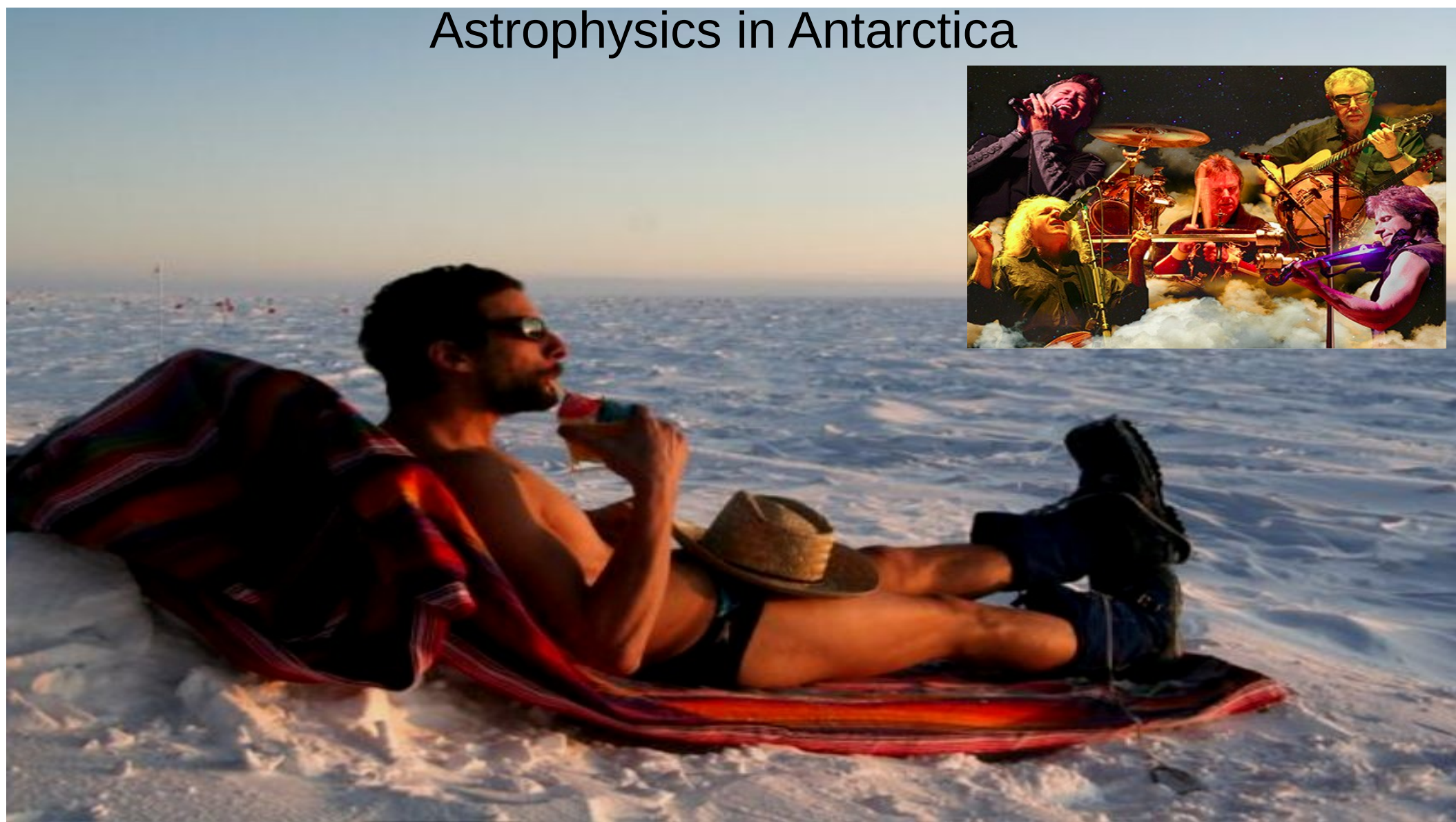


Astrophysics in Antarctica

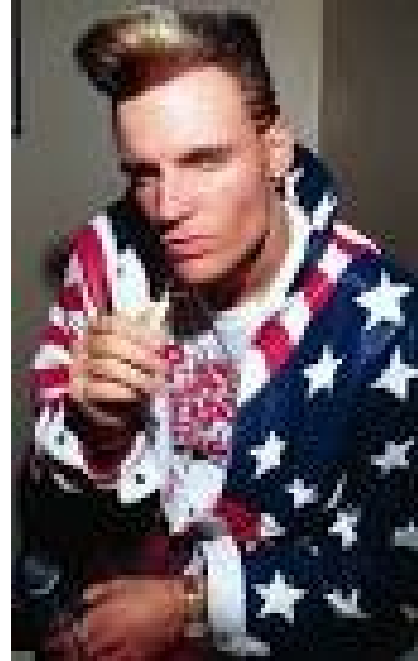


Geophysical Features

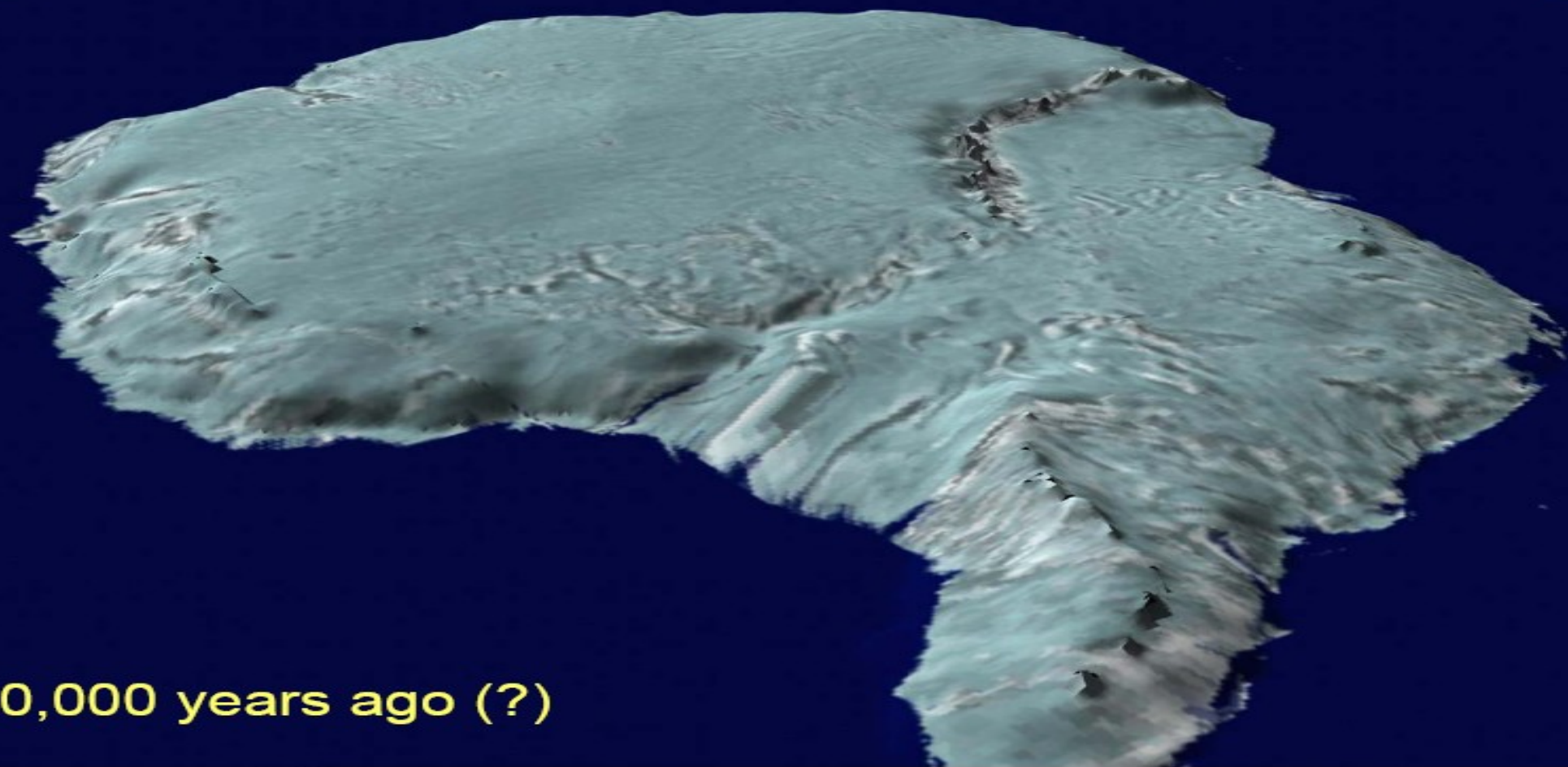


ICE, ICE, BABY

- Complex dielectric permittivity:
 $\epsilon = \epsilon' + i\epsilon''$
- ϵ' : $\text{Re}(\epsilon)$: $c = c_0/n$
- ϵ'' : $\text{Im}(\epsilon)$: gives absorption
- Permittivity is a tensor!
 - Both absorption and wavespeed depend on polarization and propagation-direction

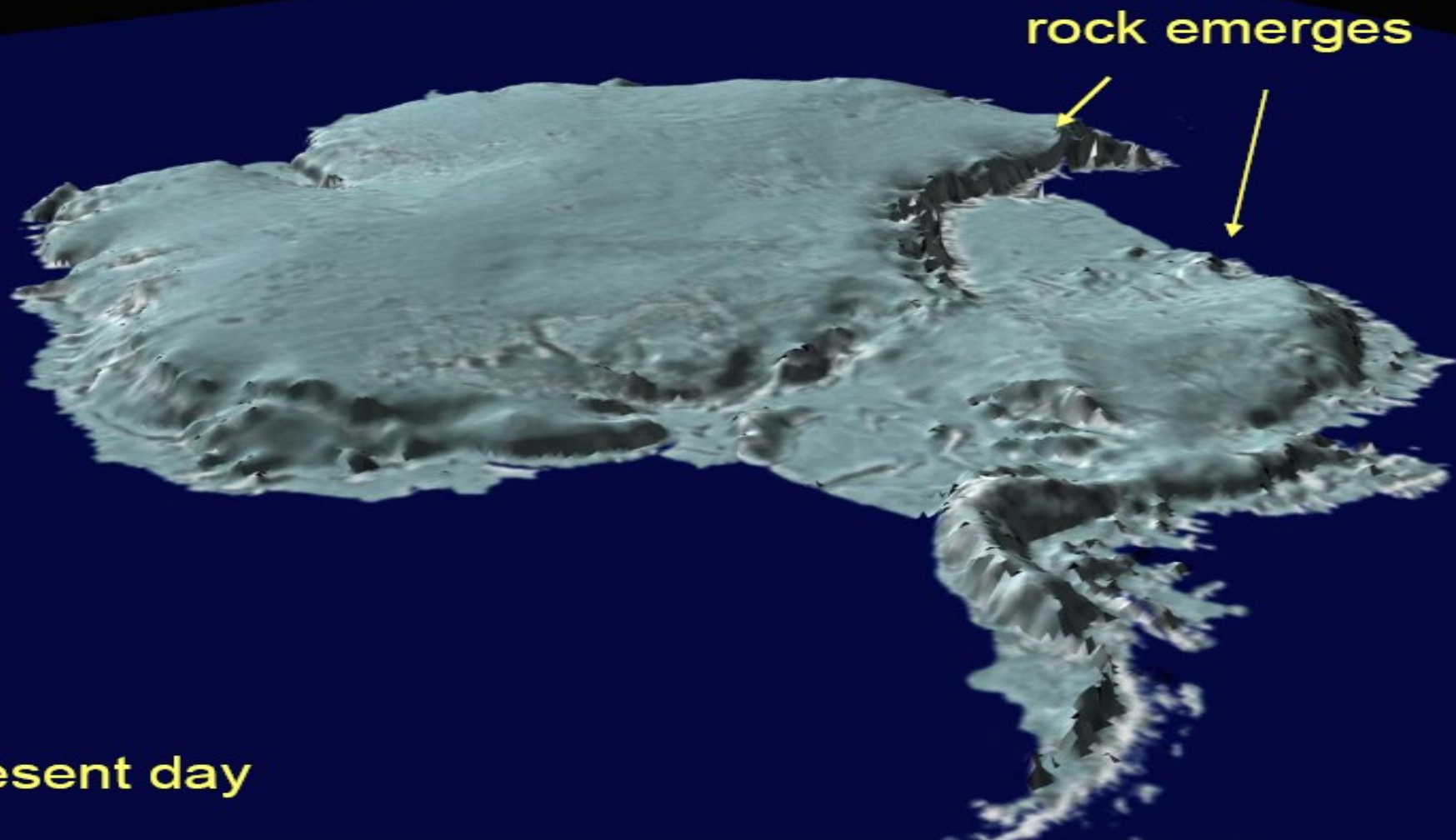


Ice history: Introduction



20,000 years ago (?)

