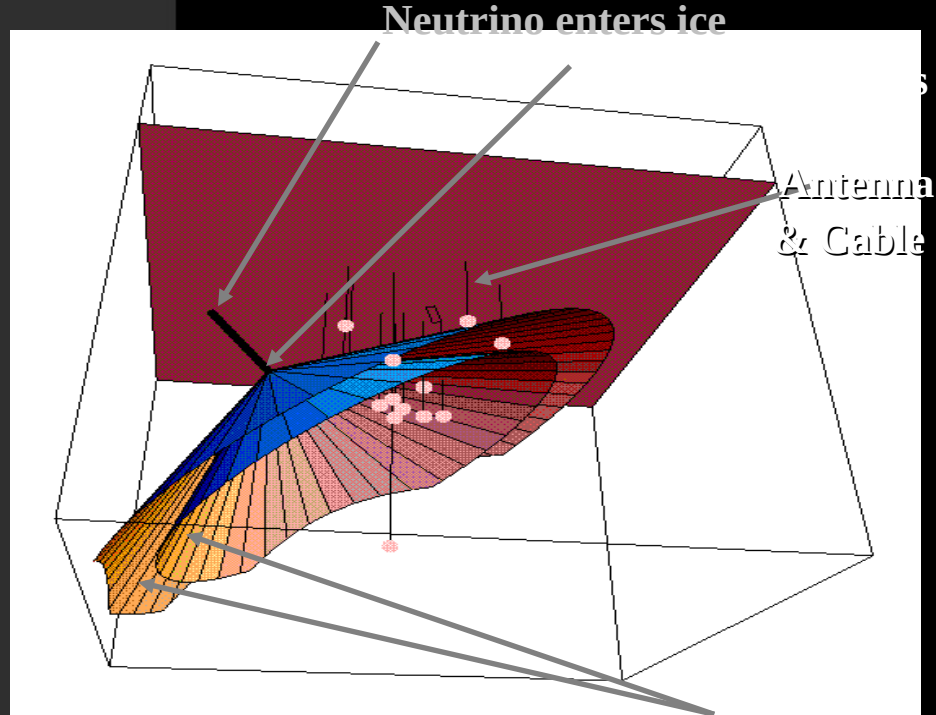
## Current Antarctic Experiments: ARA (South Pole) & ARIANNA (Moore's Bay)