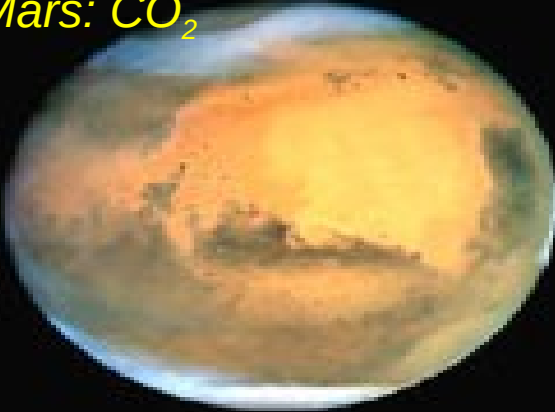
Ice history: Introduction



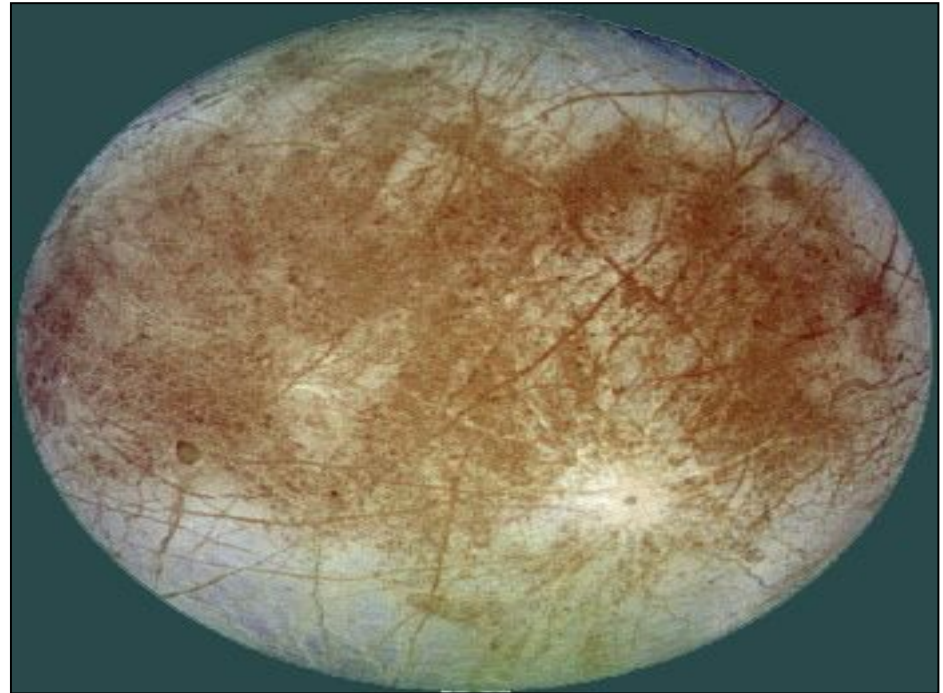
Icecaps elsewhere

*Europa: Water ice - 70K; 10-80 km thick,
Life underneath icecap??*

Mars: CO₂

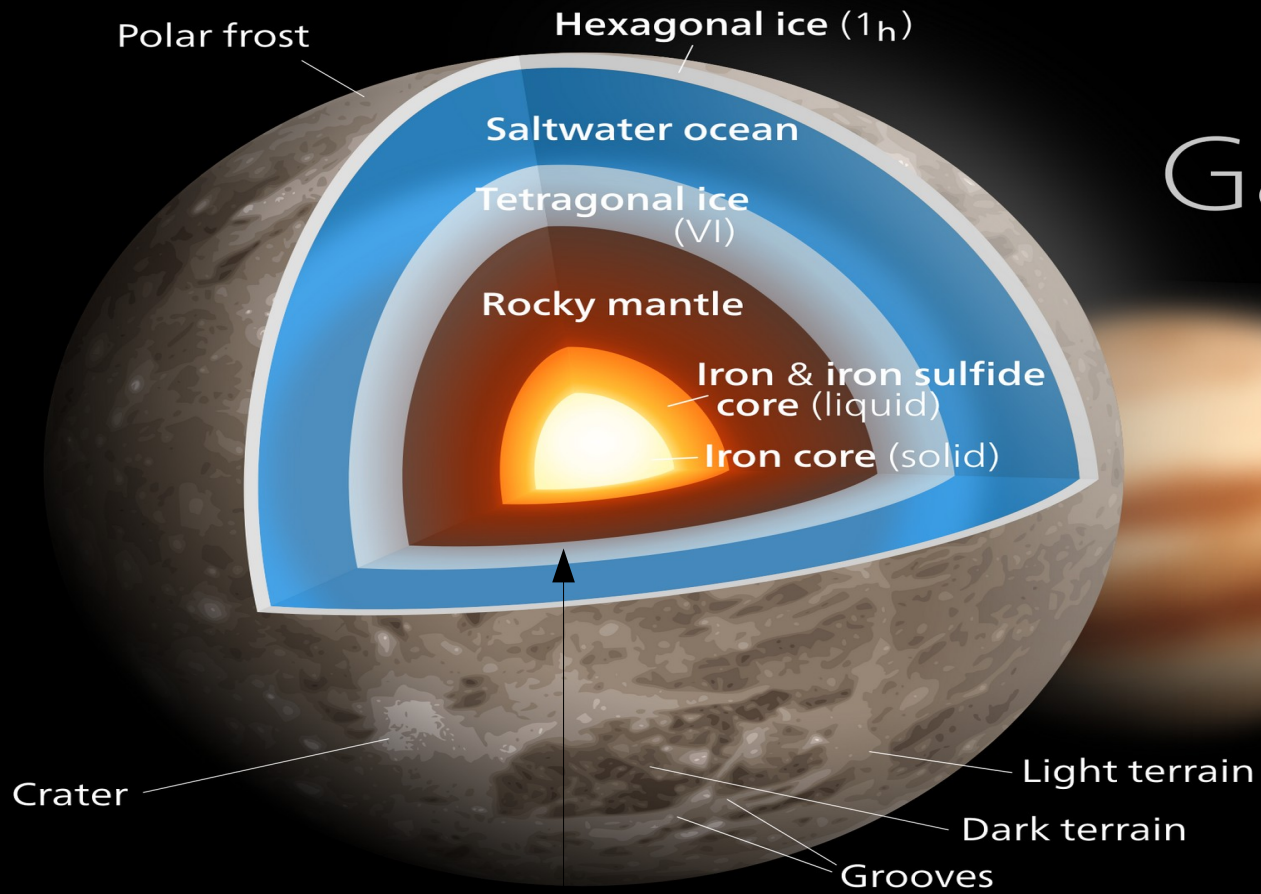


NASA

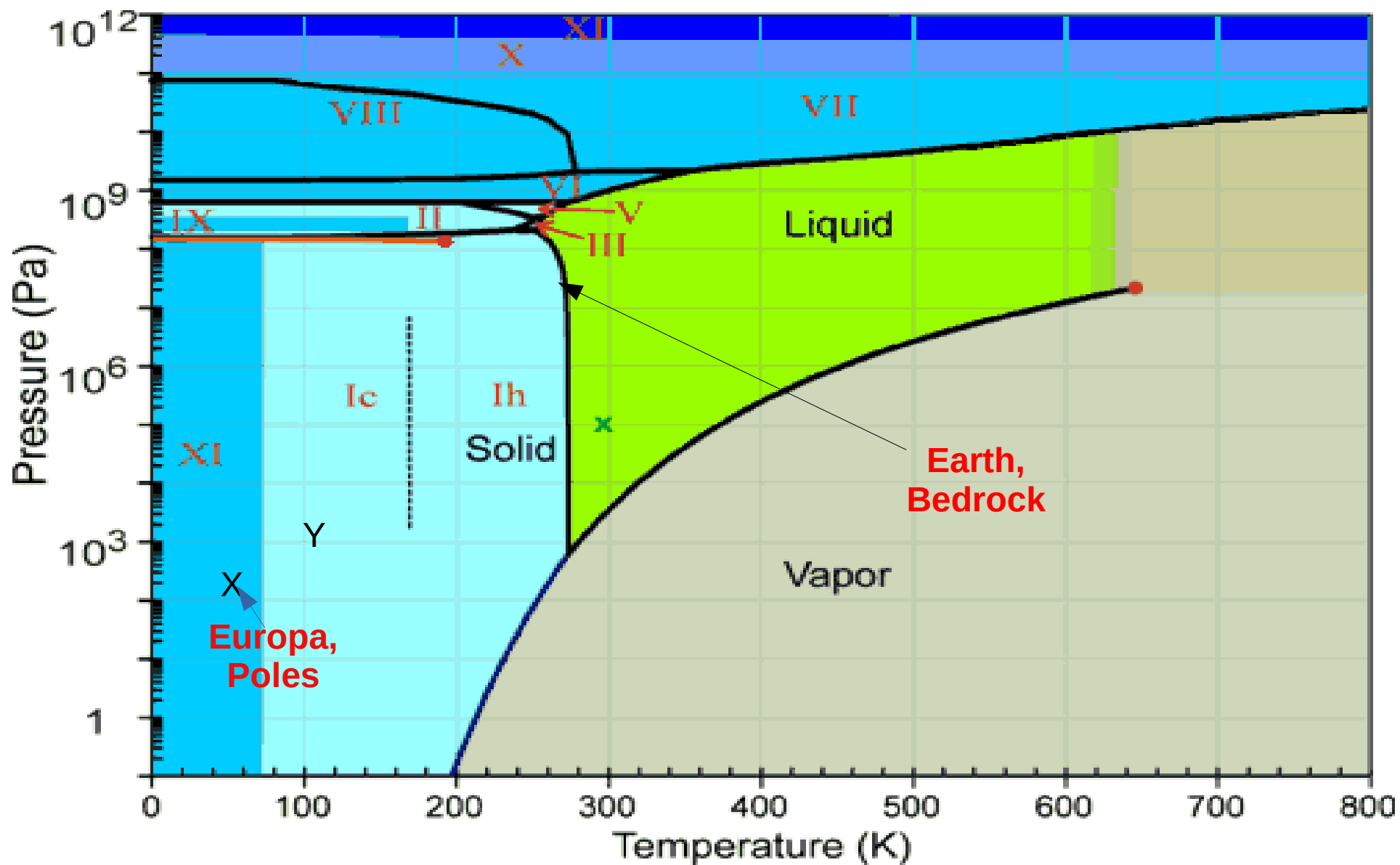


Ganymede

layers drawn to scale



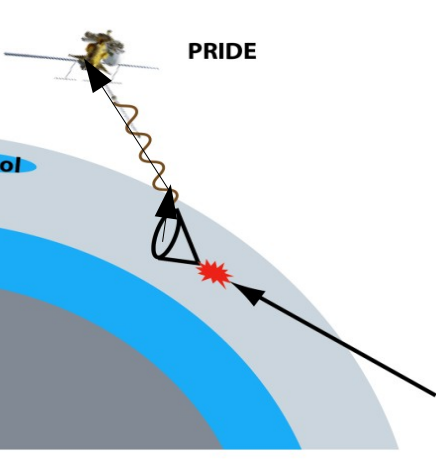
Ice VII melting point ~355 K @ 2 GPa



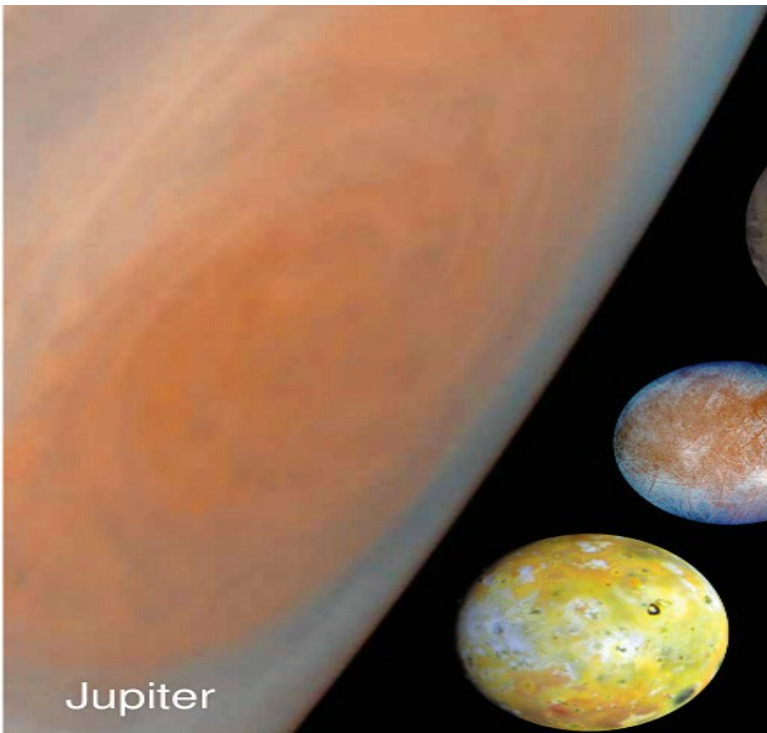
Enceladus



The Instrument Can Be Used to Map Out Large Scale Features in the Ice



PRIDE Experiment: JPL, LPI, KU, OSU, UCI, APL



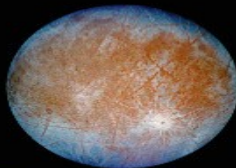
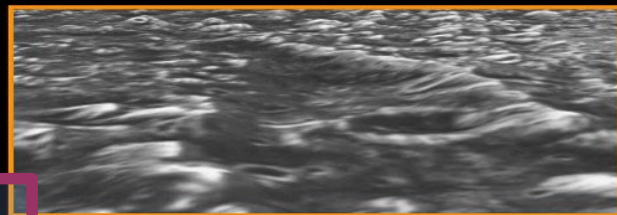
Jupiter



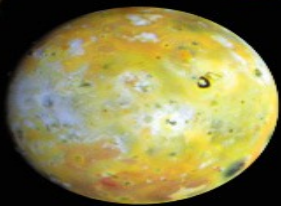
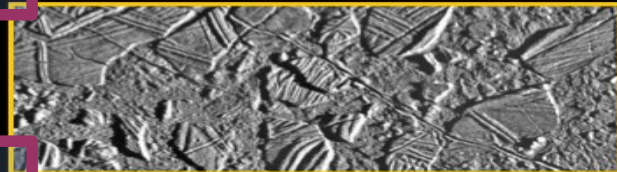
Callisto



Ganymede



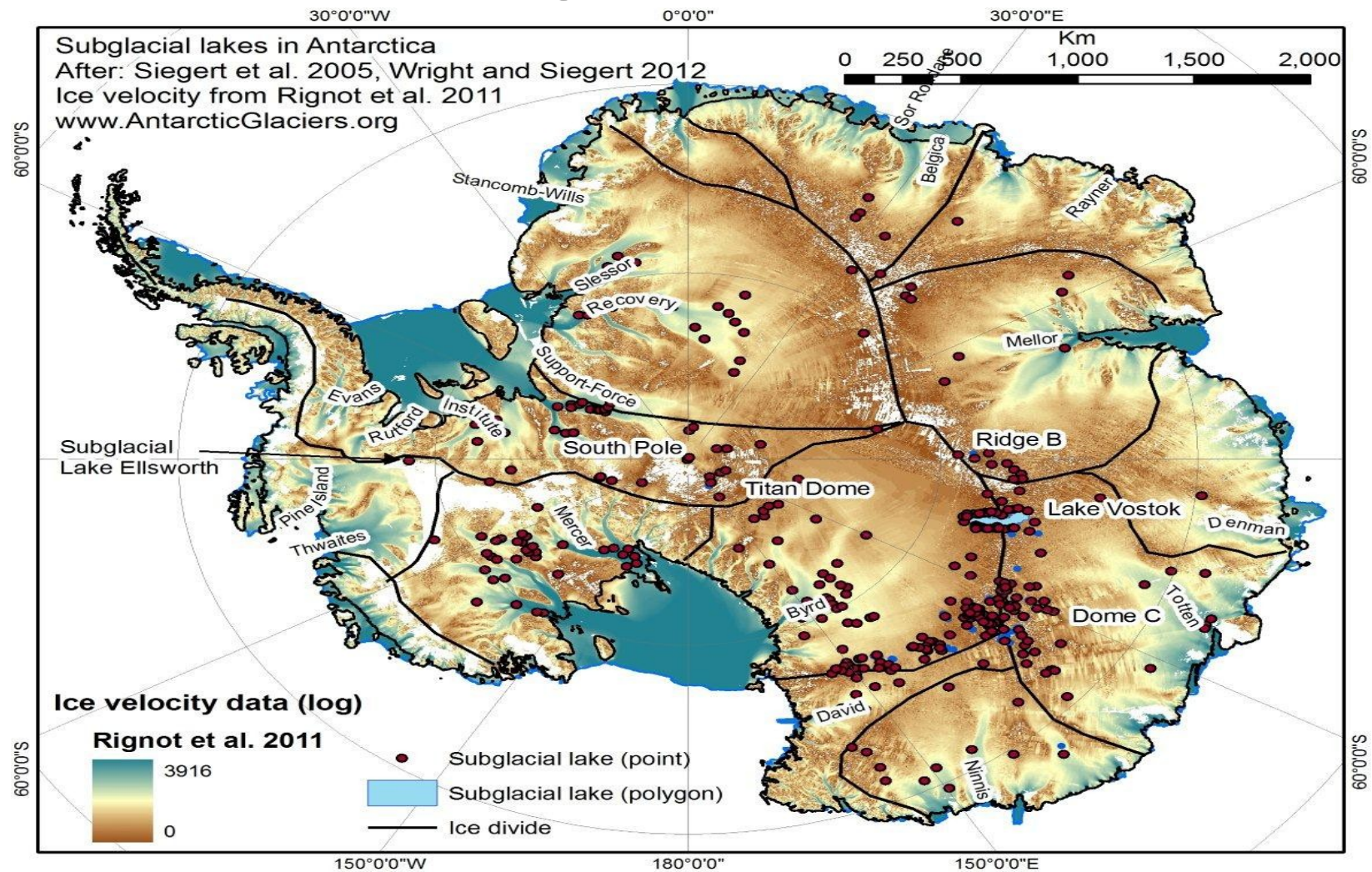
Europa

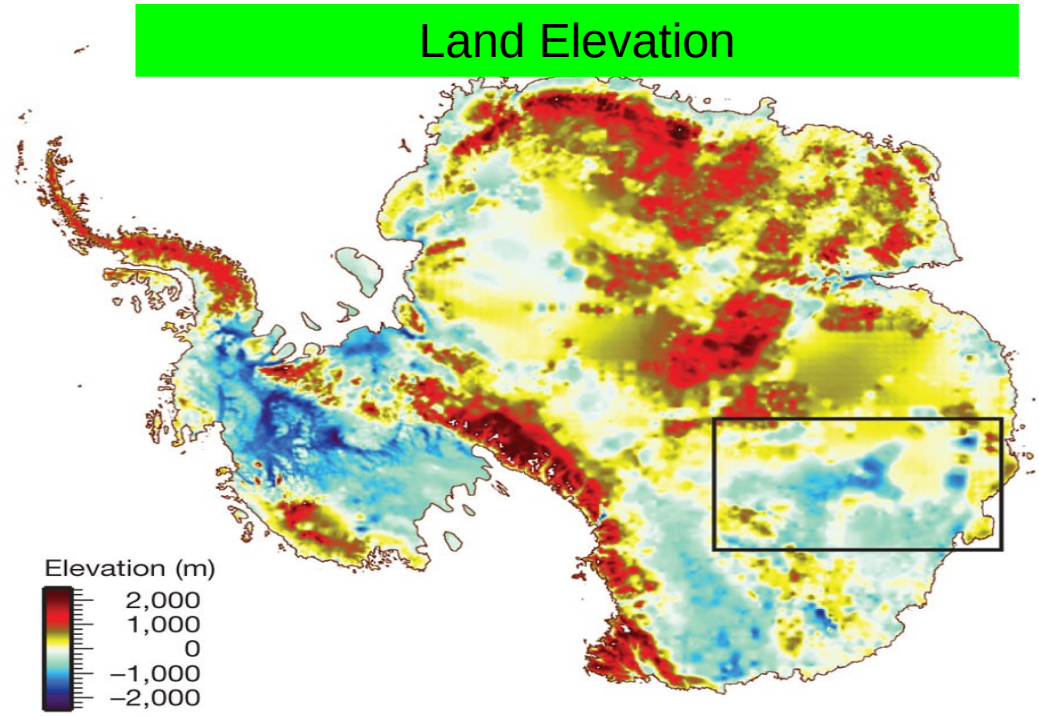
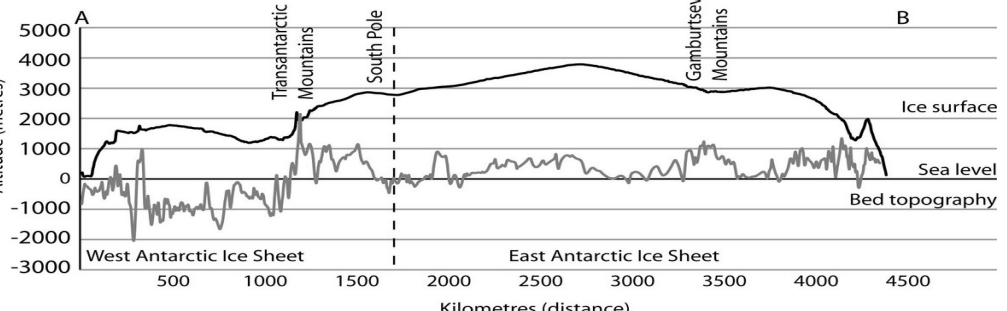
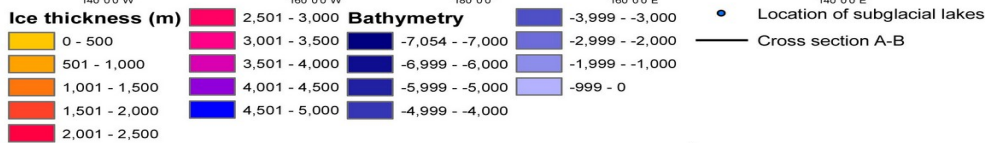
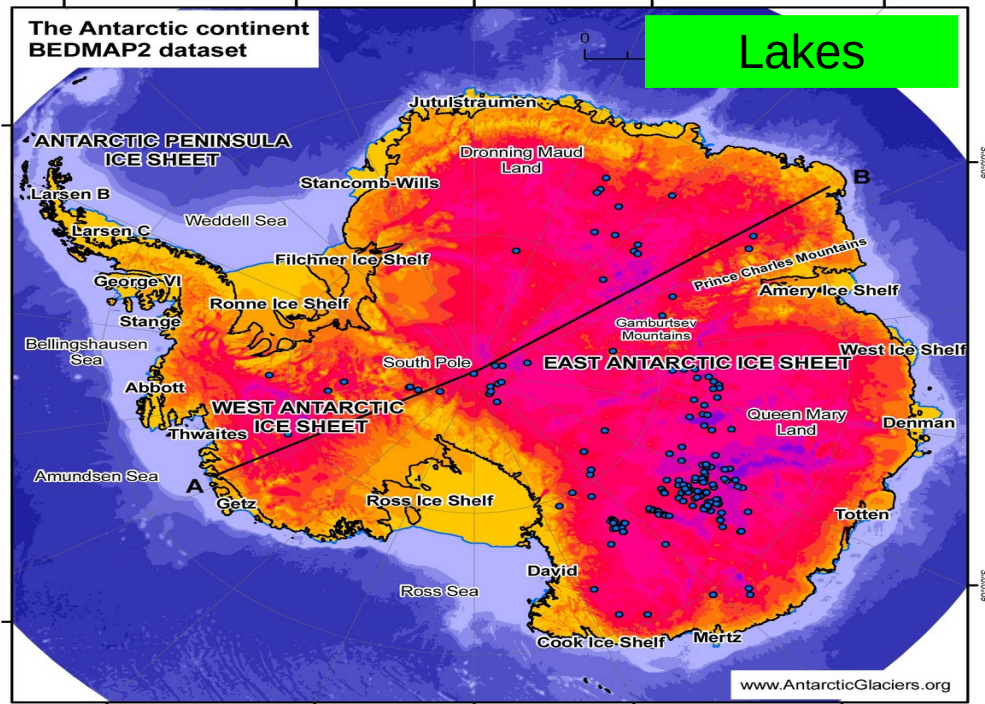


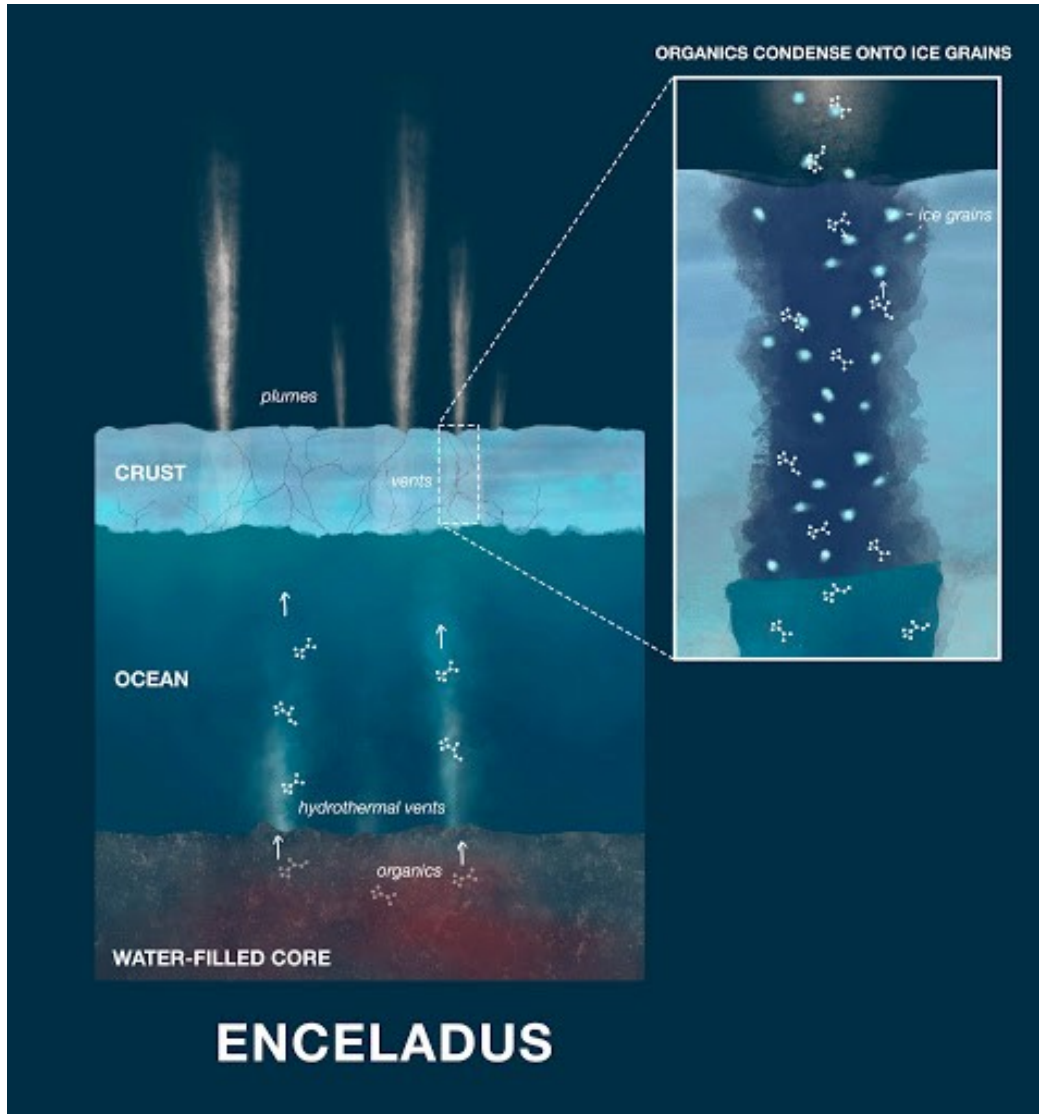
Io



Sub-glacial lakes

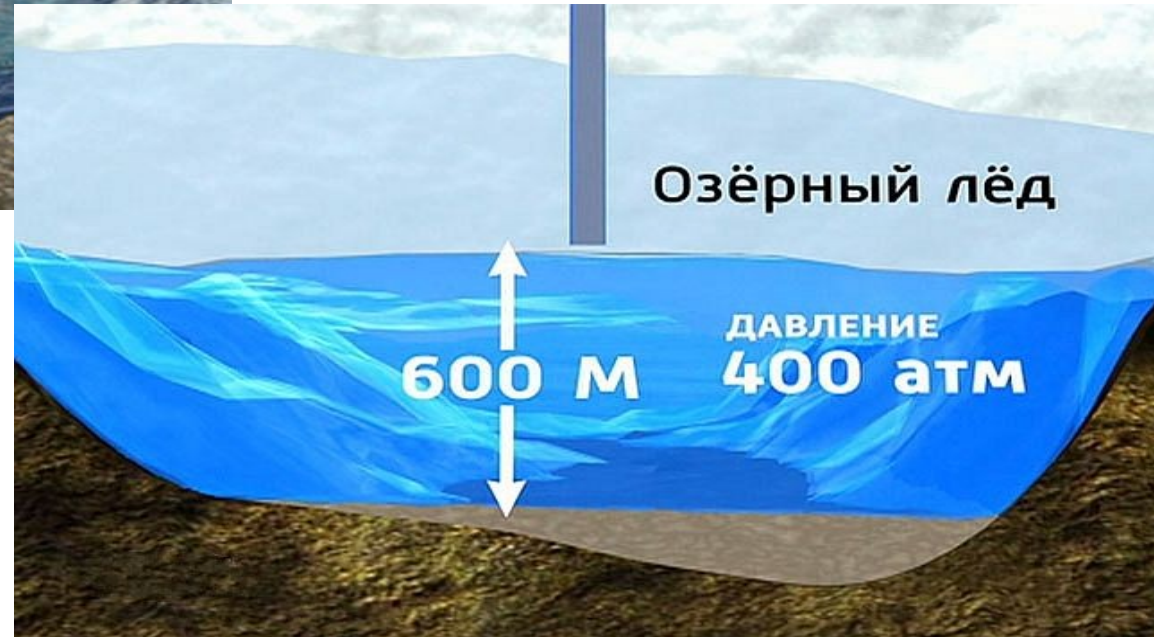
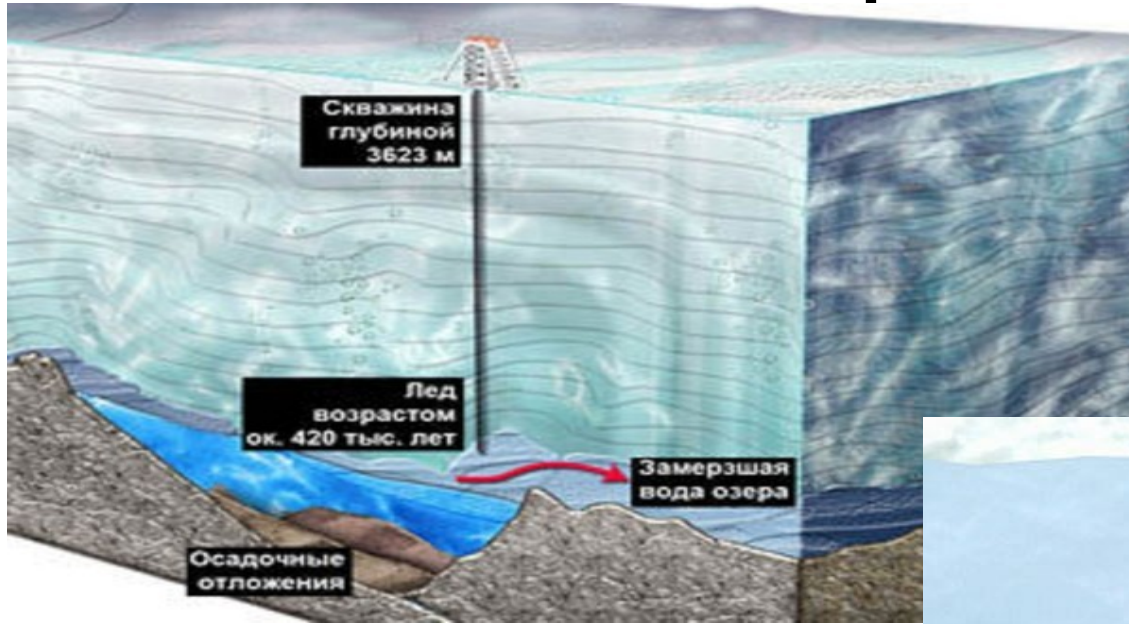