UHE  $\nu$ :  
ARA,  
ARIANNA,  
ANITA



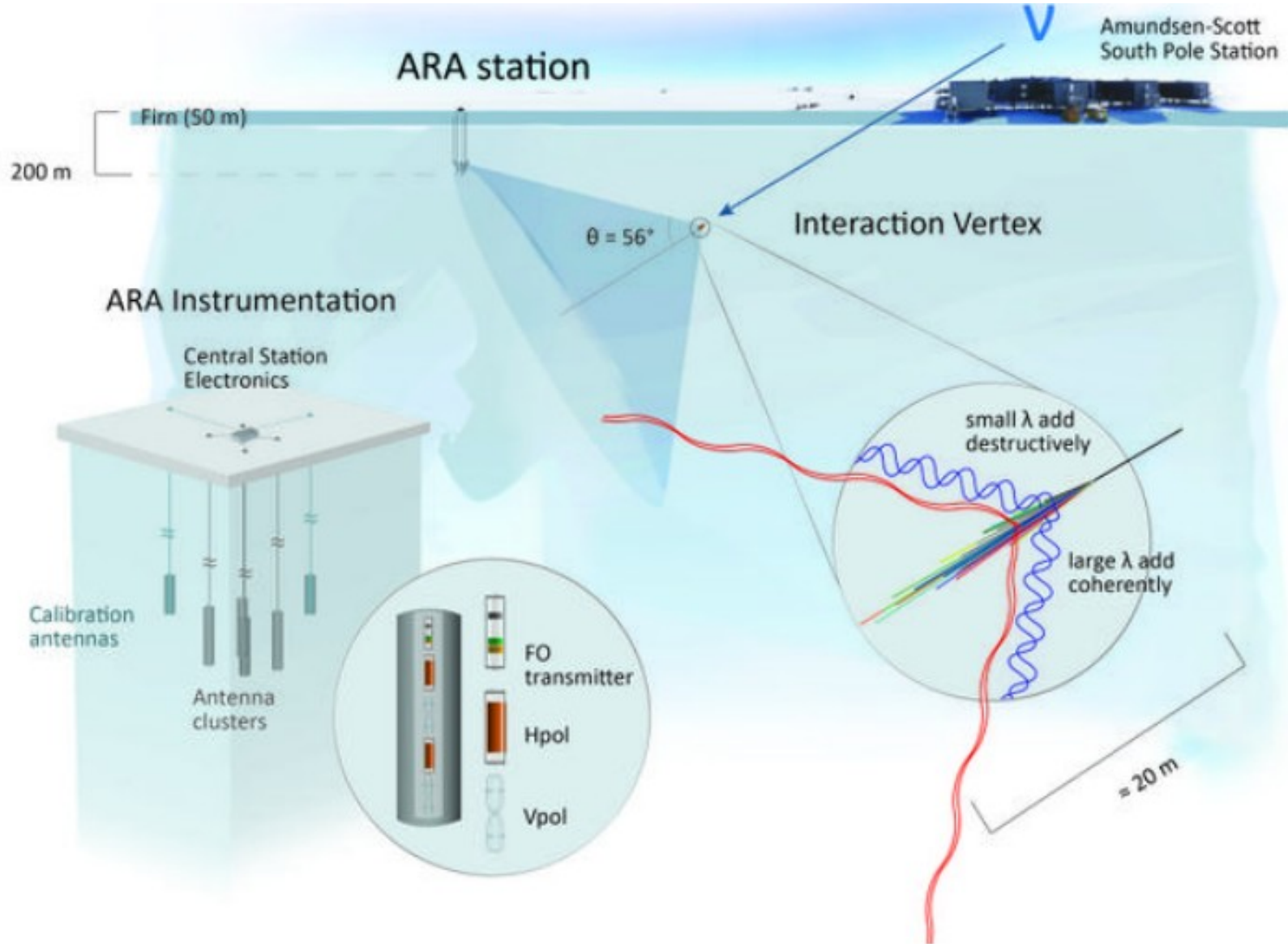


# Radio Detection in South Pole Ice

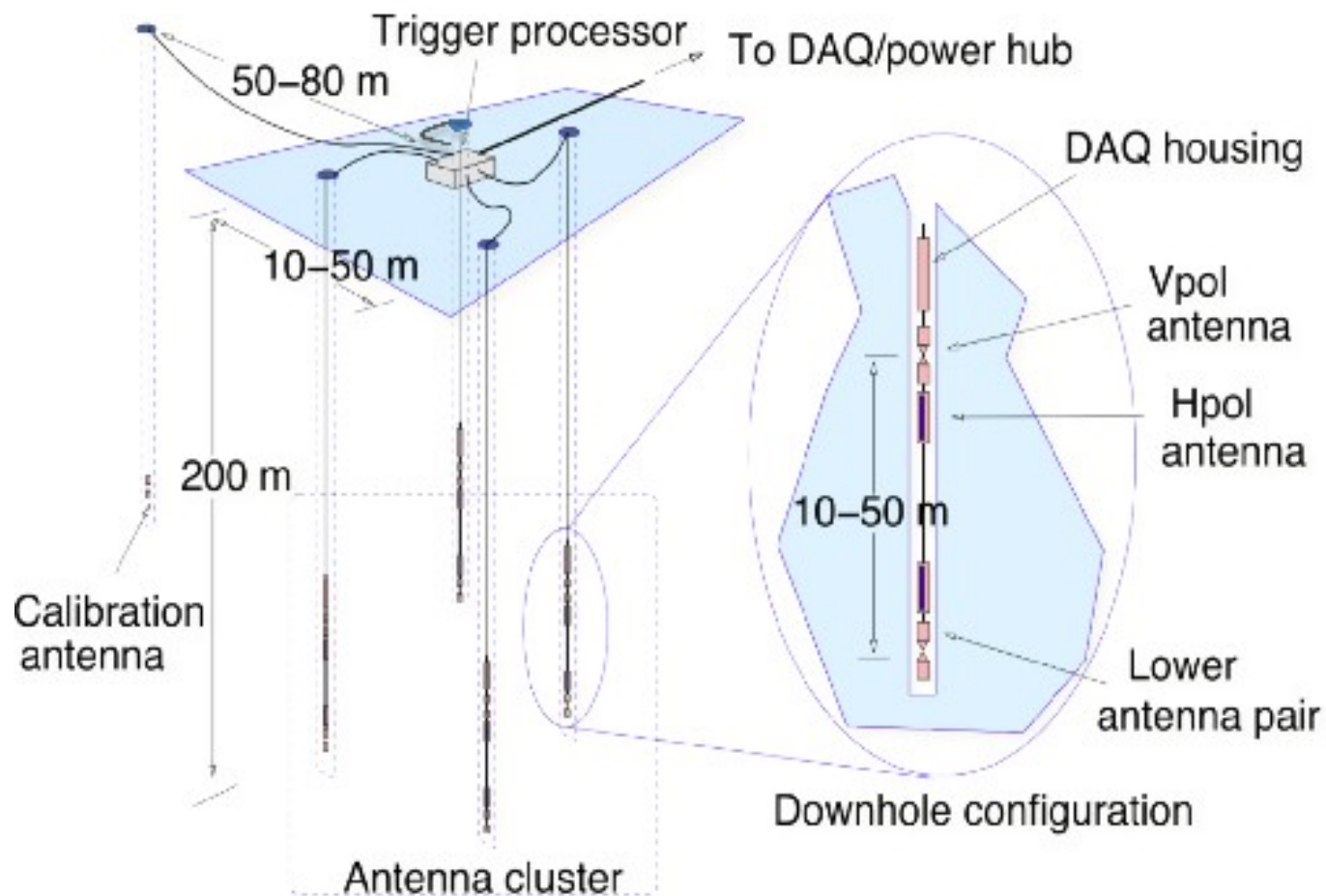


Two cones show 3 dB  
signal strength

- installed 15 antennas  
few hundred m depth with  
AMANDA strings.
- tests and data since 1996.
- most events due to local  
radio noise, few candidates.
- Askaryan Radio Array  
near IceCube : first  
deployments December 2010



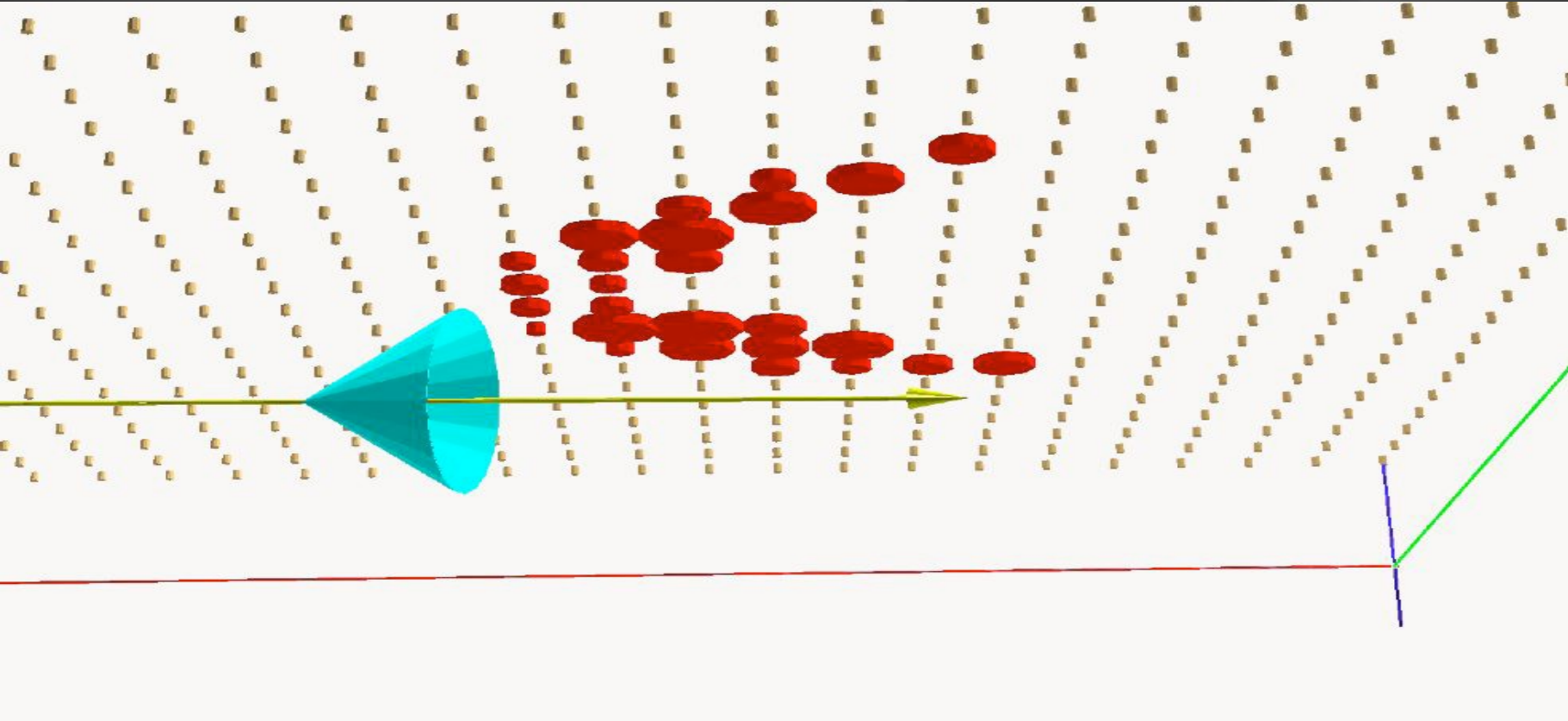
## ARA Station & Antenna Cluster



in-ice view of radio detection with antennas

→ 200m deep for ARA

→ at the surface for ARIANNA



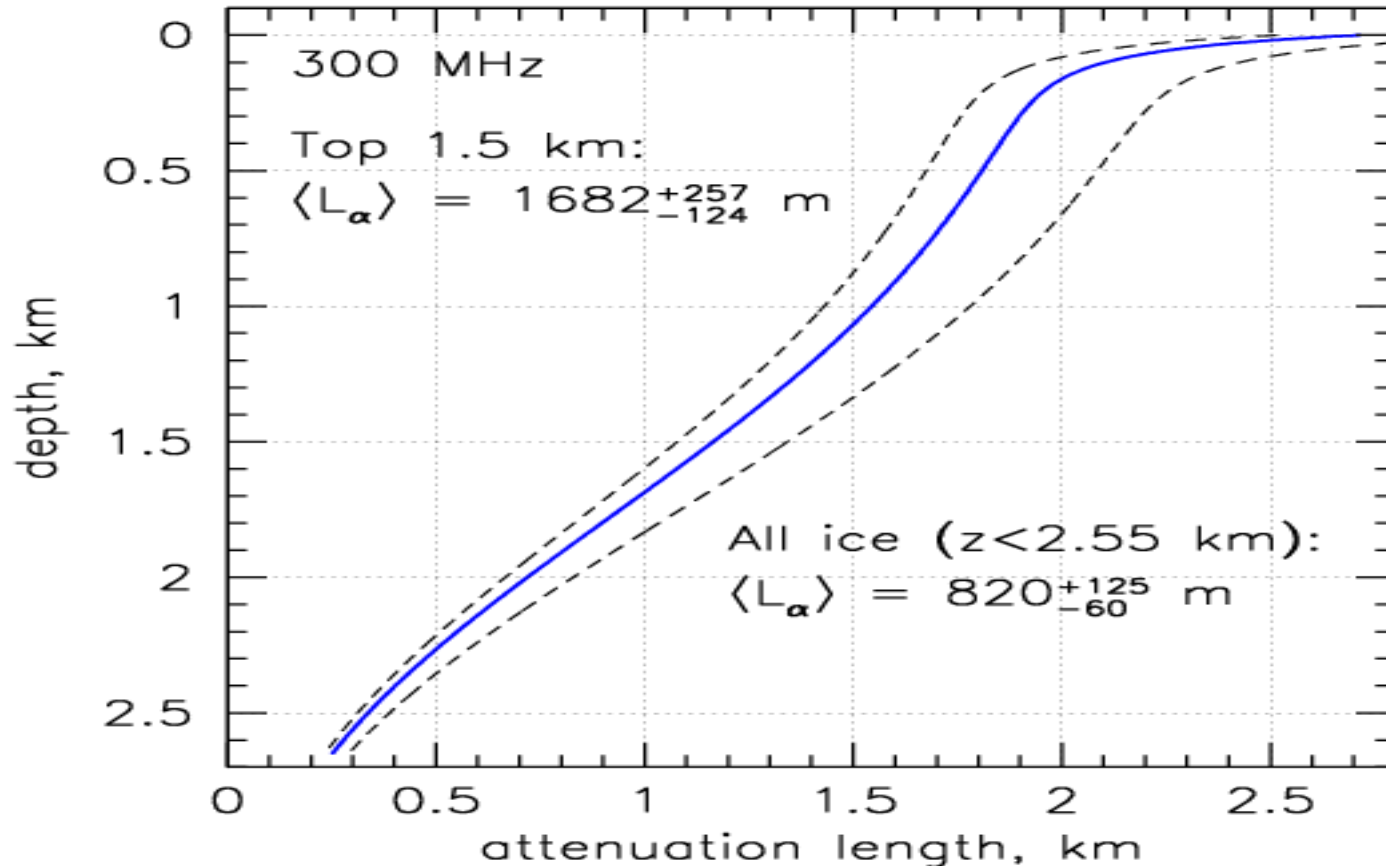
# Deployed Antennas (Jan 2011); wire bicones and bicone slot-cylinders