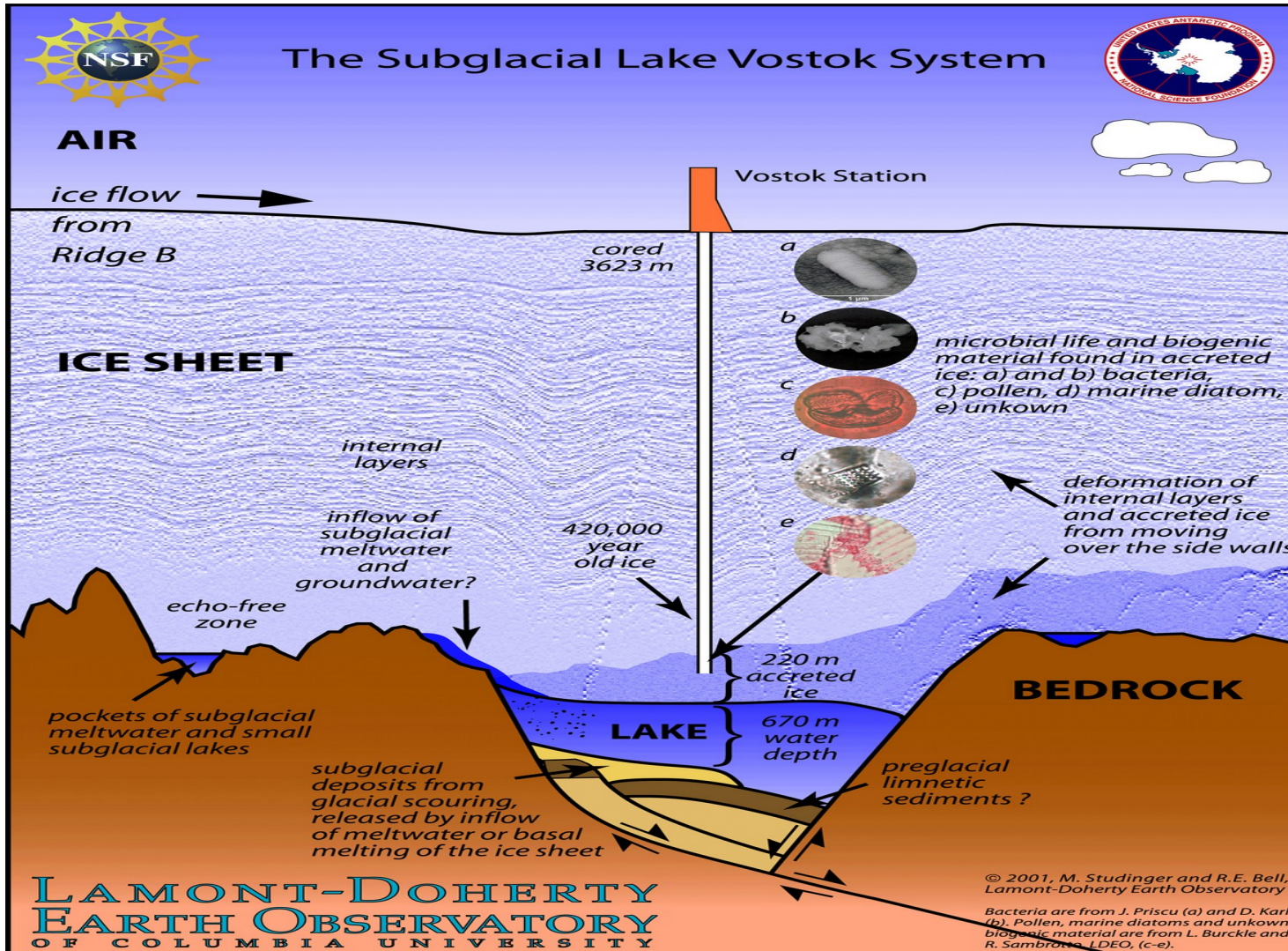


Tidal forces acting on fault lines in the moon's icy shell cause the sides of the faults to rub back and forth against each other, producing enough heat to transform some of the ice into plumes of water vapor and ice crystals.

Озеро Восток



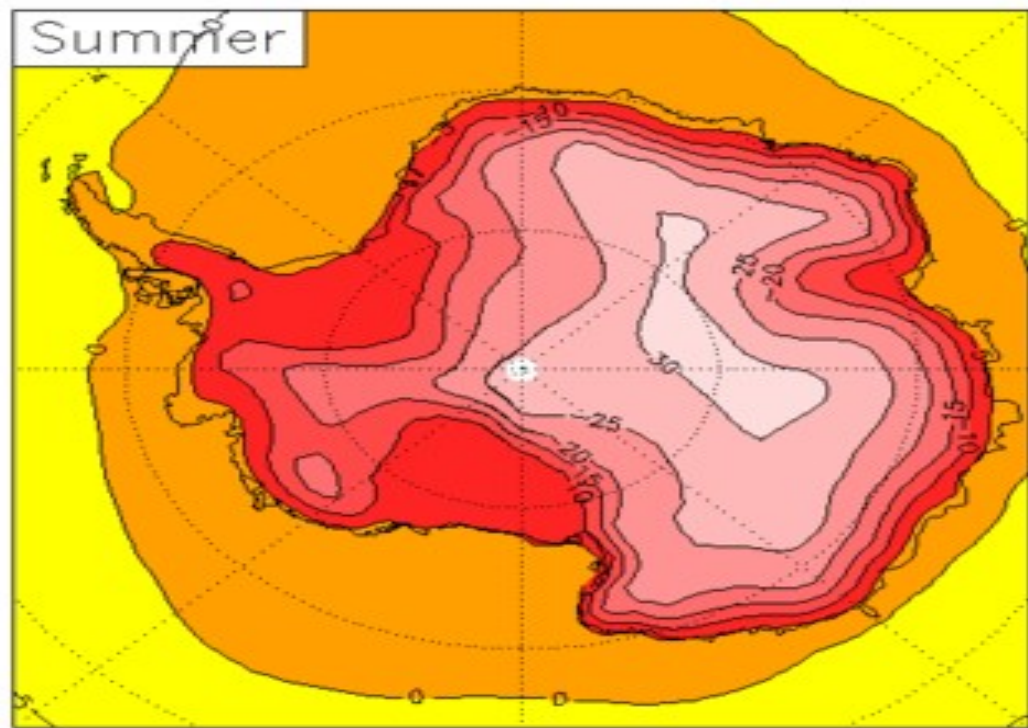
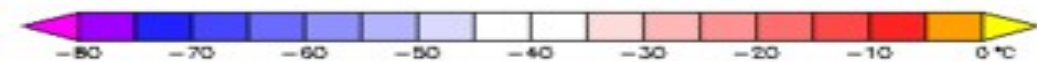
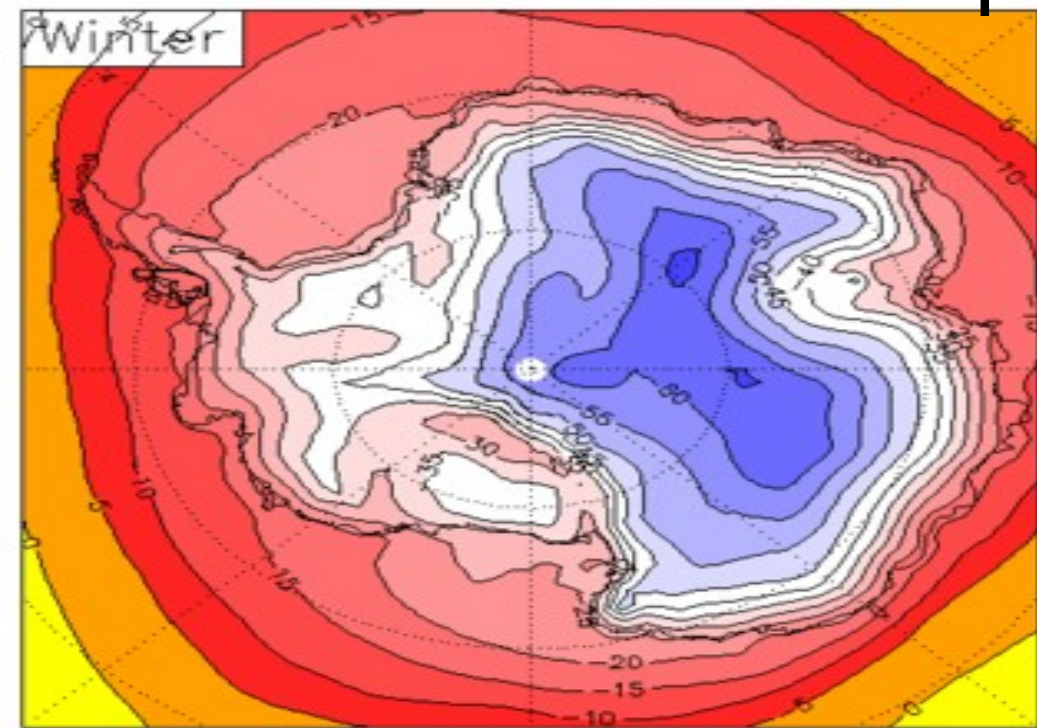
Life within the ice sheet



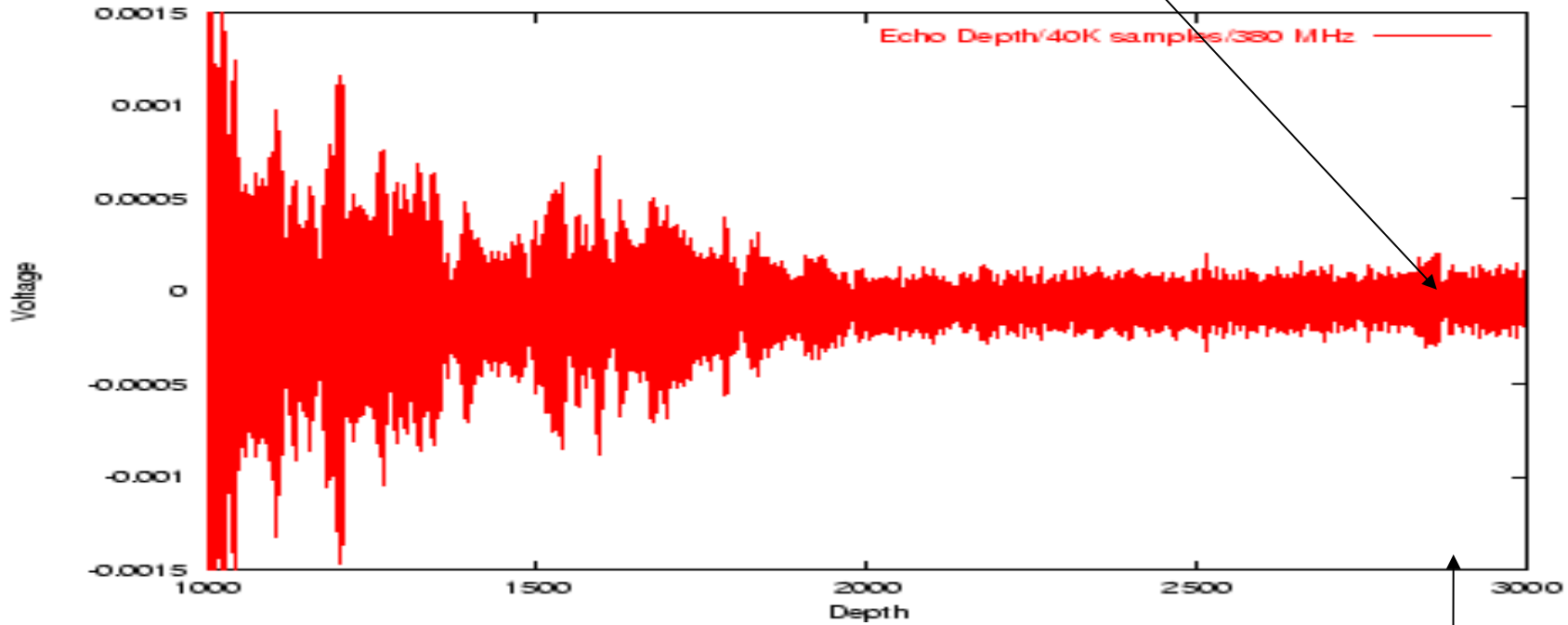
Aside: why doesn't lake freeze?

Why don't Europa/Enceladus lakes freeze?

Surface Temperature



Bottom echo visible thru 5.6 km! Attenuation Length ~ 1.6 km. cf:
Effective optical $L_{\text{atten}} \sim 40$ m (J. Glaciol. 51, 231 (2005).)

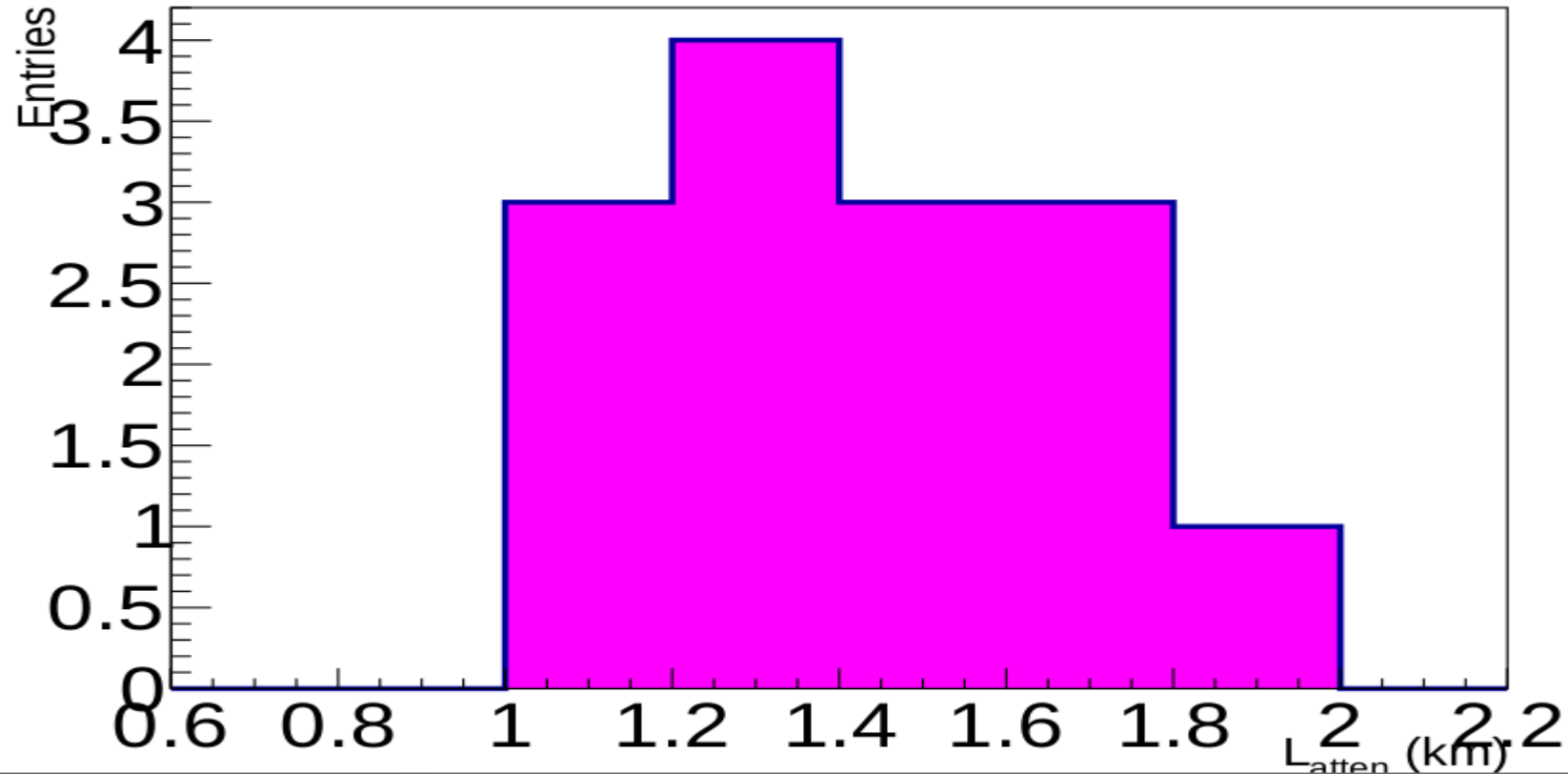


Bedrock/2850m

TAGLINE: Cold, clean ice is very RF-transparent (Europa, e.g.)