NOTE: Cable runs through center to preserve azimuthally symmetric response



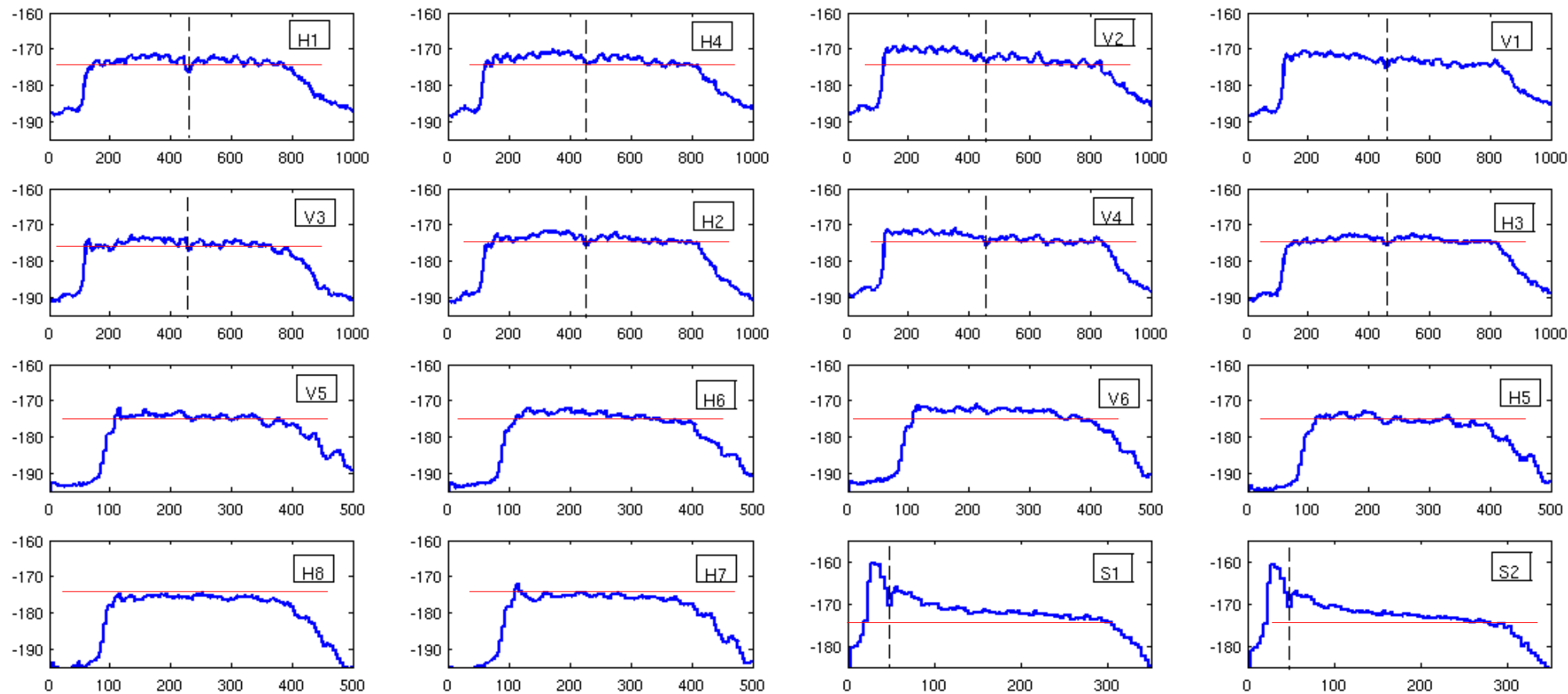
Analyze signals received @testbed from Buried Pulsar:  $\Delta z=2450$  m;  $\Delta r \sim 2000$  m  
 Calculate  $\langle L_{\text{atten}} \rangle$  via:  $P_{\text{rx}} = P_{\text{Tx}} G_{\text{Tx}} G_{\text{Rx}} \frac{\lambda^2}{4\pi r^2} \exp(-2r/L_{\text{atten}})$ ; Unfold depth dependence via lab  
 $L_{\text{atten}}(T(z))$ , using IceCube  $T(z)$

## Attenuation length (depth)



# Measured Noise floor=-174 dBm/Hz+2 dBm/Hz system

Average Noise Power, dBm/Hz



**(resistive losses in QSC)**

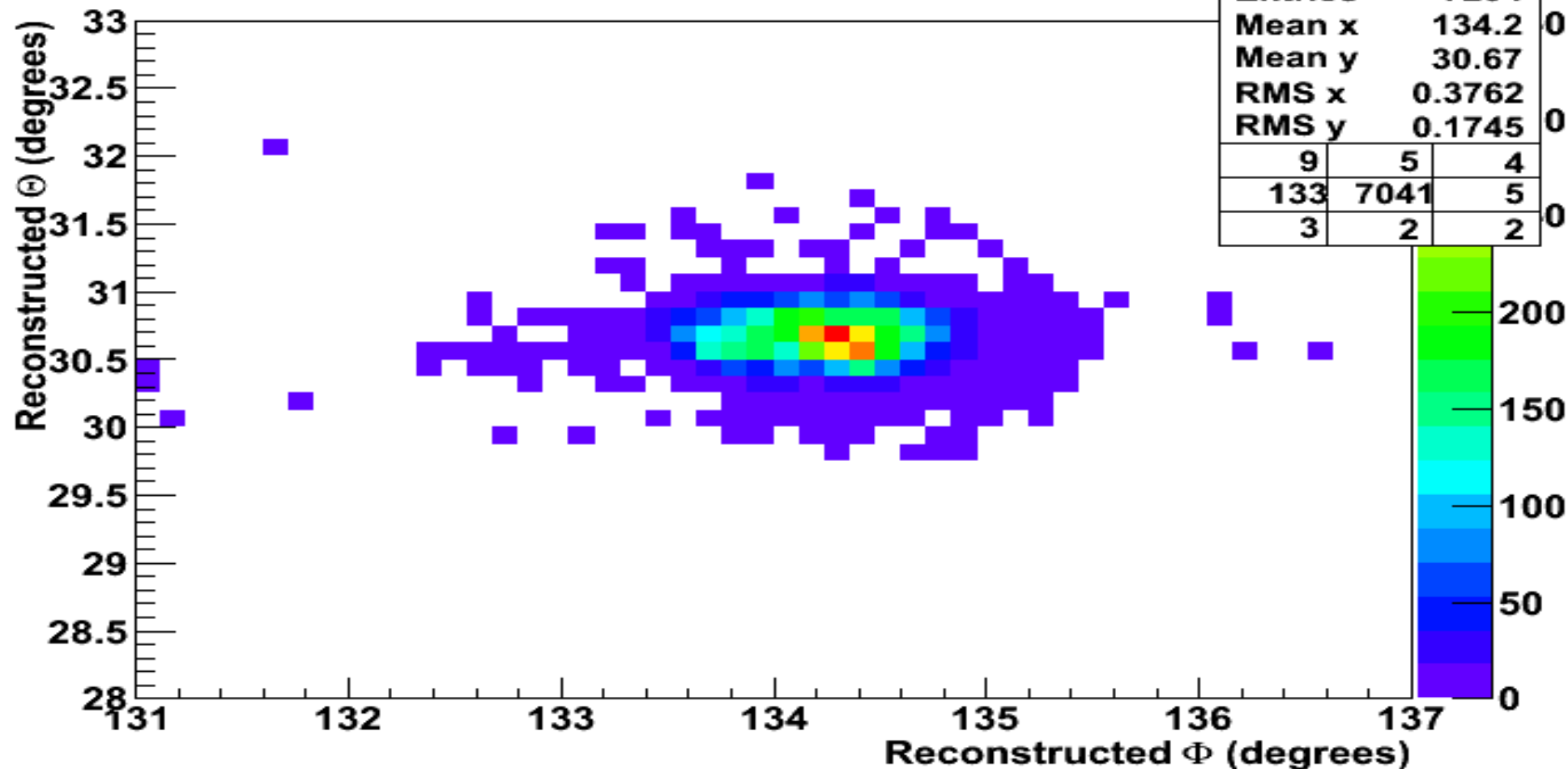
Frequency, MHz

# Thermal noise from a source at temperature T

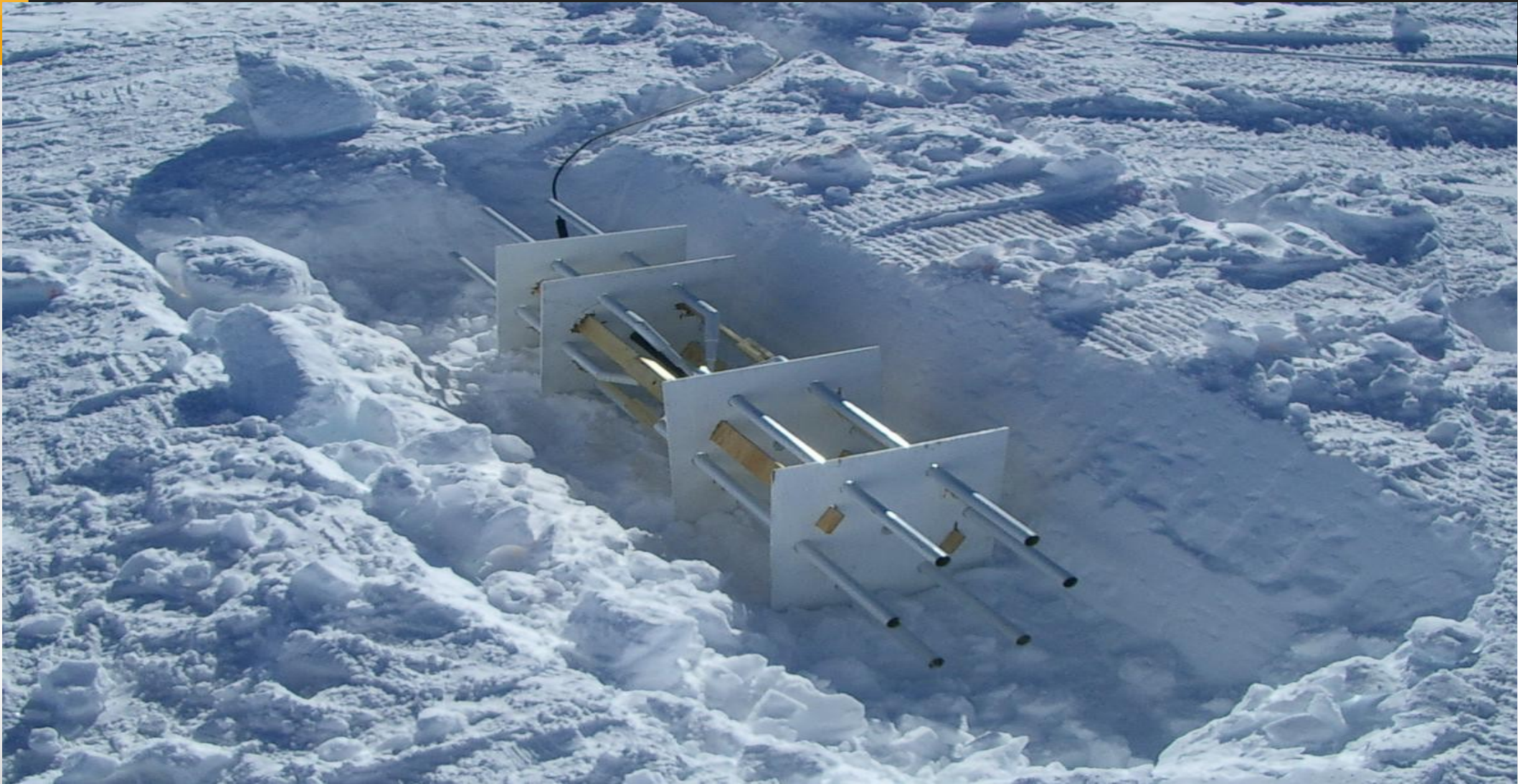
- Noise power =  $k_B T (\text{°K}) \text{Bandwidth (Hz)}$ 
  - For one Hz,  $kT = 3.31 \times 10^{-21} \text{ W} = 3.31 \times 10^{-18} \text{ mW}$ 
    - $= 10 \log_{10}(3.31 \times 10^{-18}) \text{ dBm} = -174.8 \text{ dBm}$
- Typical numbers:
  - 500 MHz bandwidth
  - Temperature = 240 K
  - $P = 1.38 \times 10^{-23} \times 5 \times 10^8 \times 240 = 1.66 \times 10^{-12} \text{ Watts}$
  - Coupled to 50Ω antenna:  $V^2/50 = 1.66 \times 10^{-12}$ 
    - $\Rightarrow 10 \text{ microVolts}$