South Pole ϵ'' Data

RF Field Attenuation Length (A4/A5 DP data): $\langle L_{\text{atten}} \rangle = 1.43 \pm 0.25$ km



N.B. Shadow Zone effective attenuation length ~ 500 m

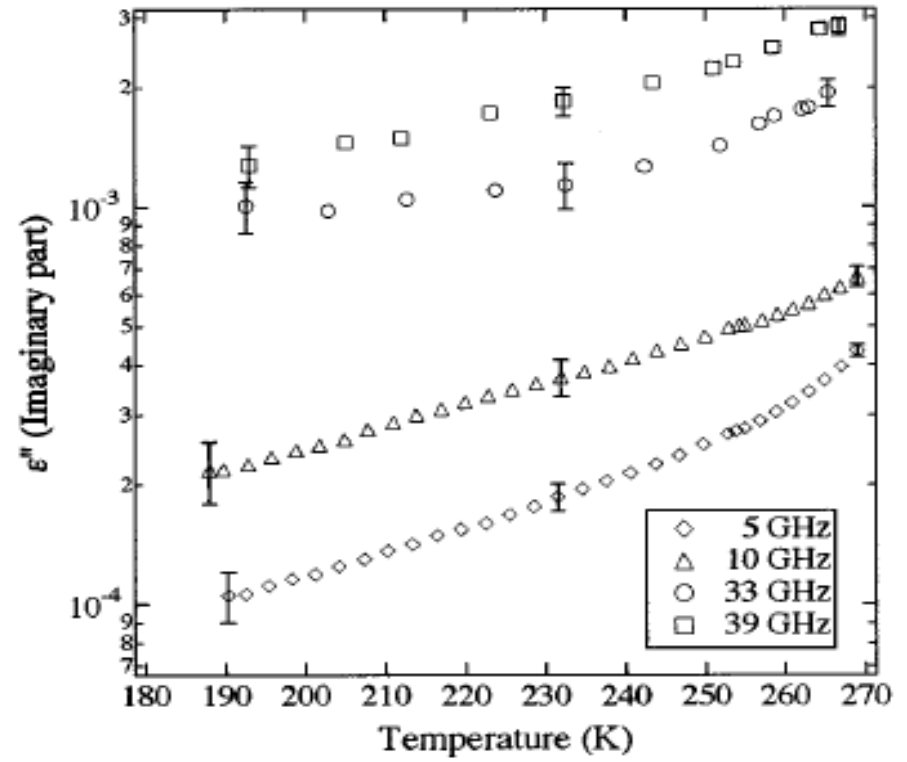
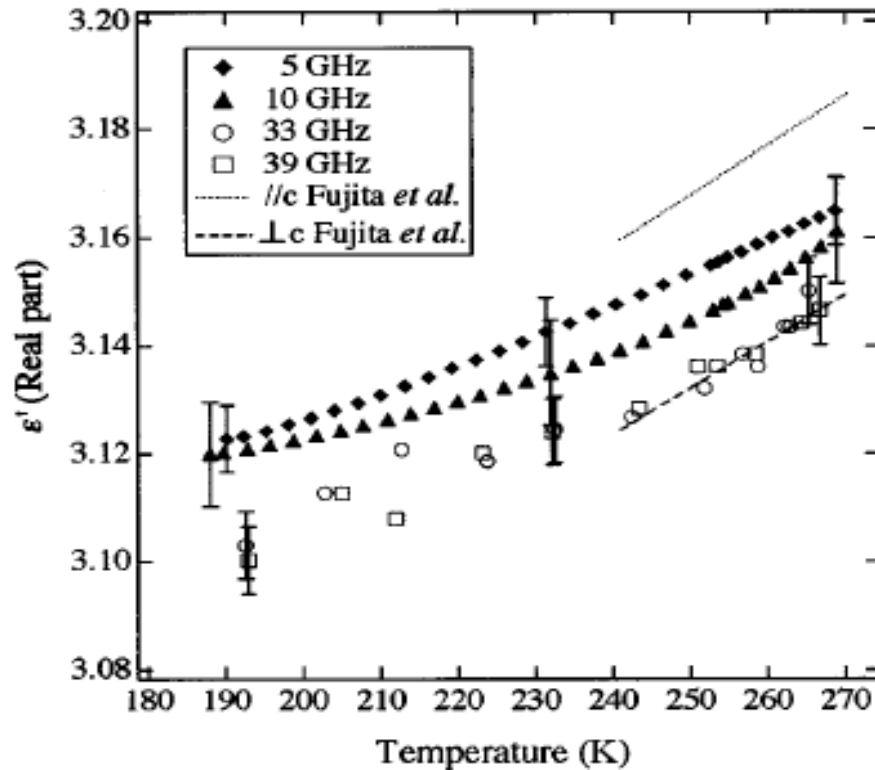
Cold Polar Ice Complex Permittivity

- $\boldsymbol{\varepsilon} = \boldsymbol{\varepsilon}' + i\boldsymbol{\varepsilon}''$
 - $\boldsymbol{\varepsilon}' = \text{Real}(\boldsymbol{\varepsilon})$: refractive-index (non-absorptive)
 - $\boldsymbol{\varepsilon}'' = \text{Im}(\boldsymbol{\varepsilon})$: attenuation length (absorption)
 - Tensor characteristics of $\boldsymbol{\varepsilon} \Rightarrow$ refractive index and attenuation length depend on polarization
 - \Rightarrow 'birefringence'

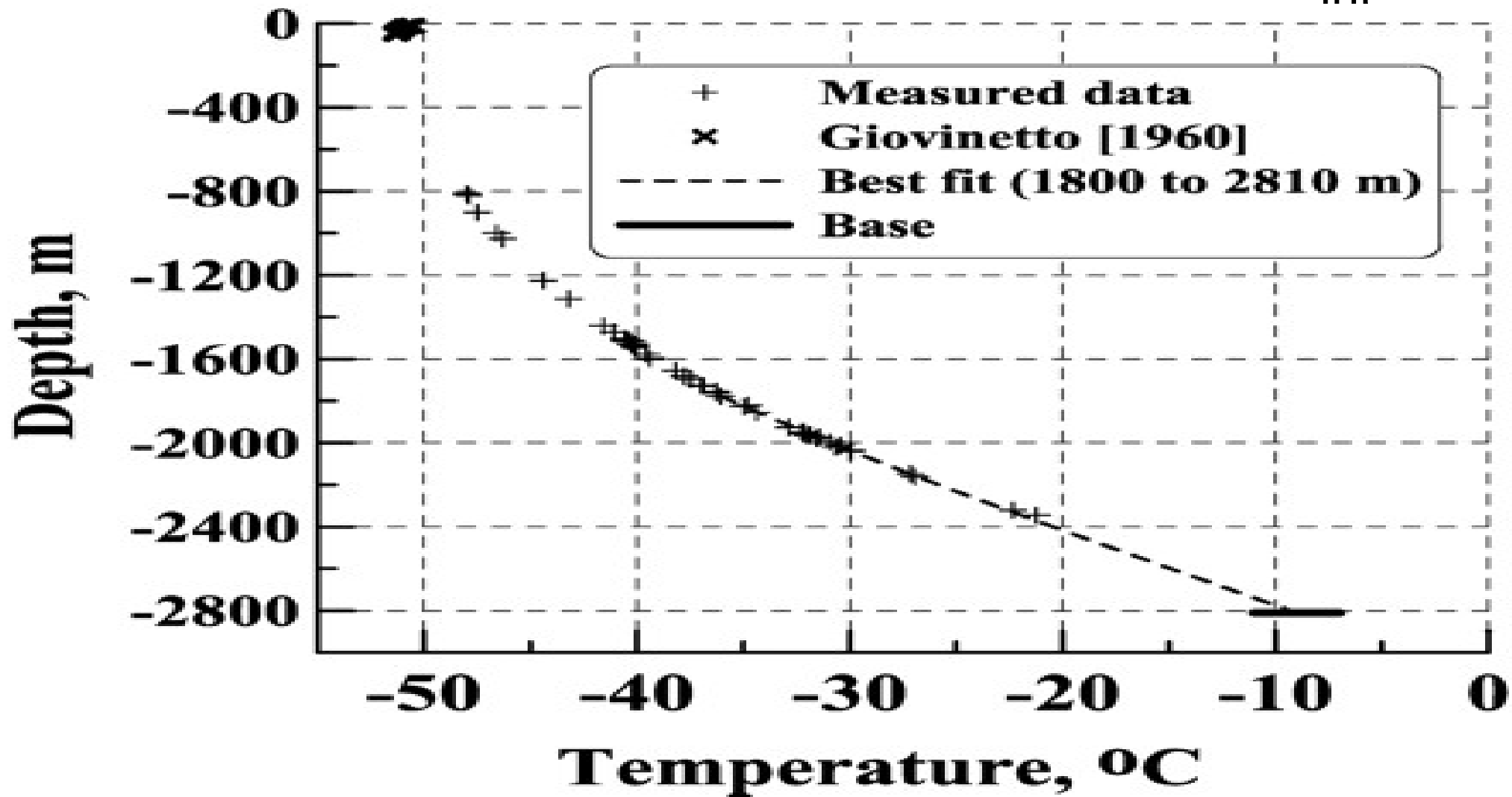
Permittivity of ice dependent on temperature.

Matsuoka, T., S. Fujita and S.J. Mae. 1996. Effect of temperature on dielectric properties of ice in the range 5–39 GHz. *J. Appl. Phys.*, **80**(10), 5884–5890.

$1/\epsilon =$ attenuation length

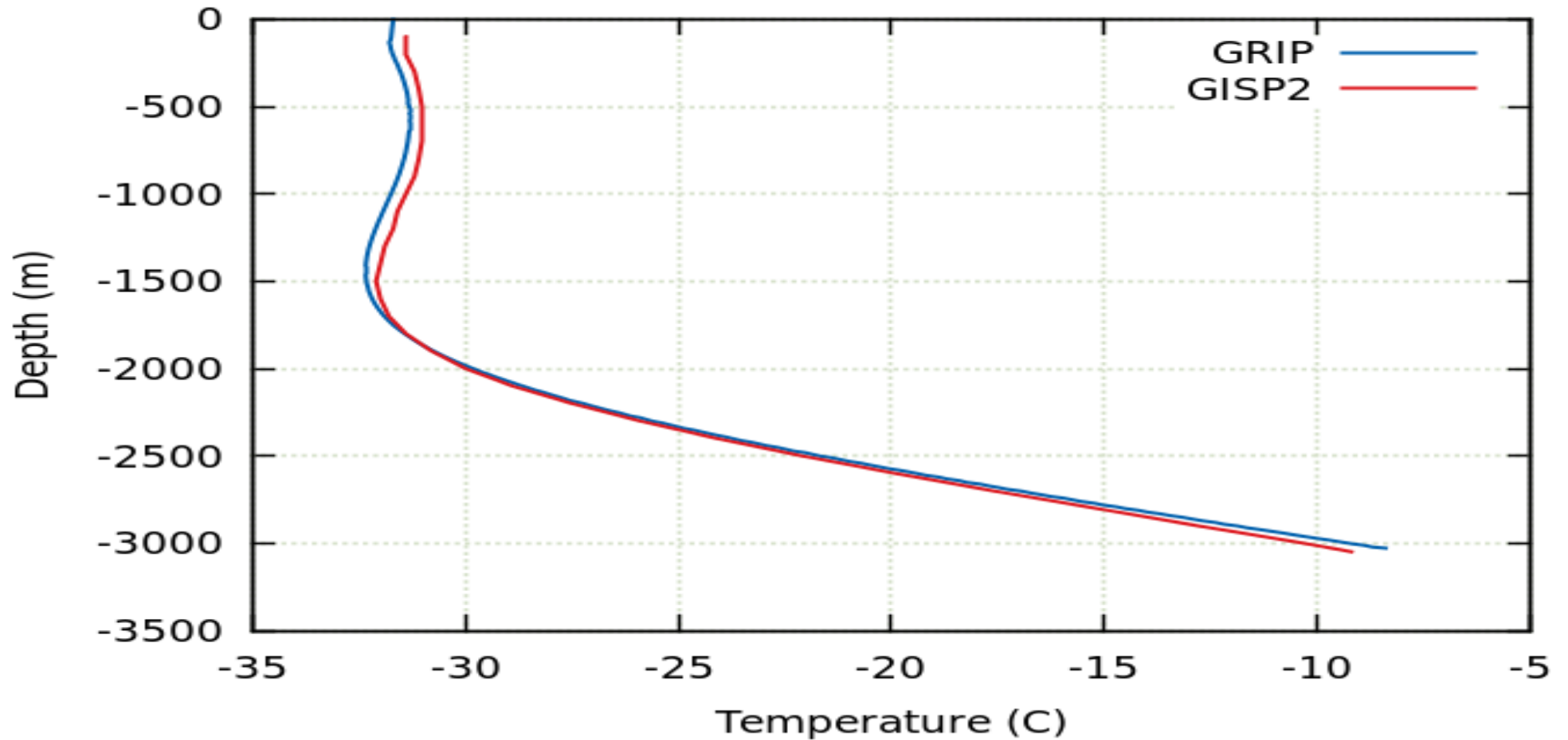


SP Temperature Profile (affects $L_{\text{int}}(z)$)

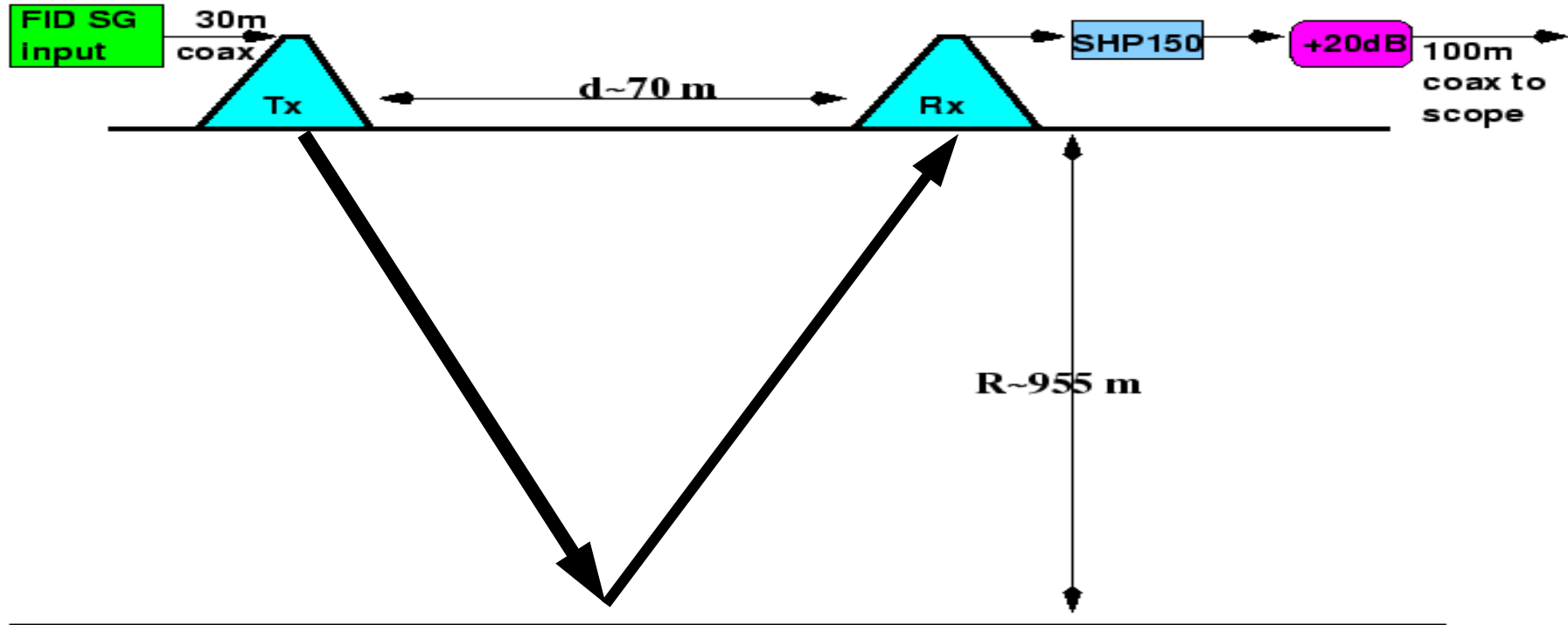


Greenland T(z)

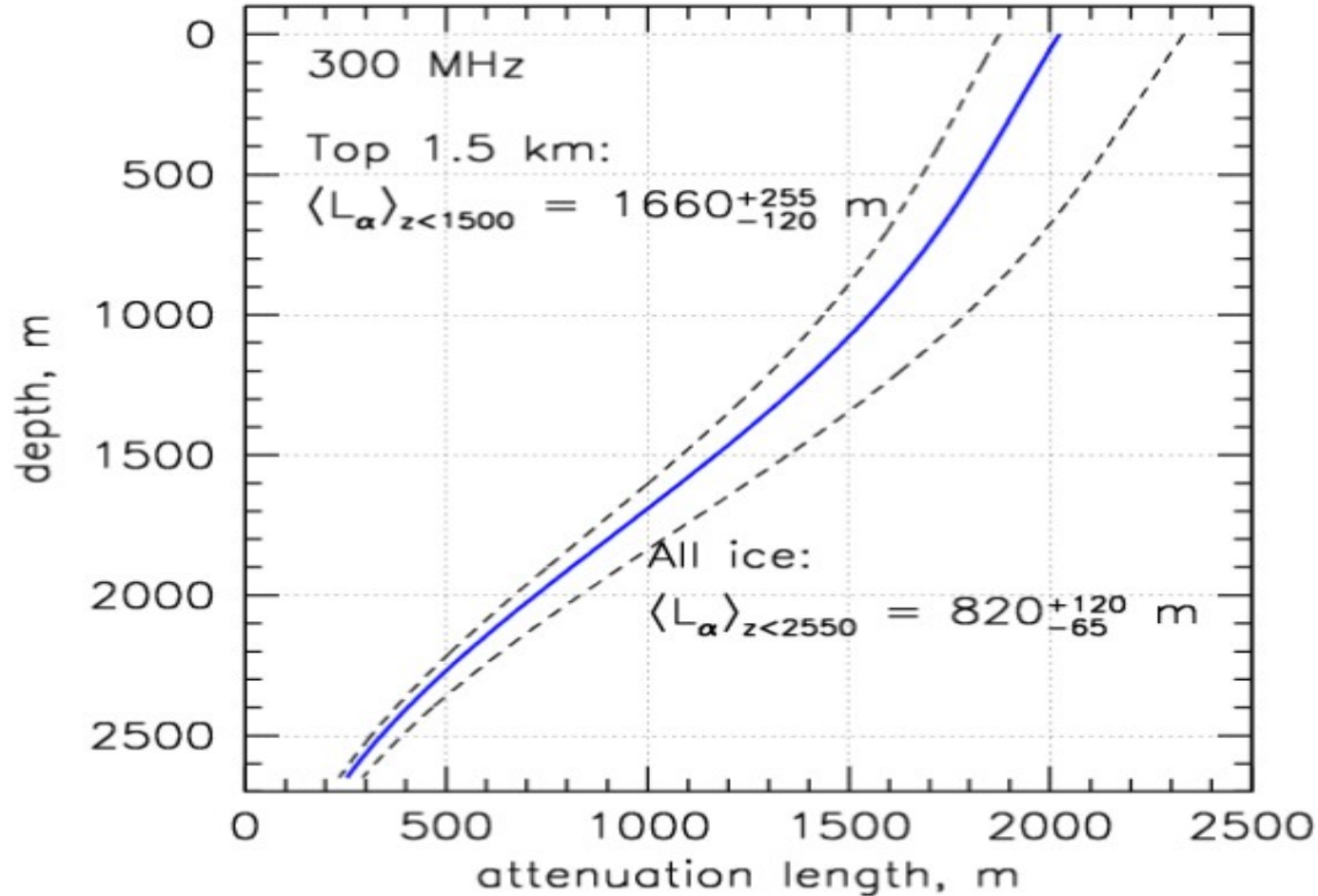
Greenland Temperature Profile at Summit



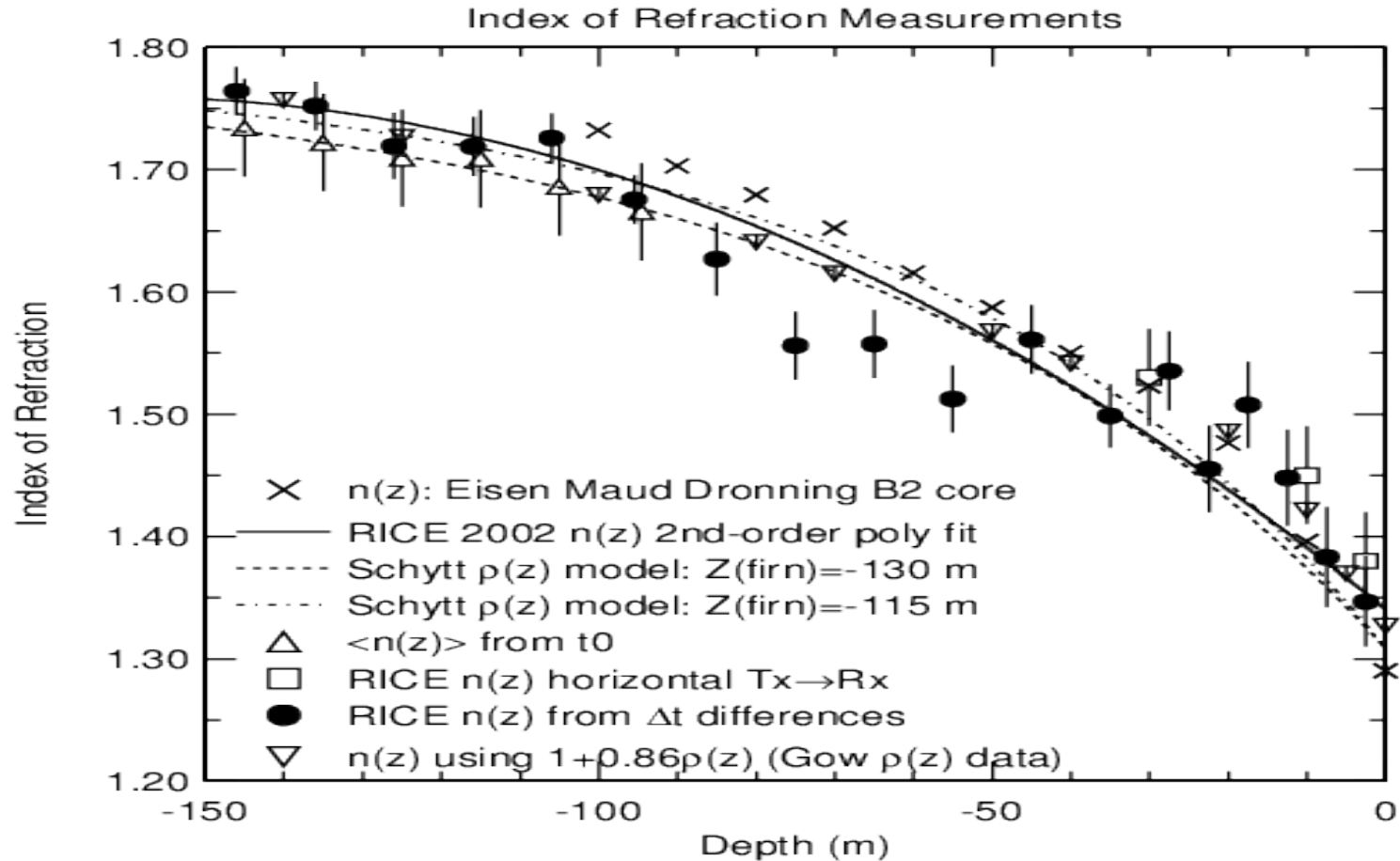
Birefringence tests: South Pole ($R=2850$ m) and Taylor Dome ($R=955$ m)