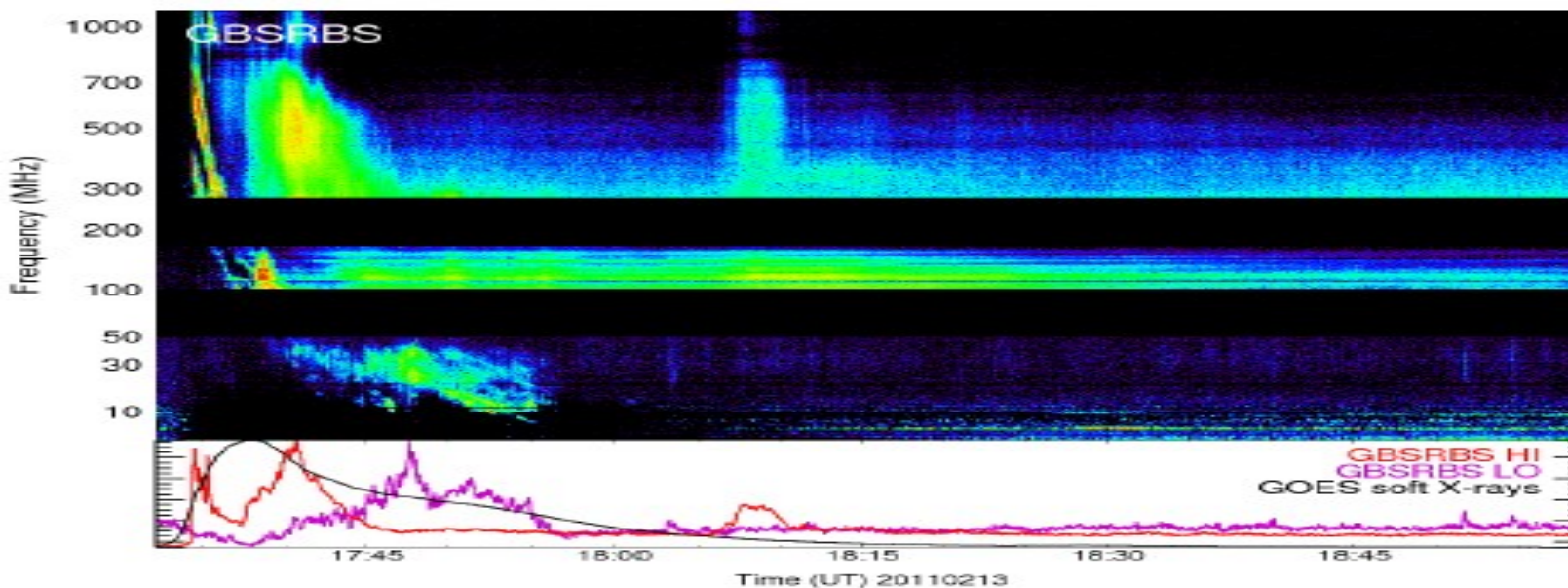
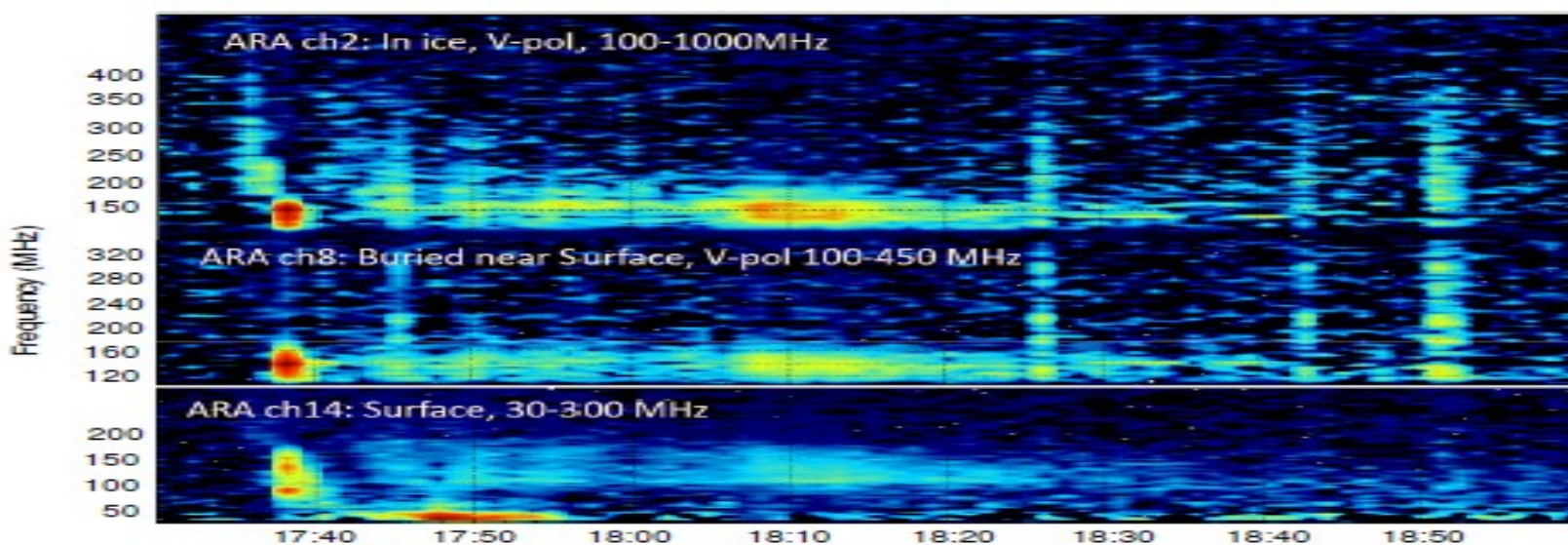
# ARA angular resolution $\sim 22'/10'$ in $\theta/\phi$

Calibration Pulsar Source Reconstruction



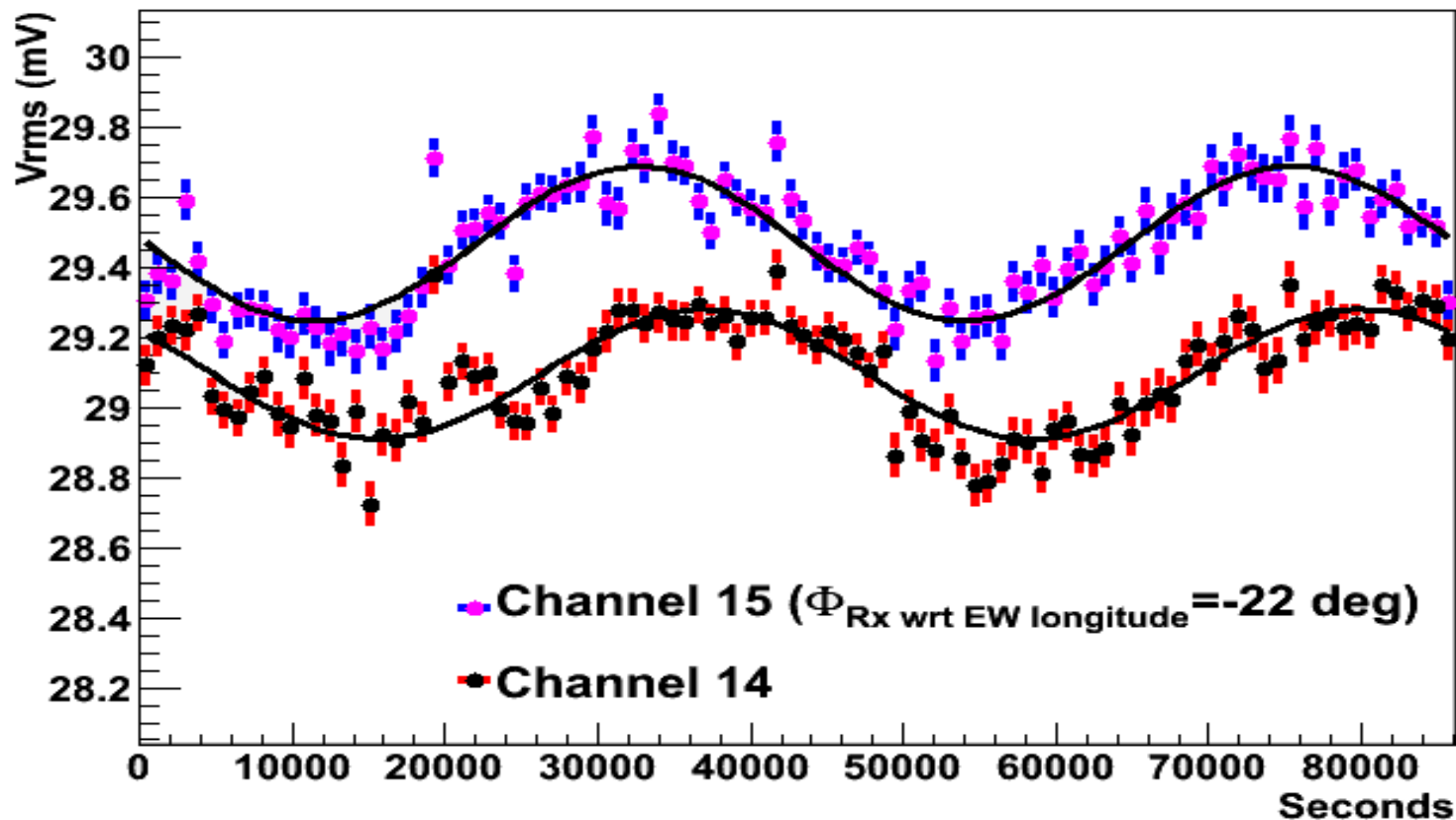
# Surface Antennas - “fat dipoles”





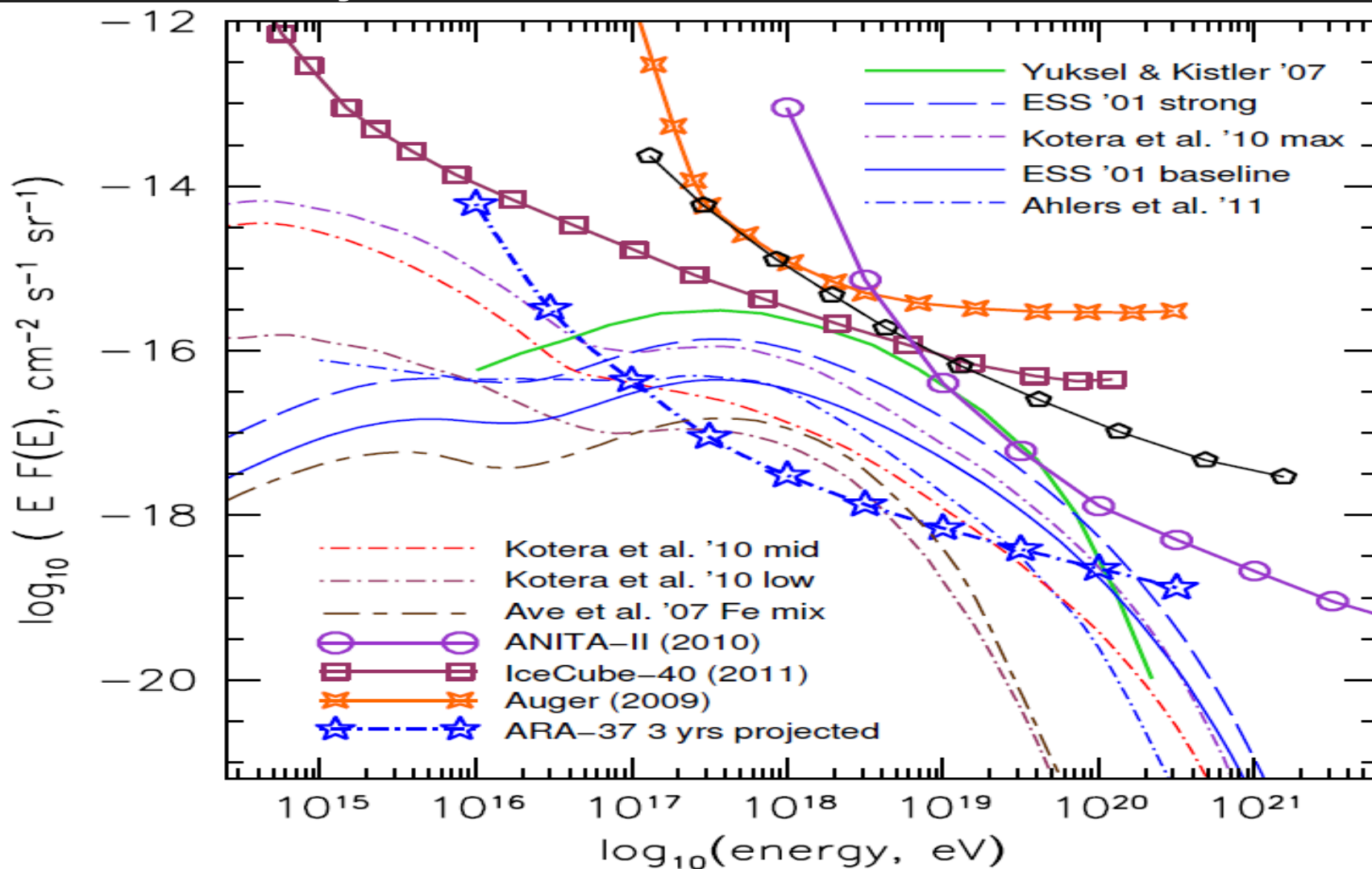
Feb 11,  
2011  
solar  
flare  
obser-  
vation