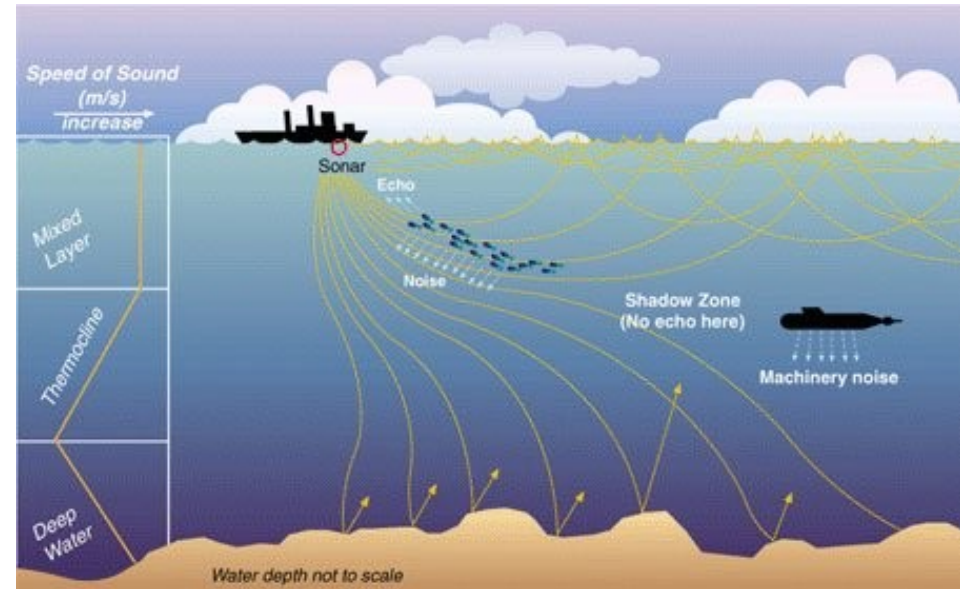
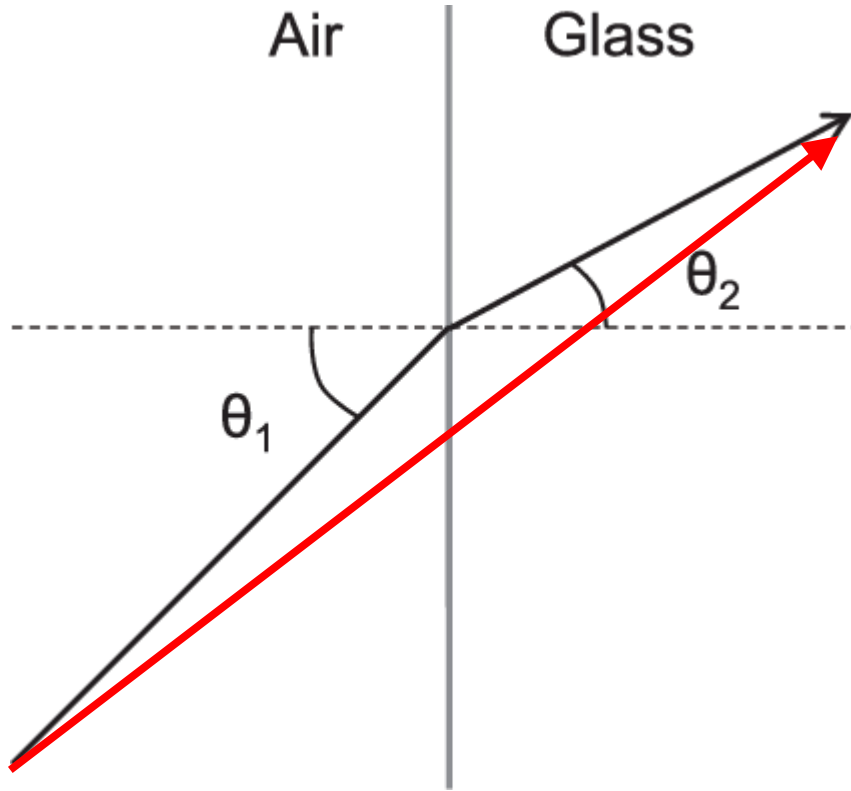
Attenuation length as f(depth)



South Pole ϵ' Data



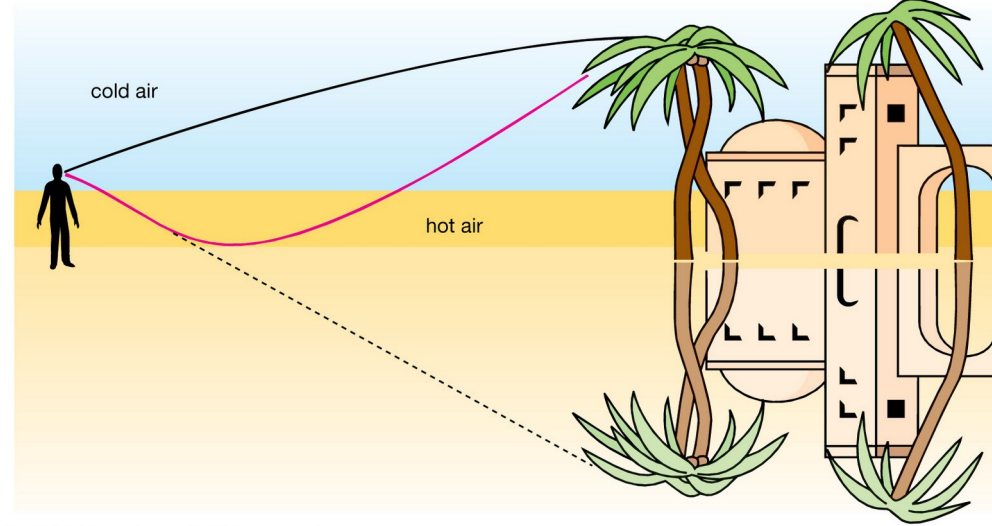
Fermat's Least time principle



Fermat's least time principle and mirages



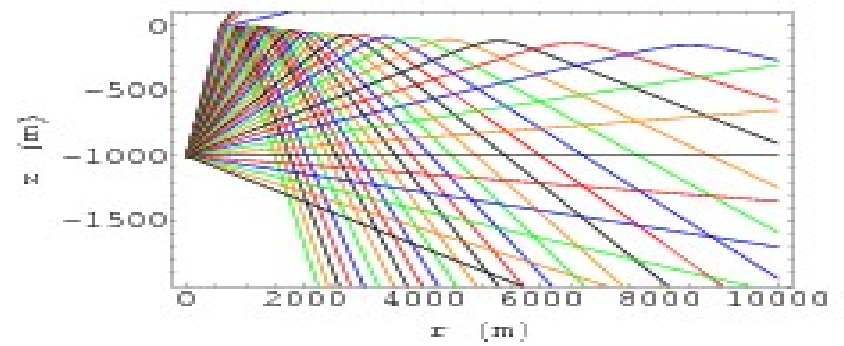
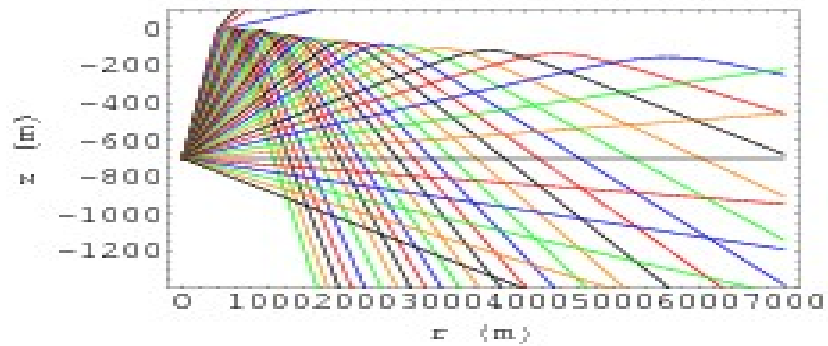
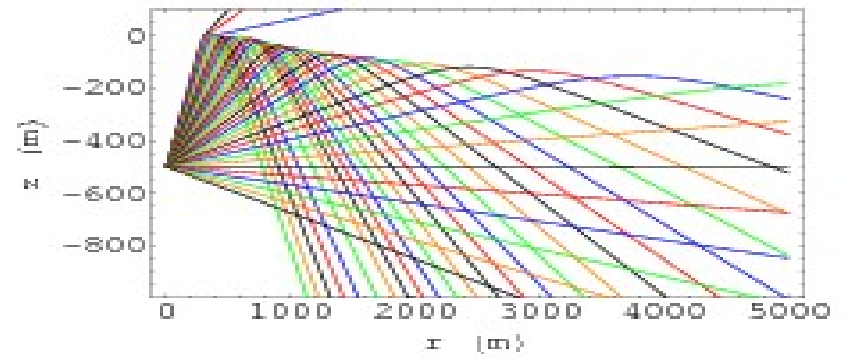
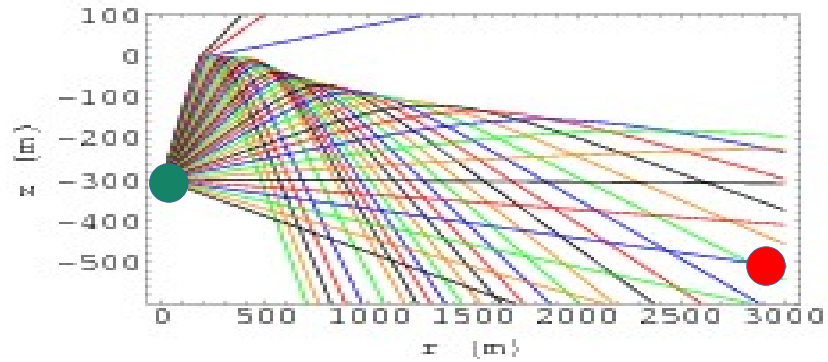
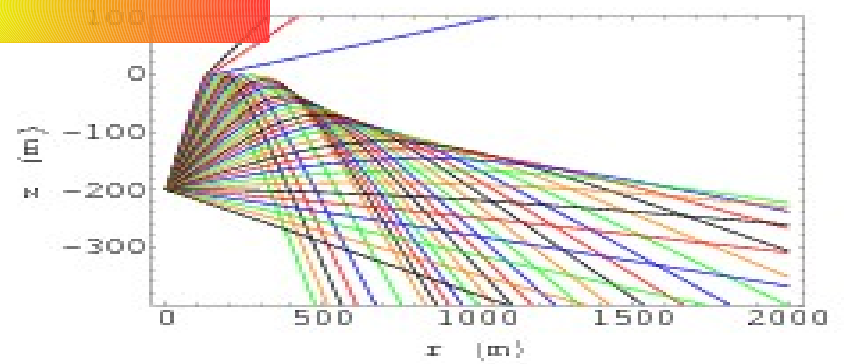
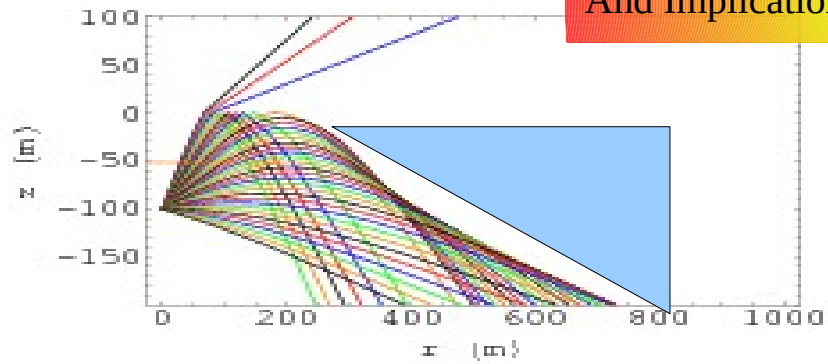
Formation of an inferior image

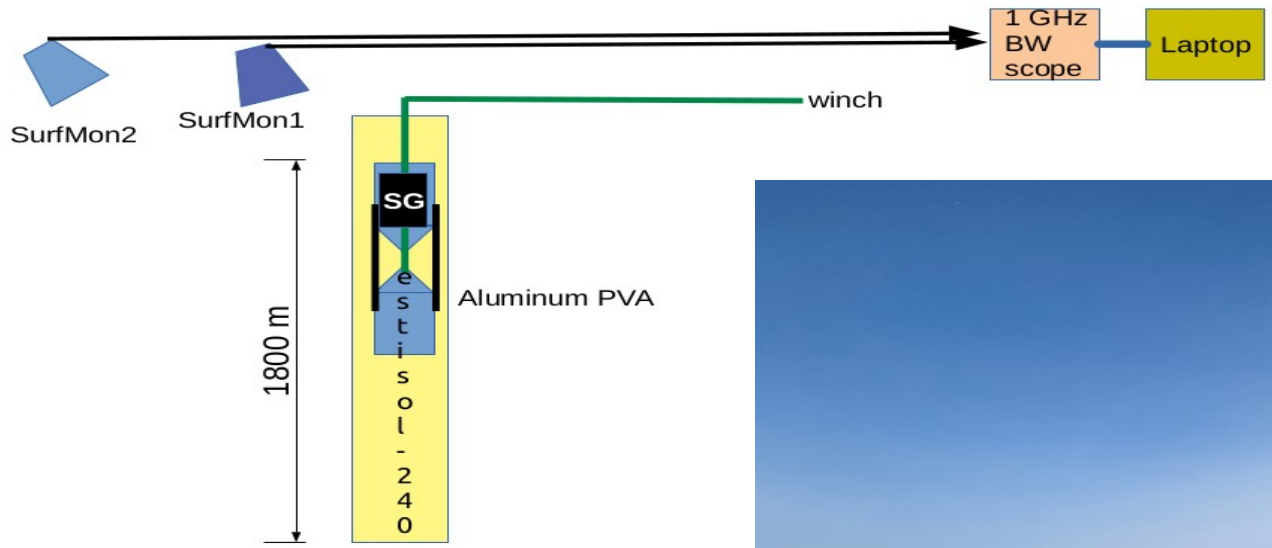


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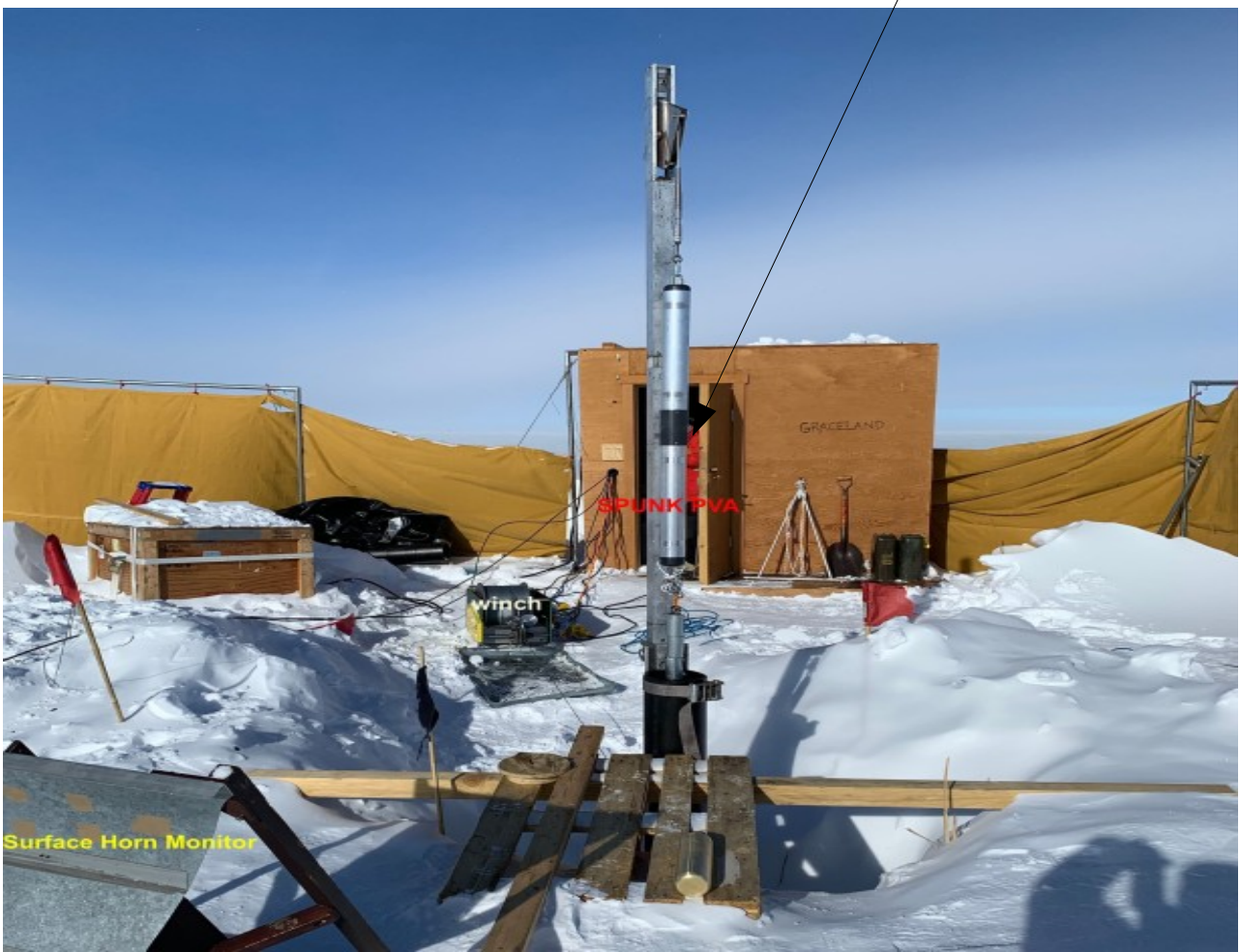
What is the effect of an index-of-refraction that changes with z ?

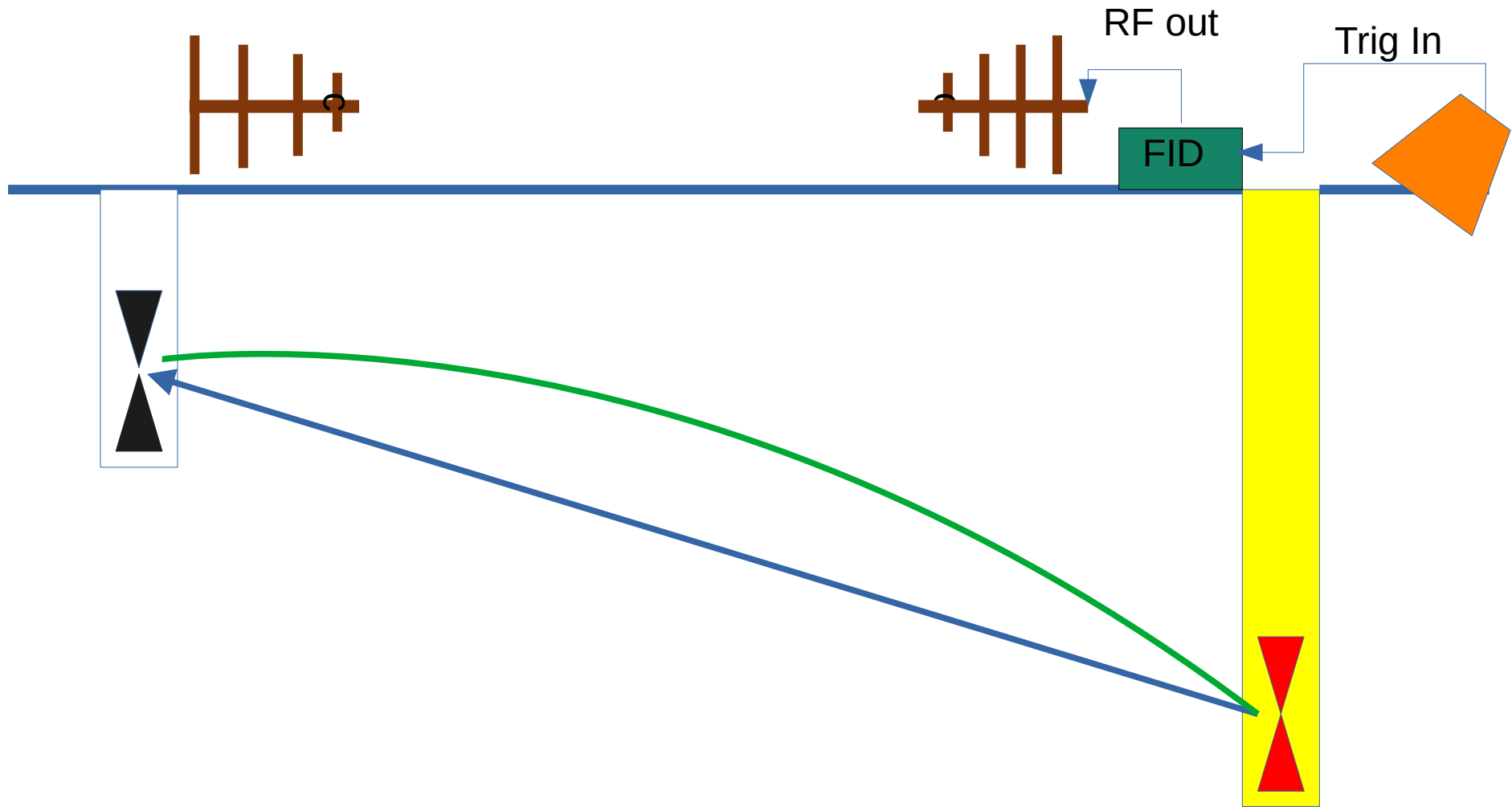
And Implications....



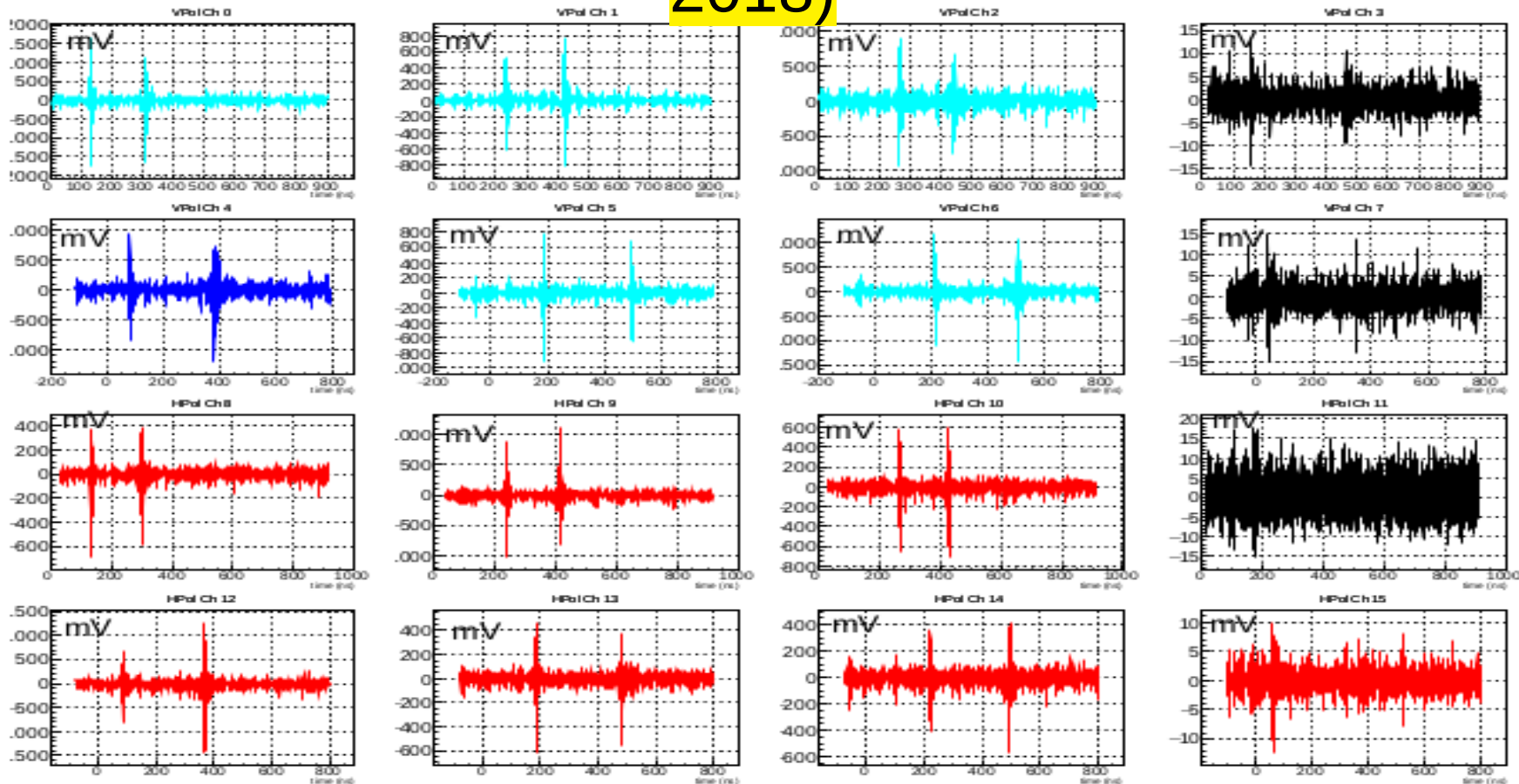


SPICE core testing (Dec., 2018)





Double pulses and refraction – SPICE core data (Dec., 2018)

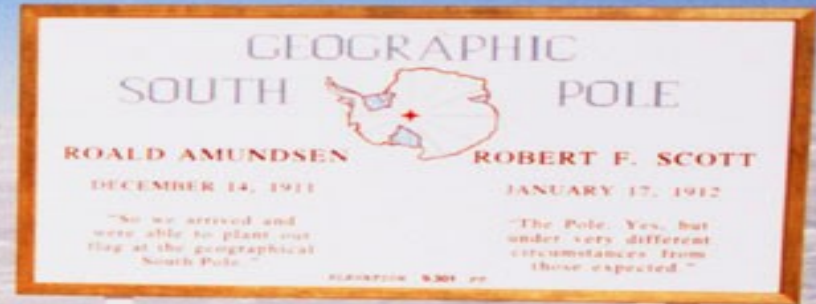


Antarctic Ice moves!!



Previous year's
South Pole
Last year's South
Pole

This year's South
Pole



ROALD AMUNDSEN

DECEMBER 14, 1911

"So we arrived and
were able to plant our
flag at the geographical
South Pole."

SCOTT 5,301 ft

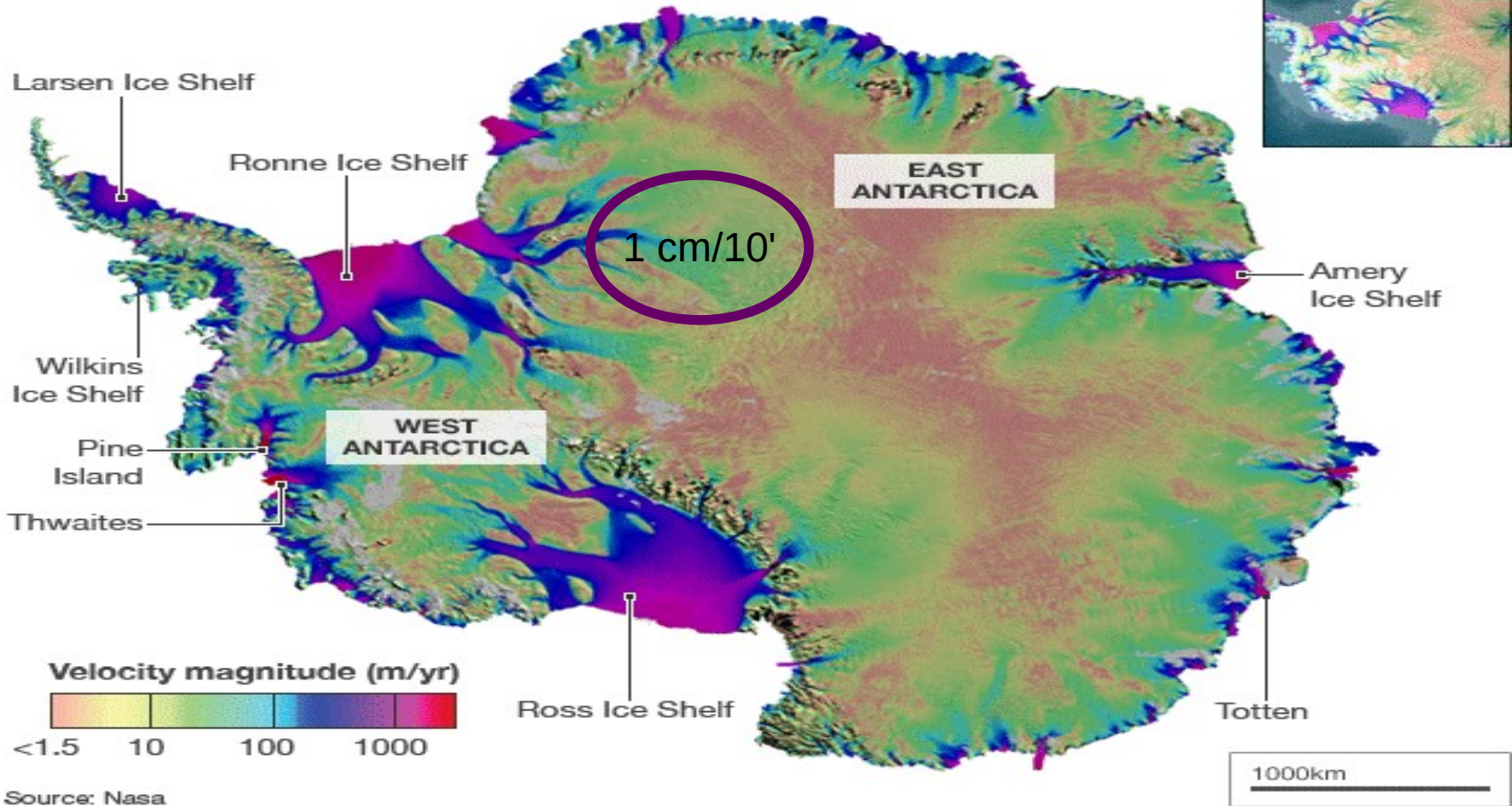
ROBERT F. SCOTT

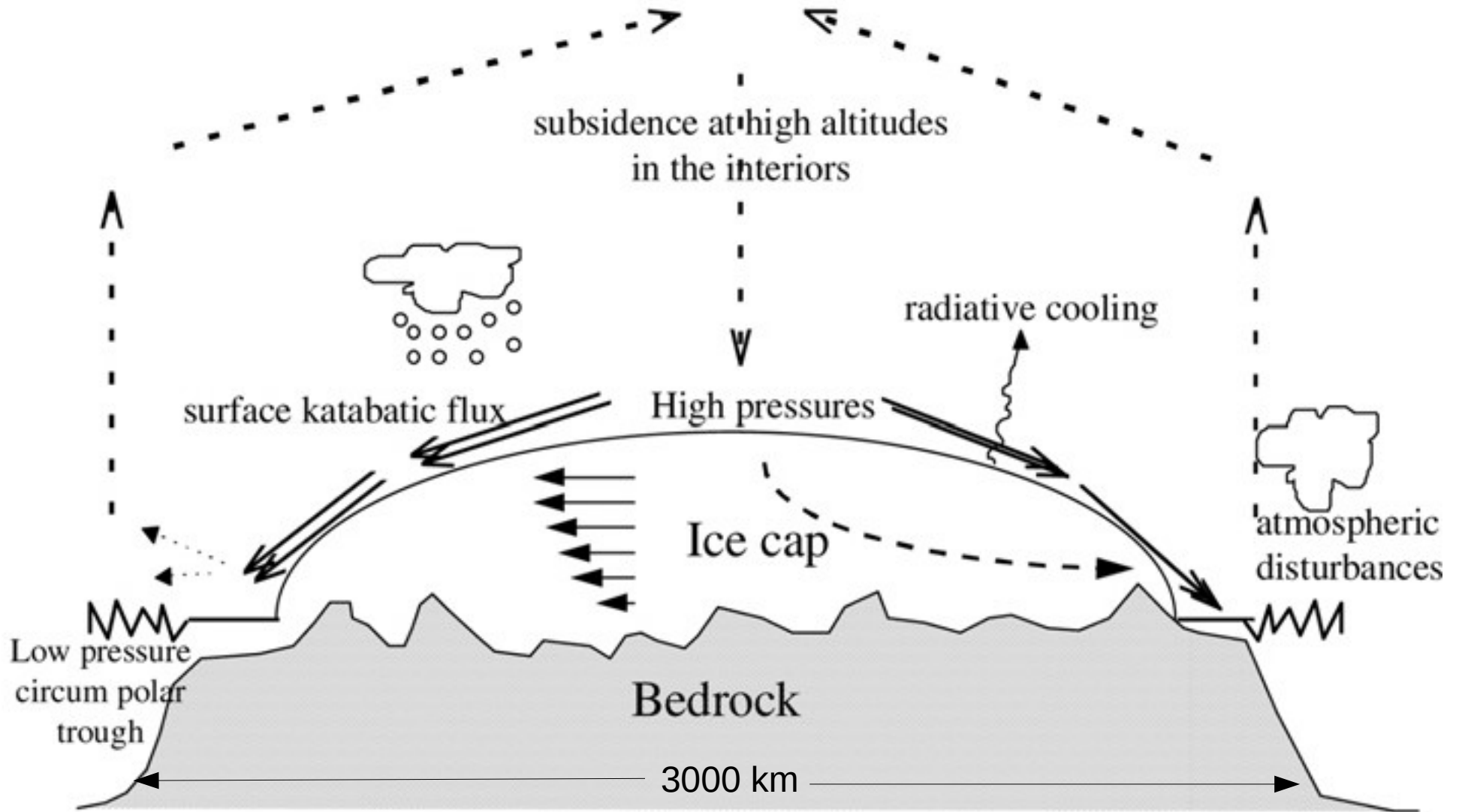
JANUARY 17, 1912

"The Pole. Yes, but
under very different
circumstances from
those expected."

Still, it Moves!

Antarctic ice on the move





Ice Streams

Present day glaciers

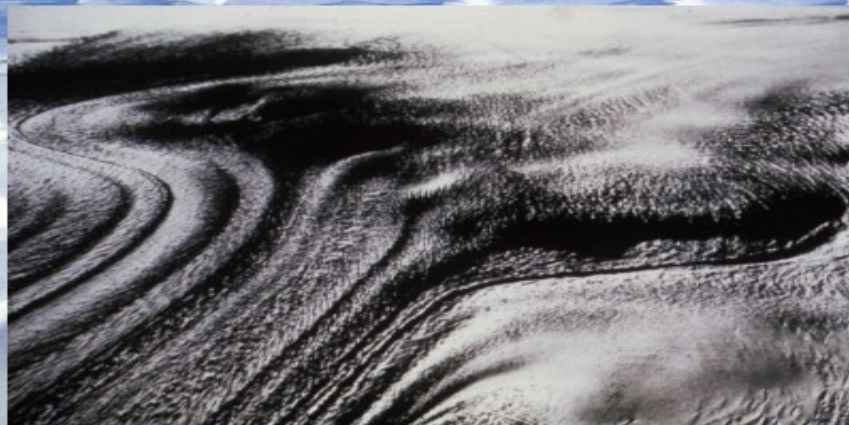
AREA (km²)

South polar region	12 535 000
North polar region	2 081 616
North America	76 880
South America	26 500
Europe	9 276
Asia	115 021
Africa	12
Pacific Australasia	1 015