# Cross-check phase offset of galactic signal...

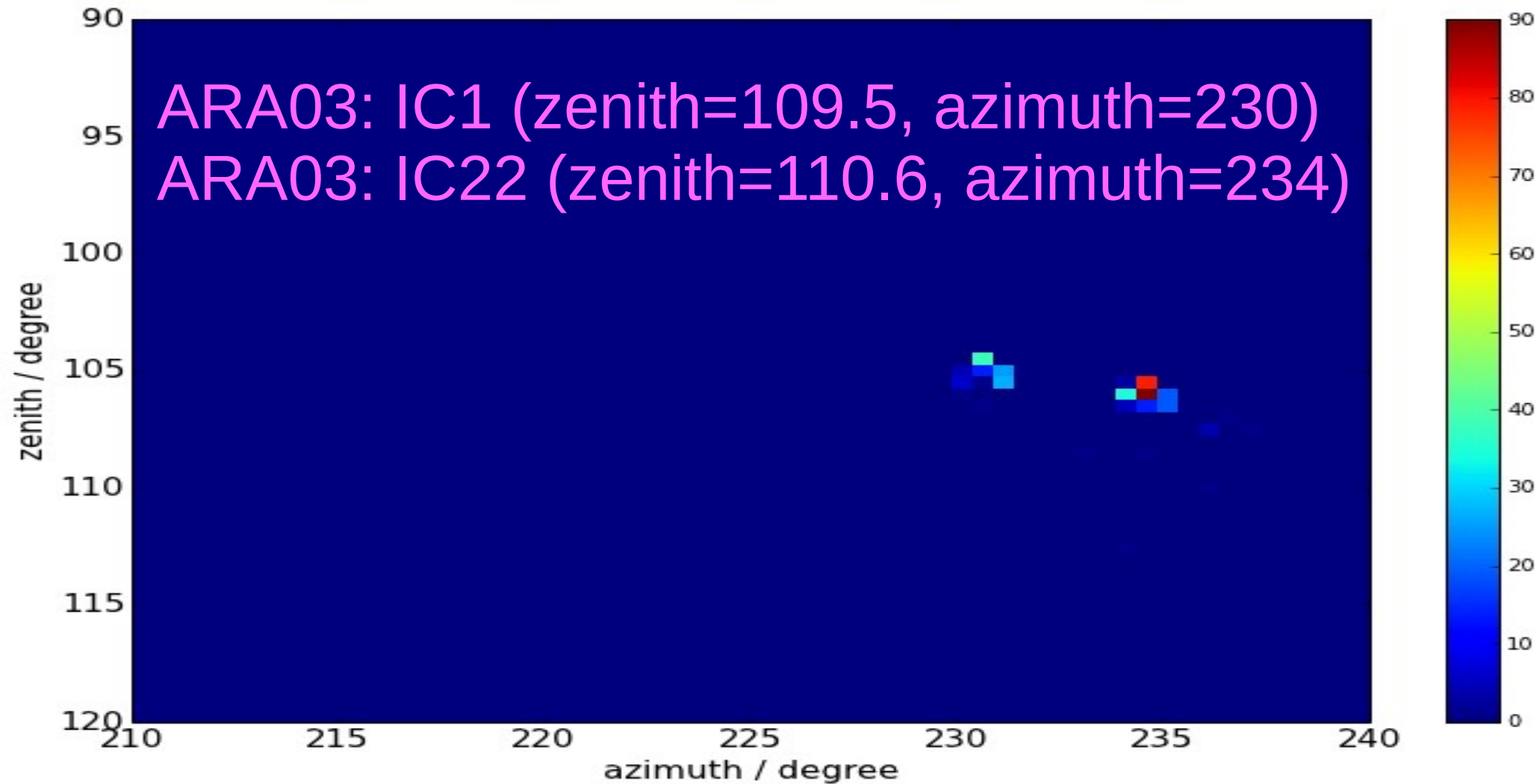


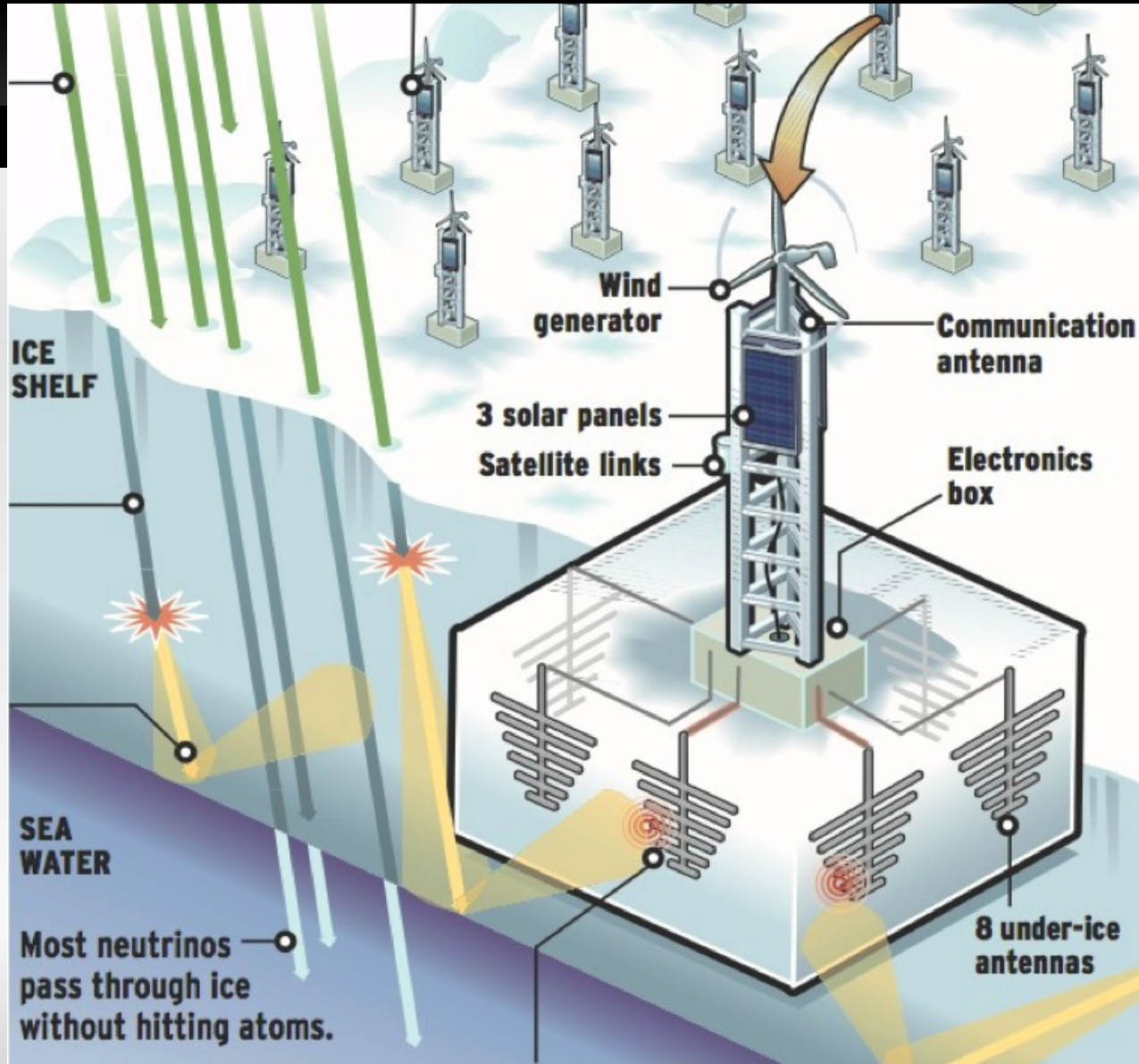
Sidereal  
Variation as  
Surface Rx  
Rotate under  
Galactic center

# Projected Performance – ARA-37



# ARA03 reconstruction





## ARIANNA experiment

Ross ice shelf (Antarctica)

autonomous station

transantarctic mountains  
additional tau neutrino target

measuring  
air-shower flux

560m

direct and reflected neutrino signal

ice

water