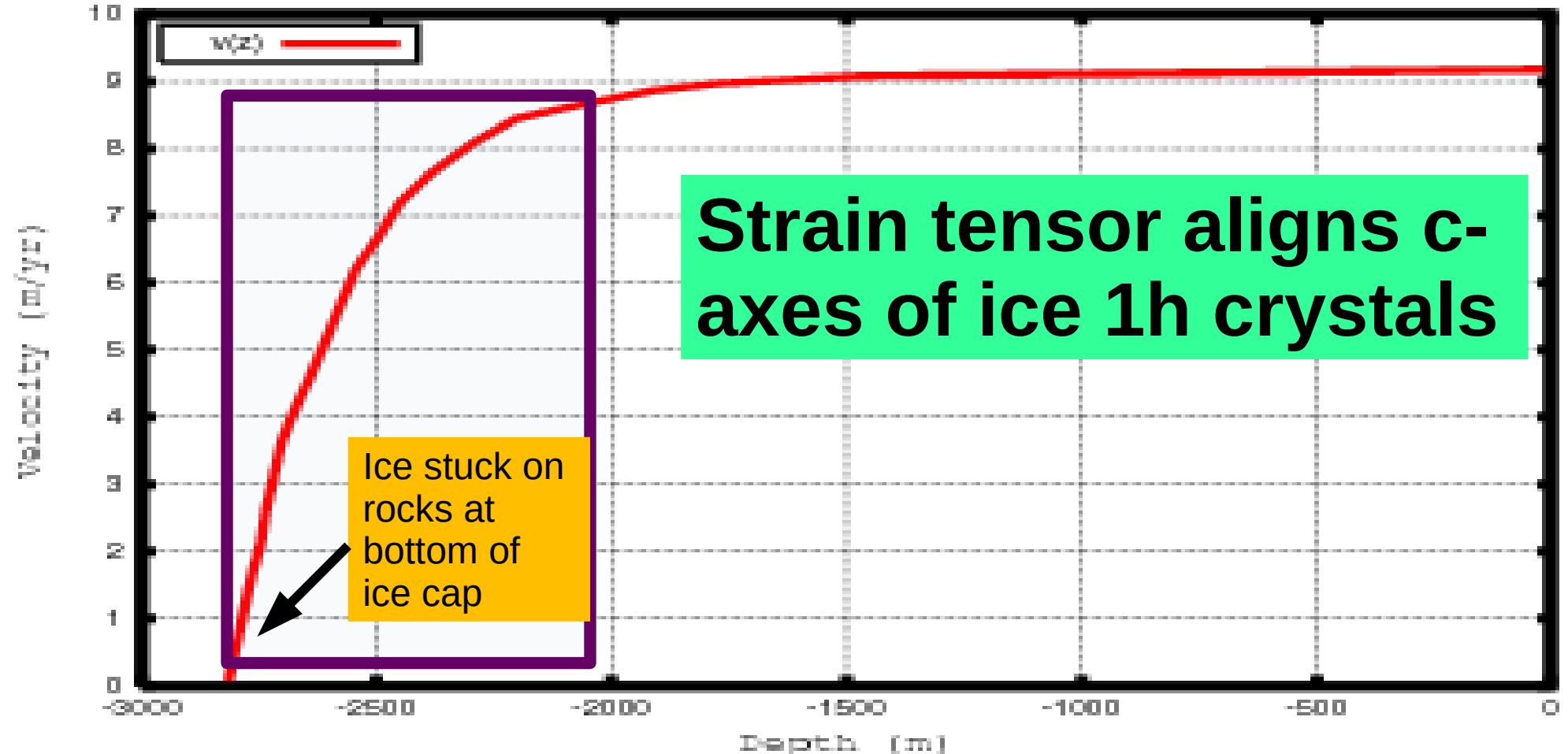
Total world coverage 14 898 320 km²

<http://ralph.swan.ac.uk/geg344/>

Ice stream

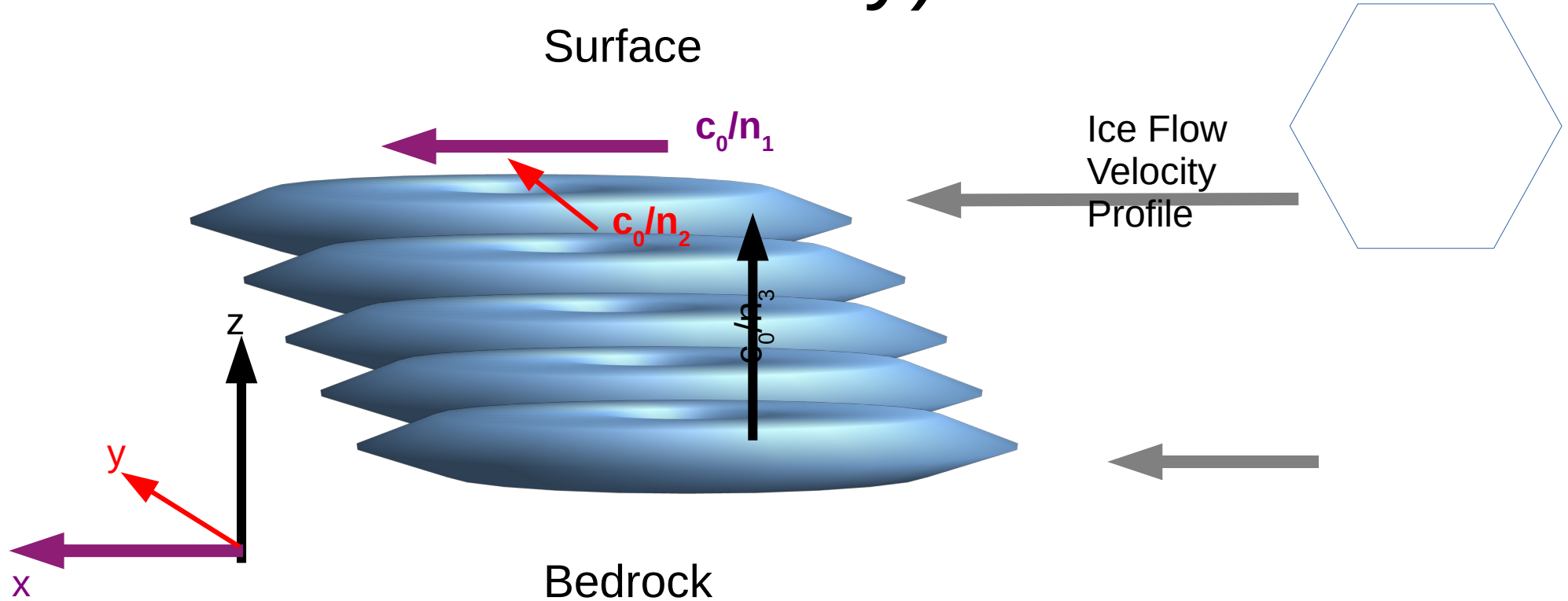


South Pole slip/stick Velocity profile

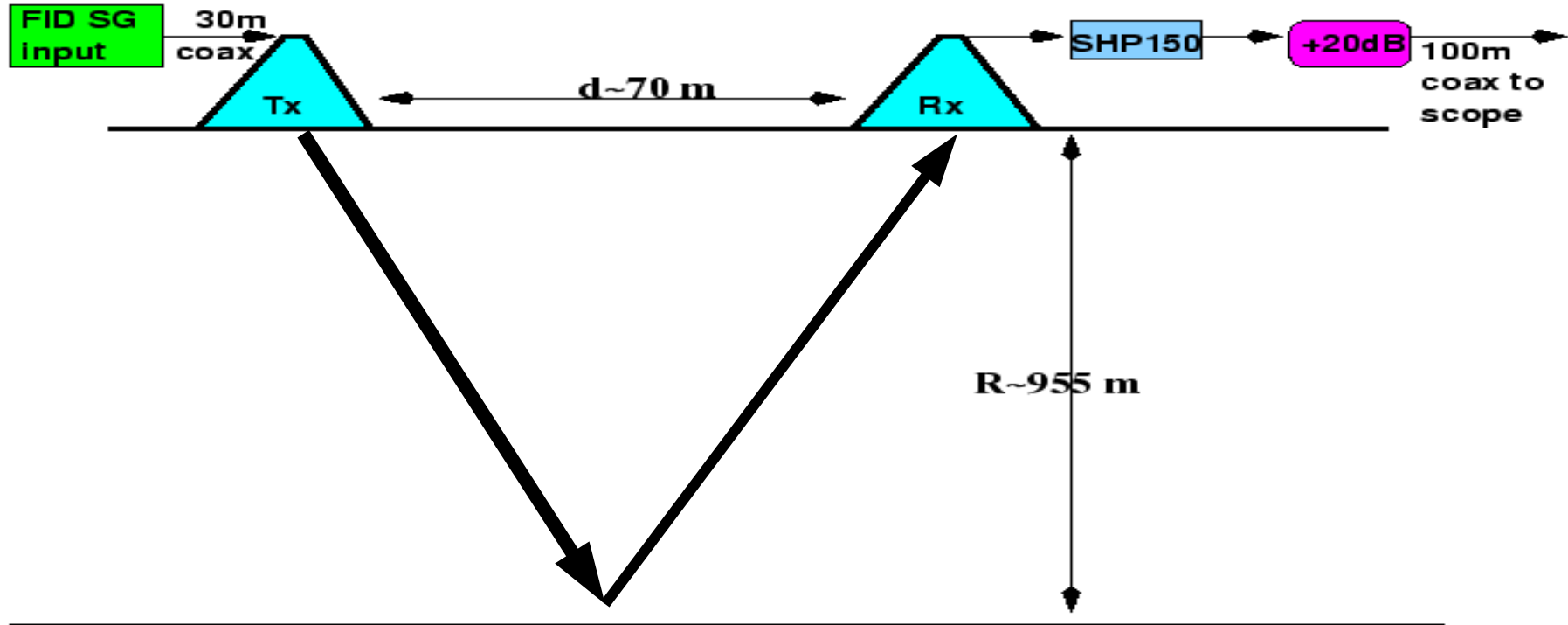


Ice c-axes (symmetry axes and light velocity)

Surface



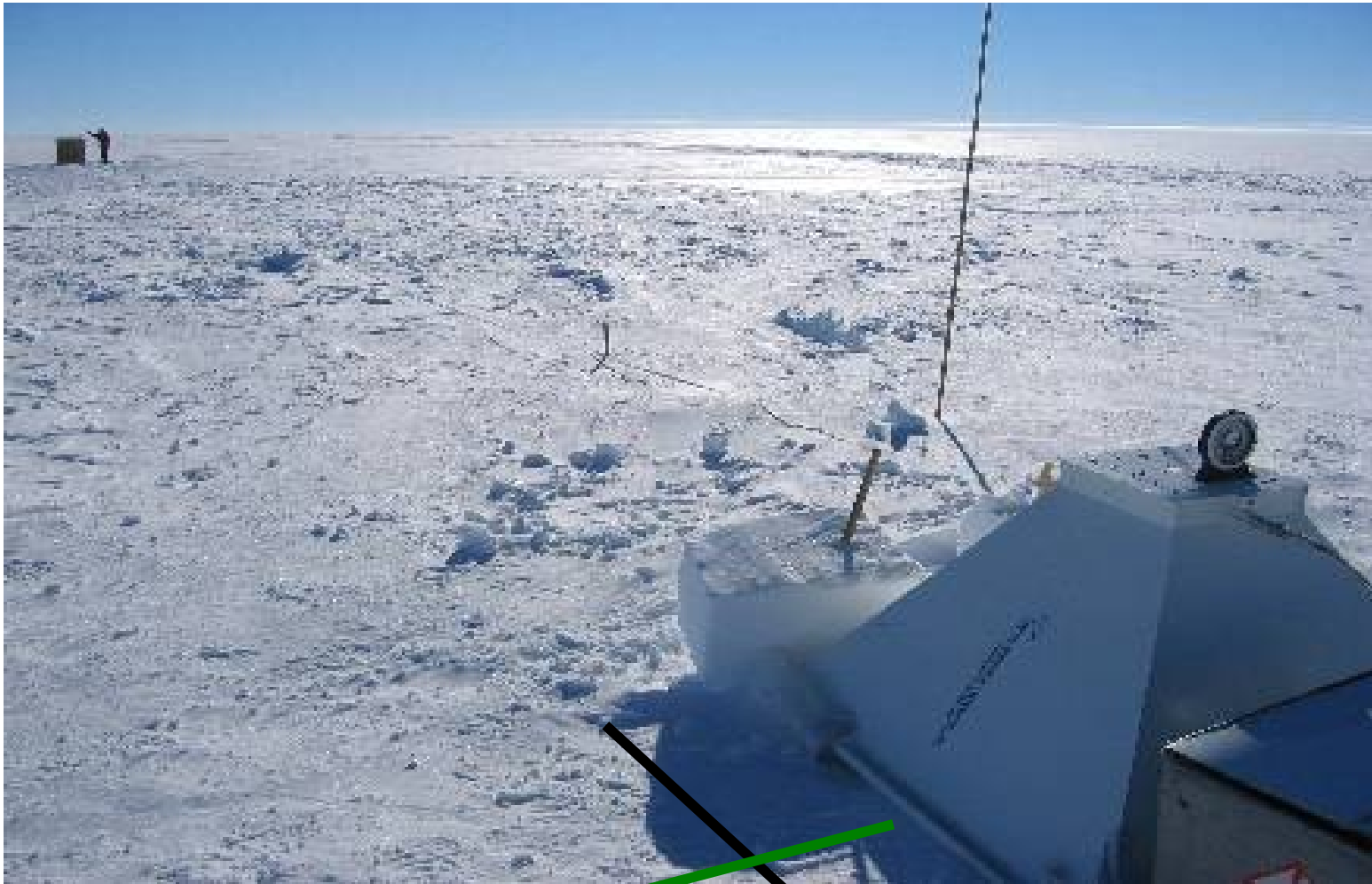
Birefringence tests: South Pole ($R=2850$ m) and Taylor Dome ($R=955$ m)



Taylor Dome Basecamp –*in situ* measurements of RF properties
Investigate birefringence (Astropart.Phys.29:130-157,2008)



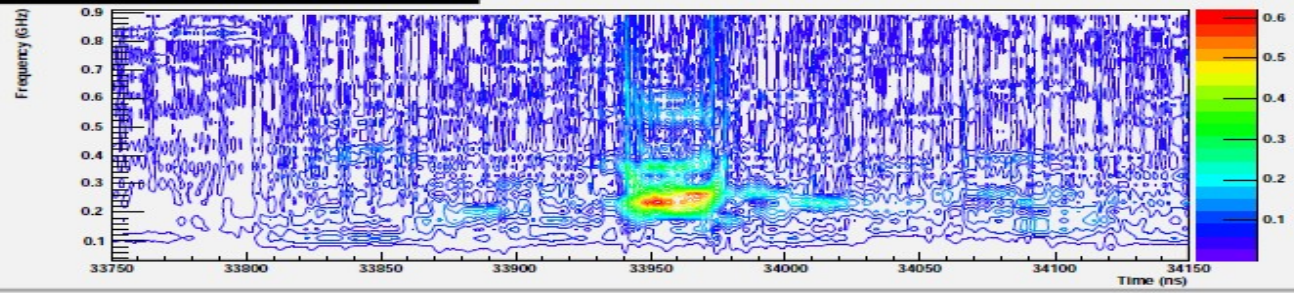




Bottom
Reflection
Studies.

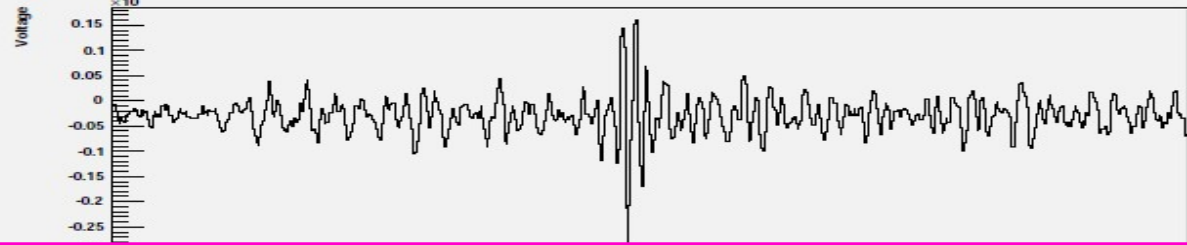
Horn antennas
send waves
polarized
along two
perpendicular
axes

Bedrock reflection (V*R*1000/(15MHz-0.5ns))

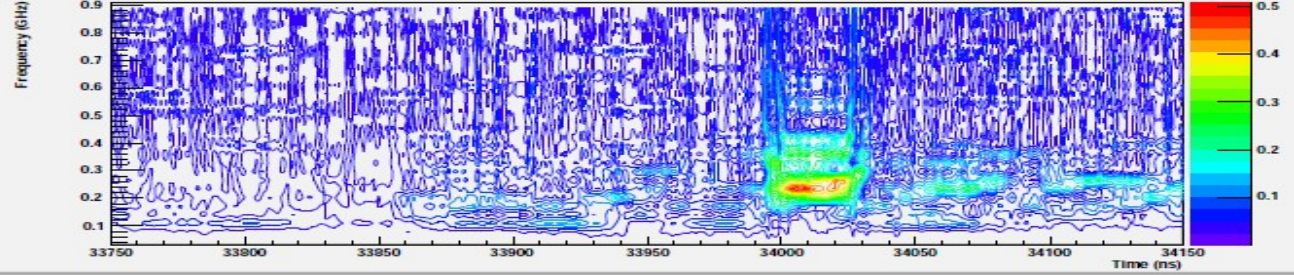


Polarization
parallel to ice-
flow direction

bed reflection/Time domain

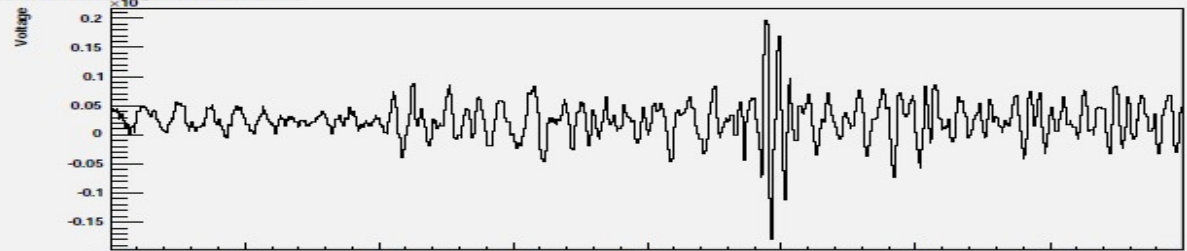


Bedrock reflection (V*R*1000/(15MHz-0.5ns))

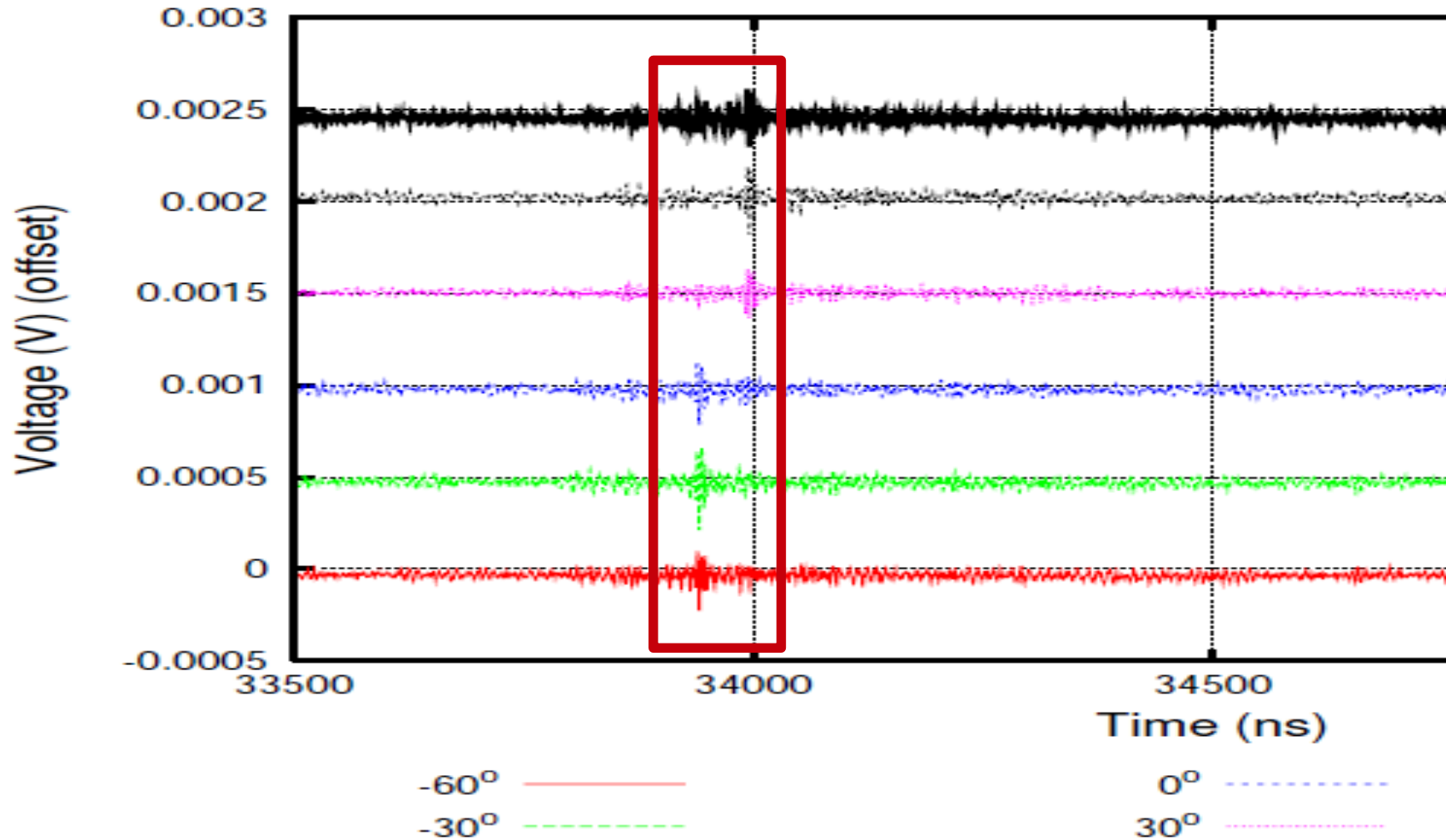


Polarization
Perpendicular to
ice-flow direction

bed reflection/Time domain

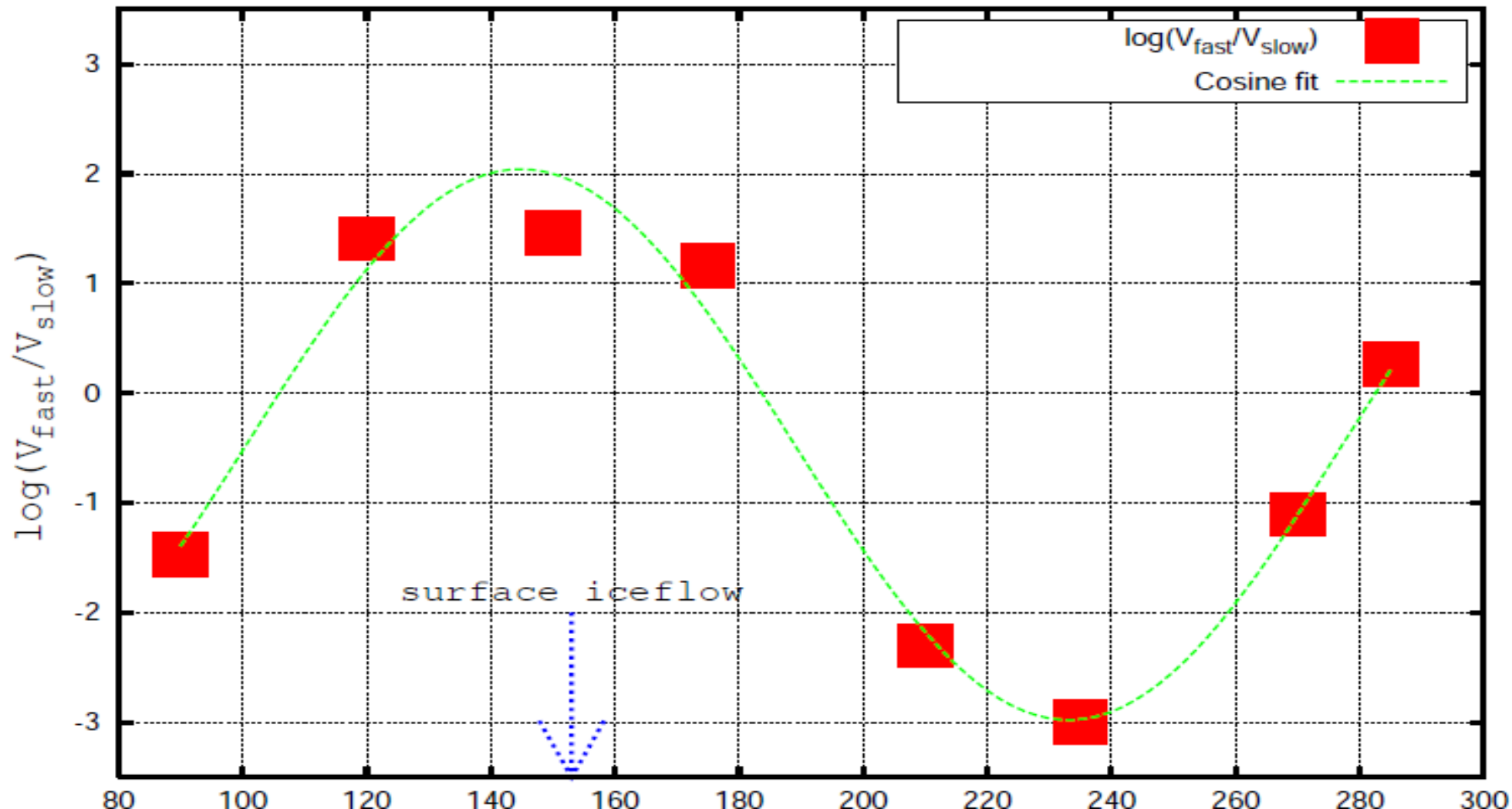


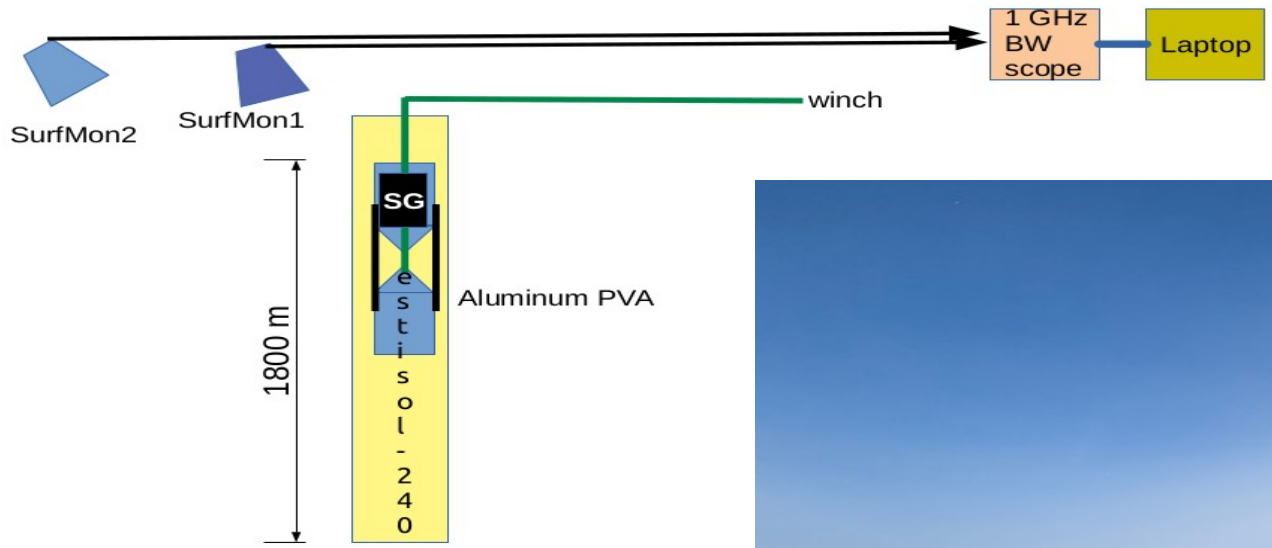
Data on ice birefringence V (polarization)



Can also derive attenuation length from this graph

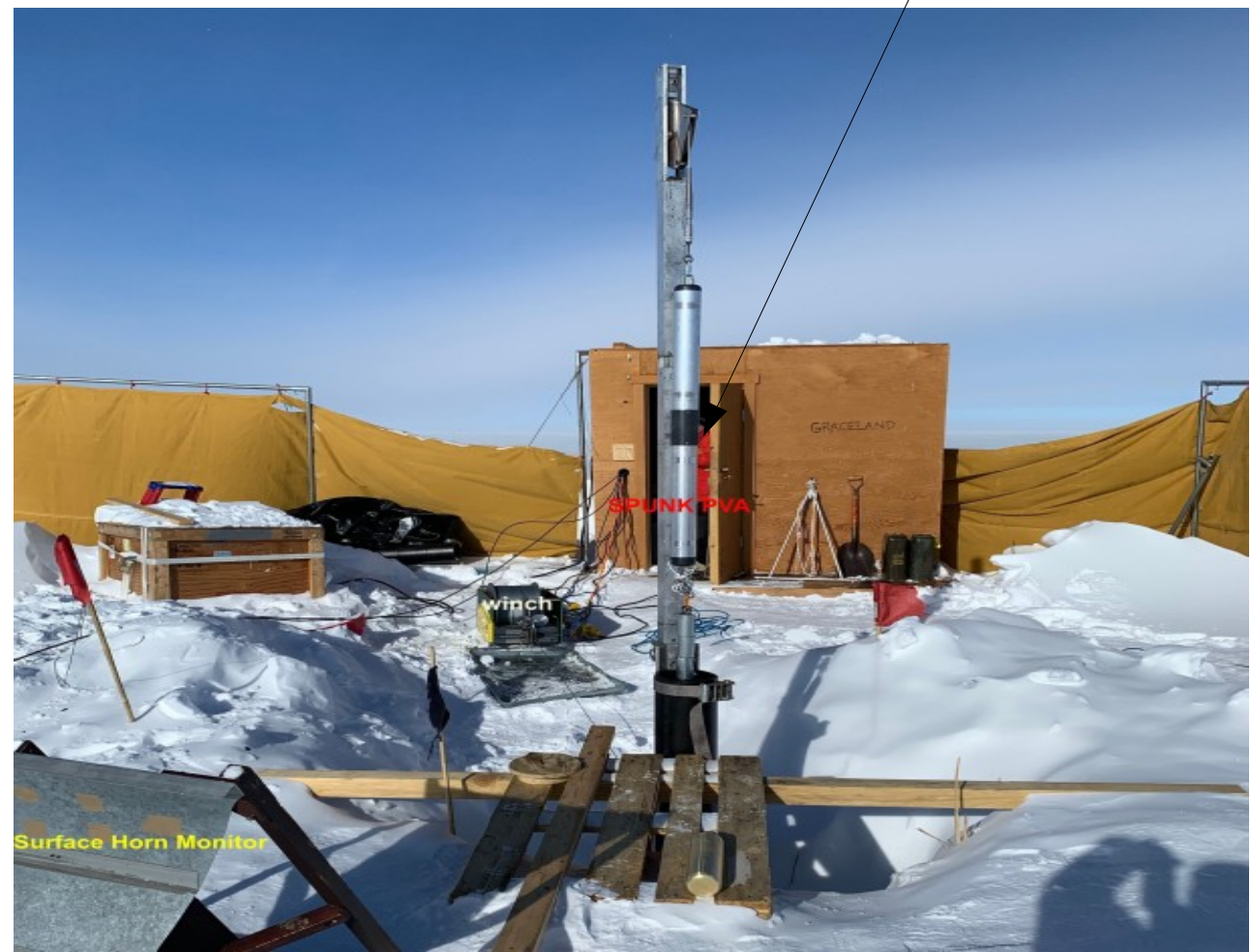
Ratio of $V(\text{fast})/V(\text{slow})$

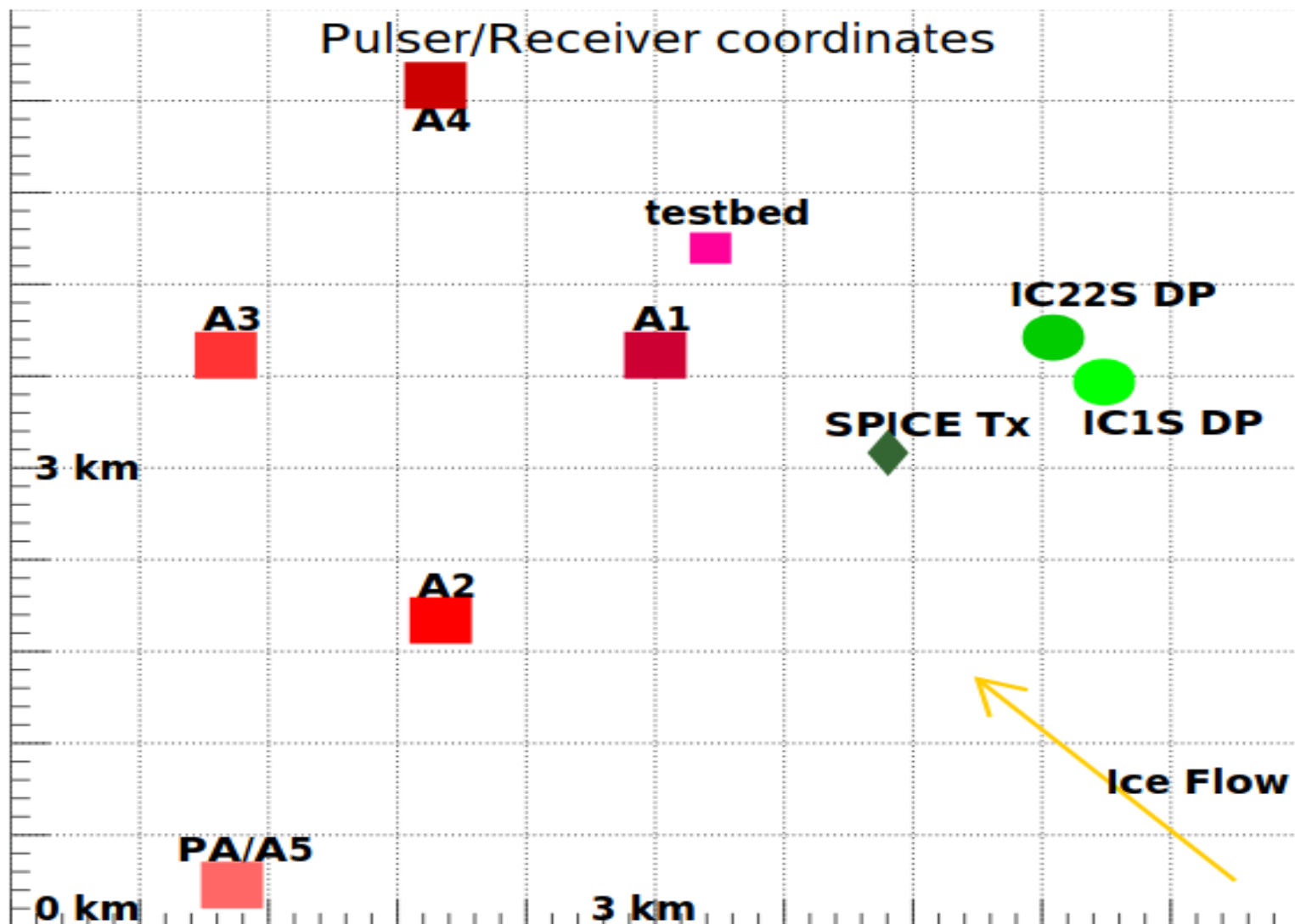




My red parka

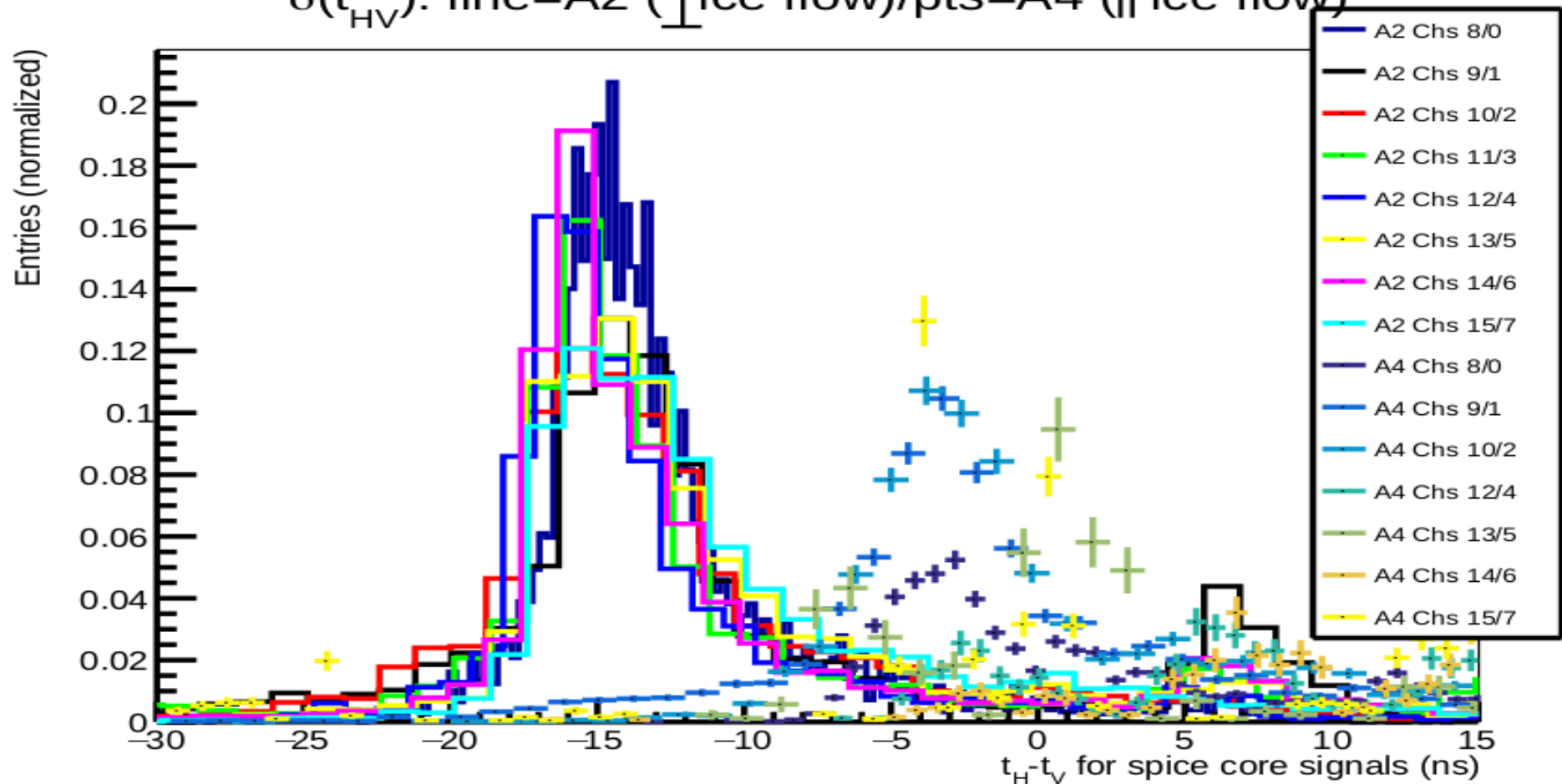
SPICE core testing (Dec., 2018)





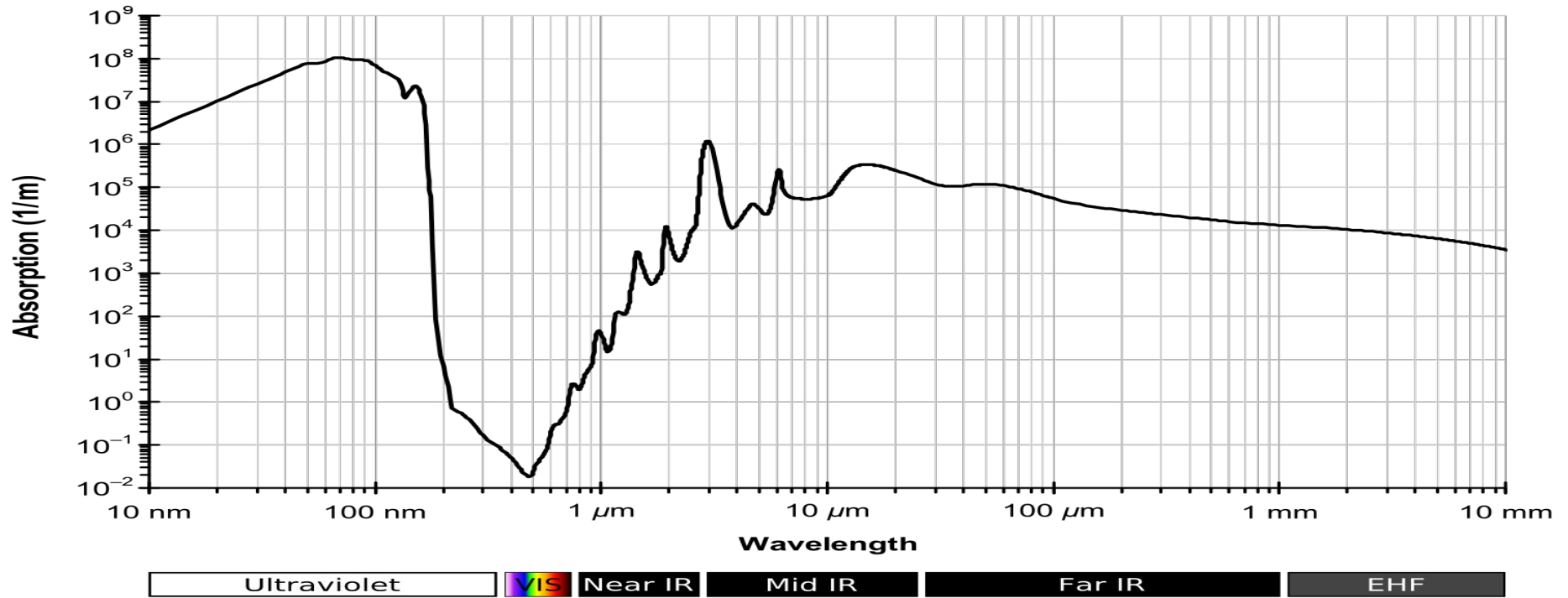
Birefringence! (Estimate $n_o - n_e$)

$\delta(t_{HV})$: line=A2 (\perp ice flow)/pts=A4 (\parallel ice flow)



Why is Antarctica a good place to do astronomy?

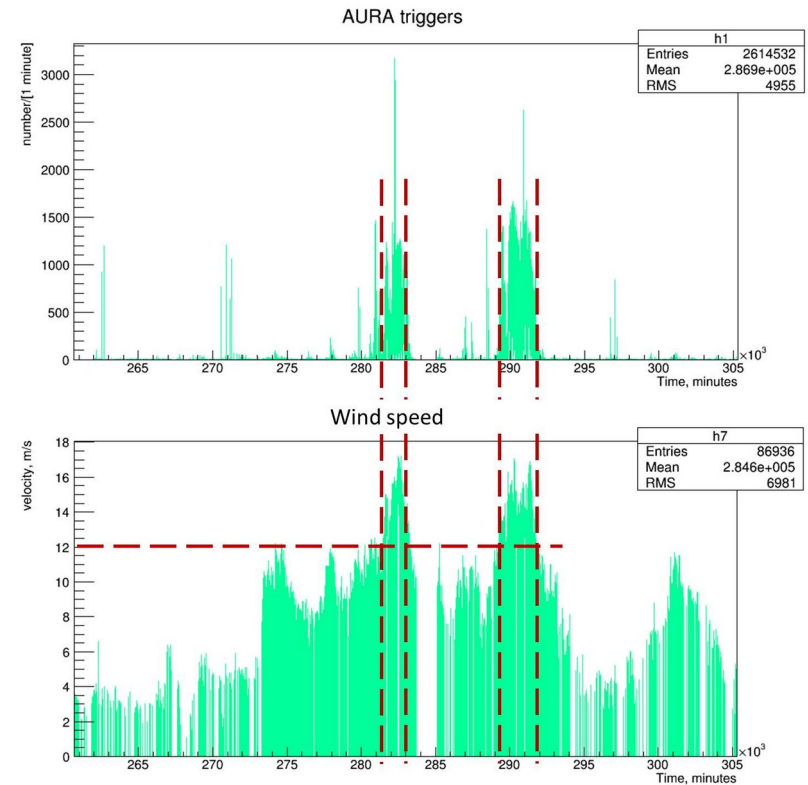
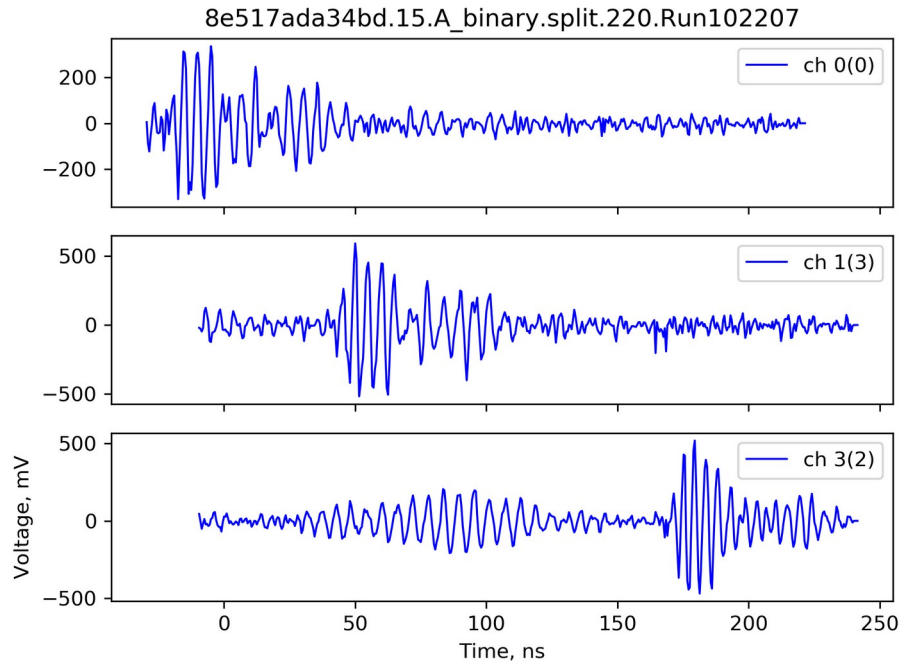
It's cold: 0.03% average humidity=>little or no absorption of CMB at 2.3 K (2.1 mm)



Cold, clean ice: Attenuation length for optical~40 m, Attenuation length for radio~2 km
Auroral Studies: Geomagnetic focusing of field lines
Meteorite Searches: Easy Peasy

Consequences of cool, dry climate

Triboelectric effect (Masha Mikhailova): Wind blowing over snow generates discharge



Ice Exercises

1. Estimate the pressure at the bottom of 3 km of ice, assuming $\rho_{\text{ice}}=0.93(\text{water})$
2. Assume Antarctic ice has an index-of-refraction $n_{\text{ice}}=1.78$. Find the time required for a radio wave to travel from the surface, reflect off the bedrock 3 km deep, and back to the surface again.
3. In fact, Antarctic ice has a property called “birefringence”, wherein different polarizations travel with different speeds.
 1. Find the time difference between two radio waves if $n_x=1.778$ and $n_y=1.782$
 2. You have a radio receiver which is only sensitive to one polarization. Sketch the amount of received power as a function of polarization angle.
4. Now take into account the fact that the first 100 meters of the ice sheet is less dense (“firn”). Assume the index of refraction in the first 100 meters follows: $n(z)=1.+0.78*(z/100)$, after which $n(z)$ is constant at $n=1.78$. Determine the time once again. Assume the ice can be modeled by the same physics that governs a scale height; calculate the expected depth $n(z)$ profile and determine the time to traverse thickness of ice and draw ray from A to B.
5. Given two receivers, separated by 10 meters.
 1. Determine the maximum time difference between the hit times for a plane wave
6. Suppose the attenuation length L of ice has a dependence on temperature that decreases linearly with temperature, such that $L(273K)=250$ m and increases by 10 meters per degree Kelvin (so, $L=350$ m at a temperature of 263K, etc). A radio signal with amplitude 2 kV is emitted spherically, broadcast through 2.8 km of ice and back again. Assuming that the surface temperature in Antarctica is 220K and that the temperature increases linearly, with depth, to 270K at the bedrock, what is the expected echo amplitude measured at the surface?
 1. Same question for Enceladus

Low temperature environment testing in Lawrence, KS the Checkers supermarket vegetable freezer

